WAVE THEORY FOR SEISMOGRAM SYNTHESIS

Chien-Ying Wang, B. S., M. S.

A Digest Presented to the Faculty of the Graduate School of Saint Louis University in Partial Fulfillment of the Requirements for the Degree of Doctor of Philosophy

DIGEST

A complete system of wave integral theory is established for purpose of synthesizing high quality and high frequency seismograms in plane layered media. The system consolidates the foundations of wave theory, and greatly facilitates its numerical application.

A classical contour integration method is extensively studied. Without introducing any sort of attenuation, the integration is taken directly along branch cuts and poles. An attempt to classify the constituents of seismograms from different integration contributions is discussed. Such a discussion proposes a new viewpoint for understanding wave fields.

The eigenfunctions of surface waves are found to have concise analytic solutions. These analytic forms not only provide a firm basis for theoretical development, but also provide a way to study high frequency signals and complicated structures.

The reflection and transmission properties of layer interfaces are reconsidered using a new approach. A simple method is proposed to decompose the wave fields, which can easily be incorporated into our system. Using this method, body as well as surface waves from a particular portion of structure are generated.

A new method for expressing seismic sources is explored, which enables us to isolate the fault orientation and receiver azimuthal dependences, thus faciliating the study of source mechanism. An inversion technique is developed to extract the instrument response coefficients. These coefficients were included in designing a recursive filter to describe the instrument effect.

Comparisons with other methods confirm that the new theory is both flexibile and reliable. The present study clarifies several ambiguities in the theory of the wave integral method and provides several new techniques for simulating wave propagation in the earth.

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ACKNOWLEDGMENTS

Foremost, I wish to thank my advisor, Dr. Robert B. Herrmann, for providing continous guidance and stimulating ideas, throughout the duration of this work. I am especially grateful for his meticulous contributions to major parts of this dissertation.

I benefited greatly from discussions with Dr. Brian J. Mitchell and Dr. Otto W. Nuttli. Their interests provided much of the impetus and direction for the success of the work.

I want to thank Dr. David G. Harkrider for providing me the idea of using analytic solutions of eigenfunctions.

Special thanks are due to Eric J. Haug, Chandan K. Saikia, and David R. Russell. Their assistance in developing and maintaining the computer programs has been of great help. Finally, I thank my wife Tai-Chi and my son Albert, for their understanding and sacrifices in behalf of this effort.

This research was supported by National Science Foundation under Grant PFR-7909795.

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CHAPTER I

INTRODUCTION

1.1 The Problem

In recent years synthetic seismograms have become increasingly useful as an aid to seismic data interpretation. Various techniques have been developed to calculate theoretical seismograms which are valid, at least, for plane-layered media. However, further improvements are still desirable. The problems of computational efficiency and accuracy, restricted use of a computer, continue to be a challenge. To solve them a thorough understanding of the nature of the a detailed reexamination of the related problem and theories are truly necessary.

The work of Thomson (1950)and Haskell (1953)first permitted the treatment of multi-layered media using matrix calculus. Prior to that time it was only possible to consider simple one- and two-layer models. Many developments based on the Thomson-Haskell technique have been pursued. The extension of their work to compound matrices and source specification are notable examples. Most of these developments, however, can be related to Haskell's work between 1953 and 1964. Because of simplicity in both concept and mathematics,

Haskell's research provided an easily understandable approach to the problem. The purpose of this dissertation is to reconcile different formalisms and to establish a complete system to treat wave propagation in layered media using Haskell's work as a starting point. Such a system will provide a foundation for handling more complicated cases and for extending the theoretical studies.

The formulation of the method developed by Haskell is constructed in the frequency domain. The complete response at a particular frequency is represented by semi-infinite integrals with respect to wavenumber so as to automatically include all types of To the boundary conditions, the responses match described in terms of a layer matrix. There still exist several questions about the properties of this matrix and its compound form. A systematic approach to the problem will provide further insight to the Haskell matrices, and will promote their use.

After stacking the layer matrices over layers, time domain seismograms can be synthesized by performing any of several integrations in complex frequency-wavenumber plane. The behavior of the integrands must be known in order to choose the proper numerical method for integrations. This analysis also provides a framework for understanding the nature of different signals

on the recorded seismogram, which in turn establishes a basis for exploring the effects of seismic source and earth structure on the seismograms.

The main application of the theory developed in this dissertation will be for synthesizing high quality seismograms at high frequencies. High frequency signals have become increasingly important in the studies of deep earth structure, earthquake source mechanism, and strong ground motion.

1.2 Historical Review

The era of the seismogram synthesis was opened by who generated the first synthetic (1904), Lamb Η. seismogram for an impulse source acting upon a infinite medium. early studies were for simple Most models such as liquid layers or one solid layer over a halfspace (see Ewing et al, 1957). With the advent of modern computers, the research rapidly grew to cover more complicated models and more sophisticated cases. In the last two decades, theoretical as well as numerical developments have progressed tremendously. We will now review the contribution of papers on some them have formed the thesizing seismogram. Most of foundation of modern theoretical seismology.

In most theoretical developments, the evaluation

of synthetic seismograms generally can be divided two parts: first, the solution of an ordinary differential equation using transform methods with appropriate radiation and boundary conditions, and second, the evaluation of the corresponding inverse transforms. Several methods exist for both parts and many combinations are possible. One important approach, called 'generalized ray theory', uses the Laplace transform technique and generates waves at discrete time points by summing hundreds of rays. The basis of this method comes from Cagniard (1962) and de-Hoop (1960).Although this theory works well in predicting particular phases, it can be inaccurate and cumbersome when modeling long time duration seismograms for a multilayered structure. Hron(1972), Kennett (1974) and Wiggins and Madrid (1974) have made great efforts to improve the efficiency of calculating the responses of a large number of rays. A recent review of the method can be found in Pao and Gajewski (1977). Helmberger (1968) and Vered and Ben-Menahem (1976) present typical applications of the method to layered media.

Another approach, called 'wave theory', uses the Fourier transform technique and calculates all the waves excited in the structure by integrating over frequency and wavenumber (or slowness). This method encompasses normal-mode theory, especially for the determination of dispersion of surface waves (Press et

al, 1961). Because of increasing complexity at higher frequencies, the method suffers from computational inefficiency and stability problems. This dissertation will thoroughly explore this theory, especially at high frequencies.

In 1953 Haskell published a corrected version of Thomson's (1950) theory of elastic waves in a plane multilayered medium. Haskell's study introduced the 'matrix method' to seismic wave studies. This method provides a systematic approach and greatly facilitates numerical computation. Since then, the theory has been extended principally by Haskell (1963, 1964) and krider (1964, 1970) to deal with the surface wave motion. Because of growing exponential terms in the matrix under certain conditions, the simple matrix method suffered from numerical problems which caused loss of precision. Knopoff (1964), Dunkin (1965), and Thrower (1965)reformulated the computational prousing compound matrix extensions in which the cedures minors of the layer matrices are propagated from interface to interface instead of the matrices themselves, so that the squared exponential terms never thus controlling the precision problem. Randall (1967), Watson (1970), and Schwab and Knopoff successful improvements in computational effimade ciency and accuracy. More recently Abo-Zena (1979) and Menke (1979)reexamined the matrix approach for

extension to very high frequencies.

The existence of dispersive surface waves has been recognized since the early days of the science of seismology. Keilis-Borok (1960) presented a detailed study of surface waves generated in a layered medium. Vlaar (1966) applied eigenfunction theory in wave generation. Saito (1967) analysis of Love developed a solution for surface wave excitation terms of mutually orthogonal eigenfunctions for the generation of free oscillations in a radially inhomogeneous earth, and for surface waves in a vertically inhomogeneous flat earth. A recent contribution by Takeuchi and Saito (1972) summarized the previous studies, considering both theory and numerical methods. They applied the calculus of variations to derive the derivative-related quantities such as group velocity, attenuation factor, etc., from the surface wave eigenfunctions. This work provided a detailed mathematical basis for surface wave eigenfunction theory.

Hudson (1969a,b) extended the work of Haskell (1964) and Harkrider (1964) to synthesize seismic signals at teleseismic distances. Hudson's analysis is applicable for large epicentral distances, as all near-field terms are ignored. However, as shown by Herrmann (1978a), the truncation of these terms causes non-causal arrivals. Herrmann (1979) and Wang and

Herrmann (1980) expressed the Haskell formalism in a more concise form, and improved the integration method of Ewing et al (1957) for synthesizing high quality seismograms. Other interesting methods to treat the problem of integration over the wavenumber-frequency domain appear in the work of Chapman (1978), Apsel (1979), and Bouchon (1979,1981).

Using another approach, Fuchs (1968) and Fuchs and (1971) recognized that the reflectivity of some layers (reflection zone) can be isolated and excited by incident waves from a 'transmission zone' to produce synthetic body wave seismograms. Their development named the 'reflectivity method'. Fuchs (1971) simplified the solution by using the stationary phase approximation. The method is largely used in the study of crust or upper mantle structures from an artificial explosive source. However Kind and Müller (1975,1977) and Müller and Kind (1976) extended the method include a double-couple point source, and after adding the earth flattening correction (Müller, 1971), they were able to simulate many real earth phases, such as Sn, ScS, SKS, etc. Kennett (1975) skillfully separated transmission response beneath the source and the receiver in order to study the laterally varying struc-Another achievement of this method was the introduction of attenuation into the layers in the form of complex velocities (Kind, 1978). This modification enables a computation for only a short time window, even if the nonattenuated seismogram has a long duration.

Extending the systematic development by Gilbert Backus (1966) of the 'propagator matrix', Kennett (1974) and his colleagues were able to express the matrix in terms of reflection and transmission properties of the stratified medium. This research bridged the gap between the wave approach and the ray method, and also provided an internal view of the layer matrix theory. Kennett et al (1978) extensively explored the symmetry properties of reflection and transmission coefficients. These symmetries not only reflect the theory of reciprocity, but provide a novel approach exploit the properties of Haskell matrices. Kennett and Kerry (1979), Kennett (1980), and Kerry (1981) gave a complete derivation for this 'reflection and transmission coefficient' method and also proposed some interesting numerical evaluation techniques.

In searching for the roots of the period equation in the complex frequency plane, Gilbert (1964) described several wave-guide generated waves as leaky modes. Alsop (1970) interpreted these waves as a constructive interference of reverberating waves within a layer. Abramovici (1968) and Cochran et al (1970) found a relationship of this kind of mode to the regular nor-

mal mode. Watson (1972) gave a very detailed discussion of leaky modes by using real frequency-complex wavenumber analysis. In the observed data, it is believed that a dispersive body wave, called the PL wave, which arrives between the direct P and S waves, arises from the contribution of leaky modes. Oliver and Major (1960), Laster et al (1965), and Su and Dorman (1965) provided some real and experimental data analysis to reveal the properties of this wave. The success of improved computation with Haskell's method may lead to further insight into this phase.

CHAPTER II

WAVE INTEGRAL THEORY

The objective of this chapter is to establish a complete, self-contained, system for wave integral theory. The solution for the surface displacements is found using procedures parallel to those of Haskell (1964). The derivations are put forth in an easily understandable, step by step, way. Some interesting symmetry properties of the layer matrix or its compound form are revealed, which are then used to simplify the numerical application or theoretical extension.

integration technique of Herrmann numerical (1978a, 1979) also is extended. The method requires contour integration in the complex wavenumber plane, the performance of which is complicated by the presence of singularities. The contour integration reduces to a consideration of pole residue contributions and branch line integrals. A detailed discussion of these aspects is undertaken, which is then used to improve the computational efficiency. The discussion also provides a framework for the investigation of different signals making up the complete seismogram.

2.1 Haskell's Matrix

In the present section a previous modification (Wang and Herrmann, 1980) of Haskell's theory (Haskell 1963, 1964) is further developed and revised. All of the derivations are made in a step-by-step manner for clarity. Several significant differences with respect to Haskell's papers are specifically noted. Such revisions are made not only to simplify the expressions, but also to compare them to other related formalisms to be discussed later. The three different types of waves, P, SV and SH, existing in the layered media are all included.

We define a semi-infinite elastic medium made up N parallel, solid, homogeneous, isotropic layers (Figure 1). Each layer is characterized by the compressional wave velocity α , the shear wave velocity β , the density ρ , and the layer thickness d . Any linear variation of elastic properties can be approximated by many small layers (Fuchs, 1968). The m'th layer is bounded by the m and m+l interfaces. Thus any quantities at the free surface are denoted by the subscript 'l'. If a water layer is placed on the top, it will be assigned a layer index '0'. The parameters in the half-space are denoted by the subscript 'N'. In order to match the boundary conditions at the horizontal layer interfaces, a cylindrical coordinate

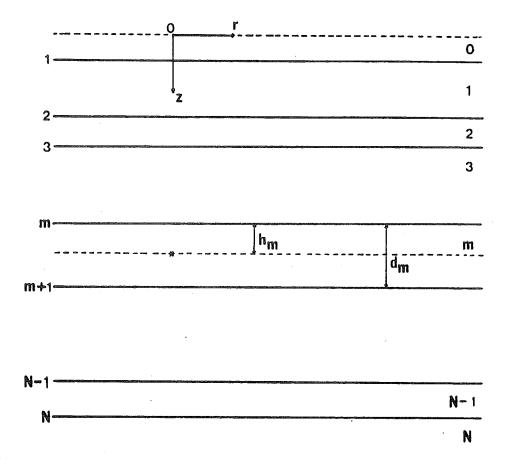


Figure 1. Direction of axes, numbering of layers and interfaces, and the depth of source in the source layer m.

system (r, ϑ, z) is chosen with origin on the free surface just above the source, and the z axis is taken positive downward.

In the following, expressions will be derived for the displacements from a point source using Haskell's notation (Haskell, 1964). Let us start from the very beginning. The displacement field can be expressed in terms of three different types of potentials, φ for P waves, ψ for SV waves, and χ for SH waves. The function φ is known as the scalar potential. ψ and χ are the vector potentials. Define

$$\mathbf{A} = \mathbf{e}_z \chi + \nabla \times (\mathbf{e}_z \psi).$$

where \mathbf{e}_z is the unit vector in the z direction. Variables with bold cases represent vectors or matrices. The displacement \mathbf{u} can be expressed as

$$\mathbf{u} = \nabla \varphi + \nabla \times \mathbf{A}$$

$$= \nabla \varphi + \nabla \times (\mathbf{e}_z \chi) + \nabla \times \nabla \times (\mathbf{e}_z \psi), \qquad (\text{II-1-1})$$

where the potential functions φ , ψ , and χ satisfy the wave equations:

$$\nabla^{2} \varphi = \frac{1}{\alpha^{2}} \frac{\partial^{2} \varphi}{\partial t^{2}}$$

$$\nabla^{2} \psi = \frac{1}{\beta^{2}} \frac{\partial^{2} \psi}{\partial t^{2}}$$

$$\nabla^{2} \chi = \frac{1}{\beta^{2}} \frac{\partial^{2} \chi}{\partial t^{2}} .$$
(II-1-2)

Expanding the gradient and curl operators in equation (II-1-1) in cylindrical coordinates, we obtain the

three components of displacement in terms of potentials:

$$u_{r} = \frac{\partial \varphi}{\partial r} + \frac{\partial \chi}{r \partial \vartheta} + \frac{\partial^{2} \psi}{\partial r \partial z}$$

$$u_{z} = \frac{\partial \varphi}{\partial z} + \frac{\partial^{2} \psi}{\partial z^{2}} - \nabla^{2} \psi$$

$$u_{\vartheta} = \frac{\partial \varphi}{r \partial \vartheta} - \frac{\partial \chi}{\partial r} + \frac{\partial^{2} \psi}{r \partial \vartheta \partial z}$$
(II-1-3)

If the time dependence is isolated by the Fourier transform factor $e^{i\omega t}$, the resulting Helmholtz equations (II-1-2) can be solved using characteristic functions of three cylindrical coordinates:

$$\varphi(r,\vartheta,z,\omega) = \frac{\cos n\,\vartheta}{\sin n\,\vartheta} J_n(kr) \left\{ Z_1(z) \right\}
\psi(r,\vartheta,z,\omega) = \frac{\cos n\,\vartheta}{\sin n\,\vartheta} J_n(kr) \left\{ Z_2(z) \right\}
\chi(r,\vartheta,z,\omega) = \frac{\cos n\,\vartheta}{-\sin n\,\vartheta} J_n(kr) \left\{ F_3(z) \right\},$$
(II-1-4)

where $J_n(kr)$ is the Bessel function of the first kind of order n. The solution involving the Bessel function of the second kind, Y_n , is not used since the solution must be valid at $r=\emptyset$ where this function becomes unbounded. The subscript index, n, indicates azimuthal mode number, and k indicates the horizontal wavenumber. Z_1 , Z_2 , and F_3 are functions of z only, satisfying

$$\frac{d^{2} Z_{1}}{dz^{2}} - \nu_{\alpha}^{2} Z_{1} = 0$$

$$\frac{d^{2} Z_{2}}{dz^{2}} - \nu_{\beta}^{2} Z_{2} = 0$$

$$\frac{d^{2} F_{3}}{dz^{2}} - \nu_{\beta}^{2} F_{3} = 0$$
(II-1-5)

where

$$\nu_{\alpha} = \begin{cases} \sqrt{k^2 - \frac{\omega^2}{\alpha^2}} & k \ge \frac{\omega}{\alpha} \\ i \sqrt{\frac{\omega^2}{\alpha^2} - k^2} & k < \frac{\omega}{\alpha} \end{cases}$$

$$\nu_{\beta} = \begin{cases} \sqrt{k^2 - \frac{\omega^2}{\beta^2}} & k \ge \frac{\omega}{\beta} \\ i \sqrt{\frac{\omega^2}{\beta^2} - k^2} & k < \frac{\omega}{\beta} \end{cases}$$

Now let us consider one homogeneous layer first. By substituting equation (II-1-4) into (II-1-3), we find

$$4\pi\rho u_{\tau}(r,\vartheta,z,\omega) = \cos n\vartheta \left[-\left\{ -\left(\frac{dZ_{2}^{c}}{dz} + Z_{1}^{c}\right)k\right\} \frac{dJ_{n}(kr)}{dkr} - \left\{kF_{3}^{c}\right\} \frac{nJ_{n}(kr)}{kr} \right] + \sin n\vartheta \left[c \to s\right]$$

$$4\pi\rho u_{z}(r,\vartheta,z,\omega) = \cos n\vartheta \left[\left\{ k^{2}Z_{2}^{c} + \frac{dZ_{1}^{c}}{dz} \right\} J_{n}(kr) \right]$$

$$+ \sin n\vartheta \left[c \to s \right]$$
(II-1-6)

$$4\pi\rho u_{\vartheta}(r,\vartheta,z,\omega) = \sin n\vartheta \left[\left\{ kF \right\}_{3}^{c} \right\} \frac{dJ_{n}(kr)}{dkr} + \left\{ -\left(\frac{dZ_{2}^{c}}{dz} + Z_{1}^{c} \right) k \right\} \frac{nJ_{n}(kr)}{kr} \right] - \cos n\vartheta \left[c \to s \right],$$

where $c \to s$ indicates the term in brackets above with c replaced by s. The constant ρ is included for simplifying the notation when the stresses are involved, and 4π for balancing the source terms which come from the Green's function. Using the transformed displacements given by equations (II-1-6), the transformed

stresses across a horizontal plane are

$$\begin{split} &4\pi T_{zr}(r,\vartheta,z,\omega) = 4\pi\mu \left[\frac{\partial u_r}{\partial z} + \frac{\partial u_z}{\partial r} \right] \\ &= \cos n\vartheta \left[\omega^2 \left[(\gamma - 1)Z_2^c + \frac{\gamma}{k^2} \frac{dZ_1^c}{dz} \right] k \right] \frac{dJ_n(kr)}{dkr} - \left\{ \frac{\mu}{\rho} \frac{dF_3^c}{dz} k \right\} \frac{nJ_n(kr)}{kr} \right] \\ &+ \sin n\vartheta \left[c \to s \right] \end{split}$$

$$\begin{split} 4\pi T_{zz}(r,\vartheta,z,\omega) &= 4\pi \left[\lambda \left(\frac{\partial u_r}{\partial r} + \frac{1}{r} \frac{\partial u_\vartheta}{\partial \vartheta} + \frac{u_r}{r} + \frac{\partial u_z}{\partial z} \right) + 2\mu \frac{\partial u_z}{\partial z} \right] \\ &= \cos n\vartheta \left[\omega^2 \left\{ (\gamma - 1)Z_1^c + \gamma \frac{dZ_2^c}{dz} \right\} J_n(kr) \right] \\ &+ \sin n\vartheta \left[c \to s \right] \end{split} \tag{II-1-7}$$

$$\begin{split} &4\pi T_{z\vartheta}(r,\vartheta,z,\omega) = 4\pi\mu \left[\frac{1}{r} \frac{\partial u_z}{\partial \vartheta} + \frac{\partial u_\vartheta}{\partial z} \right] \\ &= \sin n\vartheta \left[\left\{ \frac{\mu}{\rho} \frac{dF_3^c}{dz} k \right\} \frac{dJ_n(kr)}{dkr} + \omega^2 \left\{ \left[(\gamma - 1)Z_2^c + \frac{\gamma}{k^2} \frac{dZ_1^c}{dz} \right] k \right\} \frac{nJ_n(kr)}{kr} \right] \\ &- \cos n\vartheta \left[c \to s \right] \end{split}$$

where

$$\gamma = \frac{2\beta^2 k^2}{\omega^2} \ .$$

In equations (II-1-6) and (II-1-7), the azimuth dependent terms, the cosine and sine functions, are implied by the superscripts c and s, respectively. Since the model is plane layered, such a dependence arises solely from the orientation of the force system of the source. For homogeneous solutions, we first ignore the azimuthal terms and define the functions U_T , U_Z , T_Z , T_T , U_ϑ , and T_ϑ from the bracketed terms in equation (II-1-6)

and (II-1-7) as follows

$$\rho U_r = -\left\{\frac{dZ_2}{dz} + Z_1\right\} k$$

$$\rho U_z = k^2 Z_2 + \frac{dZ_1}{dz}$$

$$T_z = (\gamma - 1)Z_1 + \gamma \frac{dZ_2}{dz}$$

$$T_r = \left[(\gamma - 1)Z_2 + \frac{\gamma}{k^2} \frac{dZ_1}{dz}\right] k$$

$$\rho U_\vartheta = F_3 k$$

$$T_\vartheta = \frac{\mu}{\rho} \frac{dF_3}{dz} k$$
. (II-1-8)

The azimuthal terms dropped will be reconsidered when the source is introduced.

These functions are useful, since all of satisfy the transformed boundary conditions: (a) the continuity of displacement and stress across the interfaces of layers, (b) the vanishing of stresses at the free surface, and (c) no upward waves in the halfspace if the source is inside one of the upper If we further require that exponentially in the halfspace, these functions are nothing but the eigenfunctions of a boundary value In equation (II-1-8) there are extra k's compared to Haskell's (1964) definitions. The functions an extra k are those which possess $\frac{dJ_n(kr)}{dkr}$ in the corresponding equations (II-1-6) and (II-l-7). $J_n(kr)$ or $J_{n-1}(kr)$ are characteristic functions in the r direction, these k's make the functions U_r , U_z ,

 U_v have dimensions of displacement. Also U_τ differs by a minus sign and T_v by ω^2 when compared to Haskell (1964) or Wang and Herrmann (1980). Such a particular definition will allow the layer matrix α and all other Haskell's matrices defined below to be more symmetric.

From equation (II-1-8), it is easy to find ordinary differential equations for these functions. Expressing equation (II-1-8) in matrix form, we have

$$\begin{bmatrix} U_r \\ U_z \\ T_z \\ T_v \end{bmatrix} = \begin{bmatrix} 0 & -\frac{k}{\rho} & -\frac{k}{\rho} & 0 & 0 & 0 \\ \frac{1}{\rho} & 0 & 0 & \frac{k^2}{\rho} & 0 & 0 \\ 0 & (\gamma - 1) & \gamma & 0 & 0 & 0 \\ \frac{\gamma}{k} & 0 & 0 & (\gamma - 1)k & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & \frac{k}{\rho} \\ T_v \end{bmatrix} \begin{bmatrix} \frac{dZ_1}{dz} \\ Z_1 \\ \frac{dZ_2}{dz} \\ Z_2 \\ \frac{dF_3}{dz} \\ F_3 \end{bmatrix}$$
(II-1-9)

The inverse of this equation is

$$\begin{bmatrix} \frac{dZ_{1}}{dz} \\ Z_{1} \\ \frac{dZ_{2}}{dz} \\ Z_{2} \\ \frac{dF_{3}}{dz} \\ F_{3} \end{bmatrix} = \begin{bmatrix} 0 & -\rho(\gamma-1) & 0 & k & 0 & 0 \\ -\rho\frac{\gamma}{k} & 0 & -1 & 0 & 0 & 0 \\ \rho\frac{(\gamma-1)}{k} & 0 & 1 & 0 & 0 & 0 \\ 0 & \rho\frac{\gamma}{k^{2}} & 0 & -\frac{1}{k} & 0 & 0 \\ 0 & 0 & 0 & 0 & \frac{\rho}{\mu k} \\ 0 & 0 & 0 & 0 & \frac{\rho}{k} & 0 \end{bmatrix} \begin{bmatrix} U_{r} \\ U_{z} \\ T_{z} \\ U_{\vartheta} \\ T_{\vartheta} \end{bmatrix}$$
(II-1-10)

If we take the depth derivative on both sides of equation (II-1-9), replace the second z-derivatives of Z_1 ,

 Z_2 , and F_3 by equation (II-1-5), combine with equation (II-1-10), and use $\rho\alpha^2=\lambda+2\mu$, $\rho\beta^2=\mu$, we find that

$$\frac{d}{dz} \begin{bmatrix} U_r \\ U_z \\ T_z \\ T_r \end{bmatrix} = \begin{bmatrix} 0 & k & 0 & -\frac{\omega^2}{\mu} & 0 & 0 \\ -k\sigma & 0 & \omega^2 \frac{\sigma}{\lambda} & 0 & 0 & 0 \\ 0 & -\rho & 0 & k & 0 & 0 \\ \rho - \xi \frac{k^2}{\omega^2} & 0 & -k\sigma & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & \frac{1}{\mu} \\ T_v \end{bmatrix} \begin{bmatrix} U_r \\ U_z \\ T_z \\ T_r \end{bmatrix} \tag{II-1-11}$$

where

$$\xi = \frac{4 \,\mu \,(\lambda + \mu)}{\lambda + 2 \,\mu} \qquad \sigma = \frac{\lambda}{\lambda + 2 \,\mu} \,.$$

The matrix on the right hand side of this equation is denoted by A. Note that A is a skew-symmetric matrix, i.e., symmetric about the secondary diagonal axis. Such symmetric properties will reappear later. Actually, one part of the differential equations in (II-1-11) comes from the equation of motion and the other from the relationship between displacement and stress. The differential equation (II-1-11) forms the basis of the propagator matrix theory of Gilbert and Backus (1966) and the starting point for solving the eigenfunction problem by numerical integration (Takeuchi and Saito, 1972).

The z components of potentials satisfying equation (II-1-5) have the solutions

$$Z_{1} = A' e^{-\nu_{\alpha} z} + A'' e^{\nu_{\alpha} z}$$

$$Z_{2} = B' e^{-\nu_{\beta} z} + B'' e^{\nu_{\beta} z}$$

$$F_{3} = C' e^{-\nu_{\beta} z} + C'' e^{\nu_{\beta} z},$$
(II-1-12)

where the single primes represent the waves propagating in the positive z direction, i.e., downward, and the double primes for waves in the negative direction, i.e., upward. Some authors normalize the coefficients in equation (II-1-12) by the energy flux in the z direction (Kennett et al, 1978), which is $\sqrt{2\nu_\alpha/\rho}$ for A and $\sqrt{2k^2\nu_\beta/\rho}$ for B and C. Normalization is not used here, but this point should be remembered in reference to the reflection and transmission coefficients. After substituting Z_1 , Z_2 , F_3 into equation (II-1-8), the U's and T's in the layer m can be expressed by the following matrix:

$$\begin{bmatrix} U_{r} \\ U_{z} \\ T_{z} \\ T_{r} \\ U_{\vartheta} \end{bmatrix}_{m+1} = \begin{bmatrix} -\frac{k}{\rho} & -\frac{\nu_{\beta}k}{\rho} & -\frac{k}{\rho} & \frac{\nu_{\beta}k}{\rho} & 0 & 0 \\ \frac{\nu_{\alpha}}{\rho} & \frac{k^{2}}{\rho} & -\frac{\nu_{\alpha}}{\rho} & \frac{k^{2}}{\rho} & 0 & 0 \\ \frac{(\gamma-1)}{\gamma\nu_{\beta}} & (\gamma-1) & -\gamma\nu_{\beta} & 0 & 0 \\ \frac{\gamma\nu_{\alpha}}{k} & (\gamma-1)k & -\frac{\gamma\nu_{\alpha}}{k} & (\gamma-1)k & 0 & 0 \\ 0 & 0 & 0 & 0 & \frac{k}{\rho} & \frac{k}{\rho} \\ 0 & 0 & 0 & 0 & \frac{\mu k \nu_{\beta}}{\rho} & -\frac{\mu k \nu_{\beta}}{\rho} \end{bmatrix}_{m} diag \begin{bmatrix} e^{\nu_{\alpha}z} \\ e^{\nu_{\beta}z} \\ e^{-\nu_{\alpha}z} \\ e^{-\nu_{\beta}z} \\ e^{-\nu_{\beta}z} \end{bmatrix}_{m} \begin{bmatrix} A'' \\ B'' \\ B' \\ C'' \\ C' \end{bmatrix}_{m}$$

$$(II-1-13)$$

where the subscript m is the layer index. We have denoted the quantities at the top of the m'th layer by

m and at the bottom by m+1. From equation (II-1-11) and (II-1-13), it is obvious that the v component can be separated from the r,z components. This represents the SH wave with particle motion parallel to the interfaces. For ease of expression, we will keep the three transformed components of motion together.

It is convenient to introduce the matrices

$$\mathbf{B} = \begin{bmatrix} U_r, U_z, T_z, T_r, U_{\vartheta}, T_{\vartheta} \end{bmatrix}^T$$

$$K = [A'', B'', A', B', C'', C']^T$$

E for the first matrix, and Λ for the diagonal second matrix on the right hand side of equation (II-1-13). Note that the matrix Λ describes phase variation along the depth direction. With these substitutions, equation (II-1-13) becomes

$$\mathbf{B}_{m+1} = \mathbf{E}_m \, \Lambda_m(z) \, \mathbf{K}_m . \tag{II-1-14}$$

In equation (II-1-14) we have one of the most important equations of this dissertation. Matrix B , the eigenfunction vector is also called the motion-stress vector (Aki and Richards, 1980) and has the convenient property interfaces except the of being continuous across the source. Matrix K is called the potential-constant vector which is composed of potential coefficients from equation (II-1-12). These coefficients are constant everywhere inside a homogeneous layer. In the reflectivity method, the vector plays an important role.

Let B_m be the value of B at the top of the m'th layer and B_{m+1} be its value at the bottom of this layer. Taking $z=\emptyset$ in equation (II-1-14), we have

$$\mathbf{B}_m = \mathbf{E}_m \quad \mathbf{K}_m \quad , \tag{II-1-15}$$

and taking $z = d_m$, we have

$$\mathbf{B}_{m+1} = \mathbf{E}_m \ \Lambda_m(d_m) \ \mathbf{K}_m \ . \tag{II-1-16}$$

It is important to mention that although K_m is constant in the m'th layer, it can be premultiplied by the matrix Λ_m to become 'propagatable' when it is used to describe waves inside the layer. Therefore $\Lambda_m(0) K_m$ represents the potential at the top of layer m and $\Lambda_m(d_m) K_m$ is the potential at the bottom of this layer. By combining equations (II-1-15) and (II-1-16), we obtain a useful formula:

$$B_{m+1} = E_m \Lambda_m (E_m^{-1} B_m)$$

= $(E_m \Lambda_m E_m^{-1}) B_m = a_m B_m$, (II-1-17)

where we have defined

$$\alpha_m = \mathbf{E}_m \ \Lambda_m \ \mathbf{E}_m^{-1} \ , \tag{II-1-18}$$

and E_m^{-1} is

$$\mathbf{E}_{m}^{-1} = \frac{1}{2} \begin{bmatrix} -\rho \frac{\gamma}{k} & -\rho \frac{(\gamma-1)}{\nu_{\alpha}} & -1 & \frac{k}{\nu_{\alpha}} & 0 & 0\\ \rho \frac{(\gamma-1)}{\nu_{\beta}k} & \rho \frac{\gamma}{k^{2}} & \frac{1}{\nu_{\beta}} & -\frac{1}{k} & 0 & 0\\ -\rho \frac{\gamma}{k} & \rho \frac{(\gamma-1)}{\nu_{\alpha}} & -1 & -\frac{k}{\nu_{\alpha}} & 0 & 0\\ -\rho \frac{(\gamma-1)}{\nu_{\beta}k} & \rho \frac{\gamma}{k^{2}} & -\frac{1}{\nu_{\beta}} & -\frac{1}{k} & 0 & 0\\ 0 & 0 & 0 & 0 & \frac{\rho}{k} & \frac{\rho}{\mu k \nu_{\beta}} \\ 0 & 0 & 0 & 0 & \frac{\rho}{k} & -\frac{\rho}{\mu k \nu_{\beta}} \end{bmatrix}.$$
(II-1-19)

Matrix α transfers the motion-stress vector through the layer; hence it is usually referred to as the layer matrix or matrix propagator. The elements of α are defined in Appendix A. This matrix possesses several interesting properties:

- (1) $\alpha(z)$ is a function of z only, i.e., it propagates in the vertical direction, and this z-dependence arises solely from the diagonal matrix $\Lambda(z)$.
- (2) $a^{-1}(z) = a(-z)$, since

$$a^{-1}(z) = [E \Lambda (z) E^{-1}]^{-1}$$

$$= E \Lambda^{-1} (z) E^{-1}$$

$$= E \Lambda (-z) E^{-1}$$

$$= a(-z).$$

(3) $a_{ij}^{-1}=(-1)^{i+j}\,a_{ij}$, this can be seen from Appendix A where terms with i+j= odd are related to $\sinh(\nu_{\alpha,\beta}z)$ which is an odd function, and terms with i+j= even are related to $\cosh(\nu_{\alpha,\beta}z)$ which is an even function.

(4) In a homogeneous layer,

$$a_m(z_1+z_2) = a_m(z_1) a_m(z_2),$$

since the diagonal matrix Λ can be decomposed:

$$\alpha_{m}(z_{1}+z_{2}) = E_{m} \Lambda_{m}(z_{1}+z_{2}) E_{m}^{-1}$$

$$= E_{m} \Lambda_{m}(z_{1}) \Lambda_{m}(z_{2}) E_{m}^{-1}$$

$$= E_{m} \Lambda_{m}(z_{1}) E_{m}^{-1} E_{m} \Lambda_{m}(z_{2}) E_{m}^{-1}$$

$$= \alpha_{m}(z_{1}) \alpha_{m}(z_{2}).$$

This is the property of a propagator matrix (Gilbert and Backus, 1966).

- (5) All of the elements of matrix α are real for real k and ω , and most importantly the sinh functions always appear as $\sinh(\nu_{\alpha,\beta}z)/\nu_{\alpha,\beta}$ or $\sinh(\nu_{\alpha,\beta}z)\cdot\nu_{\alpha,\beta}$ which suppress the possible branch points due to the multivalued functions $\nu_{\alpha,\beta}$.
- (6) α has a type of symmetry

$$a_{ij} = a_{5-j,5-i}$$
 for $P-SV$

$$a_{55} = a_{66}$$
 for SH

Properties (3) and (6) can be combined as

$$a_{ij}^{-1} = (-1)^{i+j} a_{5-j,5-i}$$
 for $P-SV$

$$a_{ij}^{-1} = (-1)^{i+j} a_{11-j,11-i}$$
 for SH . (II-1-20)

In a similar way, the potential-constant vector can be transferred across the layer boundary by

$$K_{m+1} = E_{m+1}^{-1} E_m \Lambda_m(d_m) K_m$$
 (II-1-21)

 $\Lambda_m(d_m) \cdot K_m$ is the potential located just above the m+l interface and K_{m+1} just below this interface. Hence, $\mathbf{E}_{m+1}^{-1} \cdot \mathbf{E}_m$ contains the information about the reflection and transmission of waves passing through the m+l interface. This property was extensively used by Kennett (1974) in his study of reflection and transmission coefficients.

the above discussion, we constructed displacement-stress field in a layered medium with the aid of the motion-stress vector B which in some represents the character of the boundaries. To connect these vectors between layers, a potential-constant vecwas defined in a more direct way than Haskell (1953,1964). Haskell (1964) defined K = [A' + A'', A' - A'', B' - A'']B''. B'+B'' $]_{\bullet}^{T}$ This quantity is not convenient for application and was replaced by FK in Wang and Herrmann (1980).The potential-constant vector represents the waves inside the layers. With this new K we are able find a diagonal matrix Λ , and consequently a better form for the layer matrix $\alpha = E \wedge E^{-1}$. Matrix E consists of eigenvectors of matrix A in the differential equation (II-1-11), which will be further explored section 2.3. It is noted that use of the matrix D of Haskell (1964) and matrix F of Wang and Herrmann (1980) has been avoided.

2.2 Compound Matrix

The compound matrix theory was first introduced to seismology' by Dunkin (1965). This theory was proposed to treat the problem of loss of precision during calculation of the layer matrix in Haskell's (1953) original theory. Knopoff (1964) also developed a method using matrix representation to solve this problem. However, because of complexity in notation his method was widely applied. The reason that Haskell's layer matrix needed to be extended to a compound matrix arises exponential terms contained in matrix Λ . During the calculation, these exponential terms, when happen to be real functions as in the case of surface waves, will grow very large to obscure the significant figures of other factors. The remedy for this is to control the exponential values during the computation these elements and/or to reformulate the matrix theory of Thomson (1950) or Haskell (1953) using In this section, we will discuss pound matrix forms. the compound matrix in detail and also consider source representation to obtain complete solutions. The source function will be further explored in chapter V.

From Hudson (1969a) the source effect can be taken into account by means of a discontinuity of motion-stress vector across the source depth. Suppose that a

source is in the m'th layer at a depth h_m beneath the m interface. The source vector S is defined as

$$S = B_m^+ - B_m^- \, . \tag{II-2-1}$$

where B_m^- , B_m^+ are the motion-stress vectors immediately above and below the source depth $z_m + h_m$, respectively. Such a source definition is different from that of Wang and Herrmann (1980), which is a modification of Haskell (1964), in that the S they used comes from the discontinuity of potential-constant vector K , i.e., $\Sigma = K_m^+ - K_m^-$. The relation of these two expression is just

$$S = E_m \Sigma . (11-2-2)$$

Haskell's expression for the source is also useful when applied to other theories such as the reflectivity method (Fuchs, 1968). The introduction of revised source terms has the advantage that the factors ν_{α} and ν_{β} are no longer required.

It is straightforward to relate the motion-stress vector $\mathbf{B_i}$ at the surface to $\mathbf{B_m^-}$ by the layer matrices in between:

$$B_m = \alpha_m(h_m) \alpha_{m-1} \cdots \alpha_1 B_1 = ZB_1$$
. (II-2-3)

Similarly, B_m^+ can be related to K_N in the half-space by

$$K_N = E_N^{-1} \alpha_{N-1} \cdots \alpha_m (d_m - h_m) B_m^+ = XB_m^+.$$
 (II-2-4)

Equations (II-2-1), (II-2-3), and (II-2-4) are further combined as

$$K_N = XB_m^+ = XS + XZB_1 = XS + RB_1$$
 (II-2-5)

where

$$X = E_N^{-1} \alpha_{N-1} \cdots \alpha_m (d_m - h_m)$$

$$Z = \alpha_m (h_m) \cdots \alpha_1$$

$$R = X Z = E_N^{-1} \alpha_{N-1} \cdots \alpha_1$$
(II-2-6)

We have used property (4) of matrix α (page 24) to define matrix R.

Now we consider the boundary conditions at the free surface and the half-space. Vanishing of stresses at the free surface requires B_1 to be

$$\mathbf{B_{1}} = [\ U_{r_{1}}\,,\,U_{z_{1}}\,,\,0\,,\,0\,,\,U_{\vartheta_{1}}\,,\,0\,\,]^{T} \ .$$

Inasmuch as the source is assumed to be in one of the layers, there can be no upwardly propagating waves in the half-space. Thus $A_N^{"}=B_N^{"}=C_N^{"}=0$, and K_N takes the form

$$\mathbb{K}_{N} = [0, 0, A_{N}, B_{N}, 0, C_{N}]^{T}.$$

The first two and the fifth rows in matrix equation (II-2-5) can be written as

$$\begin{bmatrix} 0 \\ 0 \end{bmatrix} = \begin{bmatrix} X_{11} & X_{12} & X_{13} & X_{14} \\ X_{21} & X_{22} & X_{23} & X_{24} \end{bmatrix} \begin{bmatrix} S_1 \\ S_2 \\ S_3 \\ S_4 \end{bmatrix} + \begin{bmatrix} R_{11} & R_{12} \\ R_{21} & R_{22} \end{bmatrix} \begin{bmatrix} U_{r_1} \\ U_{z_1} \end{bmatrix}$$

$$0 = \left[X_{55}, X_{56} \right] \begin{bmatrix} S_5 \\ S_6 \end{bmatrix} + R_{55} U_{\vartheta_1}.$$

Here the P-SV and the SH components are separated. The free surface displacements are easily found to be

$$\begin{bmatrix} U_{\tau} \\ U_{z} \end{bmatrix}_{1} = (-1) \begin{bmatrix} R_{22} & -R_{12} \\ -R_{21} & R_{11} \end{bmatrix} \begin{bmatrix} X_{1i} & S_{i} \\ X_{2i} & S_{i} \end{bmatrix} / R \mid_{12}^{12} \qquad i = 1, \dots, 4$$

$$U_{\vartheta_{1}} = (-1) (X_{5j} S_{j}) / R_{55} \qquad j = 5, 6 ,$$
(II-2-7)

where the use of summation convention for the subscripts and the compound matrix or second order subdeterminant definition $R \mid_{kl}^{ij} = R_{ik}R_{jl} - R_{il}R_{jk}$ are understood. This is the extent of Haskell's (1964) development. Let us consider the P-SV waves further. Since R = XZ, i.e., $R_{ij} = X_{ik}Z_{kj}$, equation (II-2-7) is further developed as

$$\begin{bmatrix} U_r \\ U_z \end{bmatrix}_1 = \begin{bmatrix} -S_i \ X \mid_{ij}^{12} \ Z_{j2} \\ S_i \ X \mid_{ij}^{12} \ Z_{j1} \end{bmatrix} / R \mid_{12}^{12} . \tag{II-2-8}$$

A compound matrix can be expressed as the matrix product of its compound sub-matrices (Dunkin, 1965). Hence from equation (II-2-6)

$$X \mid_{ij}^{12} = E_N^{-1} \mid_{mn}^{12} \alpha_{N-1} \mid_{op}^{mn} \cdots \alpha_{m+1} \mid_{si}^{qr} \alpha_m (d_m - h_m) \mid_{ij}^{st}$$
(II-2-9)

$$R \mid_{12}^{12} = E_N^{-1} \mid_{mn}^{12} \cdots a_2 \mid_{st}^{gr} a_1 \mid_{12}^{st}$$
.

The elements of compound matrices $E_N^{-1}|_{ij}^{12}$ and $a|_{kl}^{ij}$ are listed in Appendix B. It is important to point out that in deriving $a \mid_{k}^{n}$ from matrix a_{ij} we are able to remove the questionable square exponential terms by the analytic equality $\cosh^2(\nu z) - \sinh^2(\nu z) = 1$. This basic reason that the compound matrix is better than the simple matrix in Haskell's theory. In Appendix B, we have factored out the quantity $1/(4k\, \nu_{\alpha_N} \nu_{\beta_N})$ when defining $E_N^{-1}|_{ii}^{12}$ to remove the possible singularities from $u_{\alpha N}$ and $u_{\beta N}$. This quantity is usually canceled out by division of similar terms in the denominator of tion (II-2-8). Therefore it will be suppressed in our calculation. However, in other cases, this term might be discarded so easily, such as when considering the stresses (Baumgardt, 1980). This point should be noted when the compound matrix $E_N^{-1}|_{ii}^{12}$ is included in the formulation.

The compound matrix $a \mid_{kl}^{ij}$ can be looked at as a 6 by 6 matrix, if we assign the indices as

$$12 \rightarrow 1$$
 $13 \rightarrow 2$ $14 \rightarrow 3$ $23 \rightarrow 4$ $24 \rightarrow 5$ $34 \rightarrow 6$.

In Appendix B, we find that the third and fourth rows or columns of compound matrix $a \mid_{kl}^{ij}$ are equivalent in a particular way. This can be explained mathematically in Appendix C by means of two properties of compound matrices. As indicated by Watson (1970), this equivalence permits us to reduce the 6 by 6 matrix to a

5 by 5 matrix by discarding the third column and row and replacing the fourth row by $\begin{bmatrix} 2a \mid_{12}^{23}, 2a \mid_{13}^{23}, 2a \mid_{23}^{23} - 1, 2a \mid_{24}^{23}, 2a \mid_{34}^{23} \end{bmatrix}$ during the matrix multiplication. The $E_{N}^{1} \mid_{12}^{12}$ matrix also drops the third component $E_{N}^{1} \mid_{14}^{12}$. Such a reduction saves some storage space and computer time. The third component dropped can be retrieved easily, since it is exactly the same as the old fourth component.

Because of the symmetry exhibited by matrices α and E^{-1} , there exist some interesting properties of their compound forms. As shown in Appendix C and equation (II-1-20), the compound matrices X and R , which consist of E^{-1} and α , possess the following properties:

$$Y^{-1}|_{34}^{ij} = (-1)^{i+j+1} q_N Y|_{5-j,5-i}^{12}$$

$$Y|_{14}^{12} = Y|_{23}^{12}$$

$$Y^{-1}|_{34}^{14} = Y^{-1}|_{34}^{23}$$
(II-2-10)

where Y might be E⁻¹ , X , or R , and the constant $q_N = \frac{4k^2\nu_{\alpha N}\nu_{\beta N}}{\rho_N^2} \ .$ This constant consists of the normalization factors we have mentioned in equation (II-1-12). Matrix Z , which consists of α 's , also has the properties:

$$\begin{split} Z_{ij}^{-1} &= (-1)^{i+j} \; Z_{5-j,5-i} \\ Z^{-1}|_{kl}^{ij} &= (-1)^{i+j+k+l} \; Z|_{5-j,5-i}^{5-l,5-k} \; . \end{split}$$

These symmetry properties are a characteristic of

Haskell's formulation, although they arise from a somewhat arbitrary choice of motion-stress and potentialconstant vectors in the last section. In Haskell's work (1964) or other similar formulations, these properties are more or less missing. These properties will be of particular interest when relating the eigenfunction theory and matrix theory in the next chapter.

The surface displacements in the forms of equation (II-2-8) are stable for calculation, since they include $X \mid \mathbb{R}$ which imposes the boundary conditions from halfspaces all the time. However, Harkrider (1964) and Harvey (1981) used another form. Since R = XZ or equivalently $X = RZ^{-1}$ we have

$$\begin{split} X \mid_{ij}^{12} Z_{jp} &= \left(X_{1i} X_{2j} - X_{1j} X_{2i} \right) Z_{jp} \\ &= R_{1k} Z_{ki}^{-1} R_{2l} \left(Z_{lj}^{-1} Z_{jp} \right) - R_{1m} \left(Z_{mj}^{-1} Z_{jp} \right) R_{2k} Z_{ki}^{-1} \\ &= R_{1k} Z_{ki}^{-1} R_{2p} - R_{1p} R_{2k} Z_{ki}^{-1} \\ &= R \mid_{kp}^{12} Z_{ki}^{-1} \end{split} .$$

Hence

$$\begin{bmatrix} U_r \\ U_z \end{bmatrix}_i = (-1) \begin{bmatrix} R \mid_{k^2}^{12} Z_{ki}^{-1} S_i \\ -R \mid_{k^1}^{12} Z_{ki}^{-1} S_i \end{bmatrix} / R \mid_{12}^{12} . \tag{II-2-11}$$

Numerically this form is not as good as equation (II-2-8), since it requires more calculations with $R \mid_{ij}^{12}$ than with $X \mid_{ij}^{12}$. However it is a useful expression for deriving other properties.

At this point, it is interesting to consider the

dispersion property of surface waves. It is known that the dispersion is determined by the velocity structure and is independent of sources. Assuming S=0 and $K_N=R\,B_1$, we obtain

$$\begin{bmatrix} 0 \\ 0 \\ A' \\ B' \end{bmatrix} = \begin{bmatrix} R_{11} & R_{12} & \dots \\ R_{21} & R_{22} & \dots \\ \dots & \dots & \dots \end{bmatrix} \begin{bmatrix} U_{r_1} \\ U_{z_1} \\ 0 \\ 0 \end{bmatrix}.$$

The first two rows give

$$\frac{U_{r_1}}{U_{z_1}} = -\frac{R_{12}}{R_{11}} = -\frac{R_{22}}{R_{21}}$$

and the last equality defines the period equation $R \mid_{12}^{12} = 0$. U_{r_1} / U_{z_1} is called the ellipticity of Rayleigh waves at the free surface, which is usually denoted by ε . By applying the arithmetic equality

$$\frac{b}{a} = \frac{d}{c} = \frac{bx - dy}{ax - cy}$$

where \boldsymbol{x} and \boldsymbol{y} are suitably arbitrary numbers, it is not difficult to find that

$$\mathcal{E} \equiv \frac{U_{r_1}}{U_{z_1}} = -\frac{R|_{23}^{12}}{R|_{13}^{12}} = -\frac{R|_{24}^{12}}{R|_{14}^{12}} \tag{II-2-12}$$

(This is the derivation of Equation 20 in Harkrider (1970)). Furthermore, another expression exists for the ellipticity in terms of R^{-1} . By using the symmetry properties of compound R (equation (II-2-10)), it is found that

$$\varepsilon = \frac{R^{-1} \begin{vmatrix} 13\\ 34 \end{vmatrix}}{R^{-1} \begin{vmatrix} 23\\ 34 \end{vmatrix}} = \frac{R^{-1} \begin{vmatrix} 14\\ 34 \end{vmatrix}}{R^{-1} \begin{vmatrix} 24\\ 34 \end{vmatrix}}.$$

The period equation in this case is $R^{-1}|_{34}^{34} = 0$.

The solutions in equation (II-2-8) include the source terms S_i . For the surface waves in which $R|_{12}^{12}=0$, the ellipticity of Rayleigh waves is known to be independent of the source. This can be shown by another form of the solution in equation (II-2-11):

$$\mathcal{E} = \frac{U_{r_1}}{U_{z_1}} = -\frac{R \mid_{k_2}^{12} Z_{k_1}^{-1} S_i}{R \mid_{k_1}^{12} Z_{k_1}^{-1} S_i}$$

$$= -\frac{R \mid_{23}^{12} \left(Z_{3i}^{-1} S_i + \frac{R \mid_{24}^{12}}{R \mid_{23}^{12}} Z_{4i}^{-1} S_i \right)}{R \mid_{13}^{12} \left(Z_{3i}^{-1} S_i + \frac{R \mid_{14}^{12}}{R \mid_{14}^{12}} Z_{4i}^{-1} S_i \right)}.$$
(II-2-13)

Using an equality for any compound matrix $R \mid_{\psi}^{12}$ (Abramovici, 1968),

$$R \mid_{12}^{12} R \mid_{34}^{12} + R \mid_{14}^{12} R \mid_{23}^{12} = R \mid_{13}^{12} R \mid_{24}^{12}$$

however since $R \mid_{12}^{12} = 0$, equation (II-2-13) becomes

$$\varepsilon = -\frac{R \mid_{23}^{12}}{R \mid_{13}^{12}}$$

Thus we have shown that the ellipticity is really a quantity determined from the layer response only and is independent of the source.

The solutions for the Fourier transformed displacements as expressed in equation (II-1-6), after

summing the mode numbers over horizontal and azimuthal directions, have the forms

$$\begin{split} u_z(r,\vartheta,0,\omega) &= \sum_{n=0}^{\infty} \int_0^{\infty} dk \; \{\; (U_z^{nc} \cos n\vartheta + U_z^{ns} \sin n\vartheta) \; J_n(k\tau) \, / \, F_R \; \} \\ -u_r(r,\vartheta,0,\omega) &= \sum_{n=0}^{\infty} \int_0^{\infty} dk \; \{\; (U_r^{nc} \cos n\vartheta + U_r^{ns} \sin n\vartheta) \; J_{n-1}(k\tau) \, / \, F_R \\ &- (\frac{n}{k\tau}) \; (U_r^{nc} \cos n\vartheta + U_r^{ns} \sin n\vartheta) \; J_n(k\tau) \, / \, F_R \\ &+ (\frac{n}{k\tau}) \; (U_r^{ns} \cos n\vartheta - U_{\vartheta}^{nc} \sin n\vartheta) \; J_n(k\tau) \, / \, F_L \; \} \\ -u_{\vartheta}(r,\vartheta,0,\omega) &= \sum_{n=0}^{\infty} \int_0^{\infty} dk \; \{\; (U_{\vartheta}^{ns} \cos n\vartheta - U_{\vartheta}^{nc} \sin n\vartheta) \; J_n(k\tau) \, / \, F_L \\ &- (\frac{n}{k\tau}) \; (U_r^{ns} \cos n\vartheta - U_{\vartheta}^{nc} \sin n\vartheta) \; J_n(k\tau) \, / \, F_L \\ &+ (\frac{n}{k\tau}) \; (U_r^{ns} \cos n\vartheta - U_{\vartheta}^{nc} \sin n\vartheta) \; J_n(k\tau) \, / \, F_R \; \} \;\; , \end{split}$$

where $U_z^{\text{nc,s}}$, $U_z^{\text{nc,s}}$, $U_z^{\text{nc,s}}$, F_R and F_L are

$$\begin{aligned} U_{z}^{nc,s} &= S_{i}^{nc,s} X \big|_{ij}^{12} Z_{j1} \\ U_{r}^{nc,s} &= - S_{i}^{nc,s} X \big|_{ij}^{12} Z_{j2} \\ U_{0}^{nc,s} &= - X_{5j} S_{j}^{nc,s} \\ F_{R} &= R \big|_{12}^{12} \\ F_{L} &= R_{55} . \end{aligned}$$

There is now no doubt that an earthquake can be represented by a double-couple source without moment model (Aki and Richards, 1980, p.43). Haskell (1963) used $\bf n$ for the vector normal to the fault and $\bf f$ for the direction of force to describe the source. For practical applications, the dip $\bf d$, strike $\bf \varphi$ of the fault plane and the slip $\bf s$ of movement on this plane are usu-

ally preferred. The relation of these quantities to n and f are just

$$n_1 = -\sin d \sin \varphi$$
 $f_1 = \cos s \cos \varphi + \sin s \cos d \sin \varphi$ $n_2 = \sin d \cos \varphi$ $f_2 = \cos s \sin \varphi - \sin s \cos d \cos \varphi$ $n_3 = -\cos d$ $f_3 = -\sin s \sin d$.

The Fourier transformed displacements generated by such a dislocation source, for which $n = \emptyset, 1, 2$ in equation (II-2-14), have the forms:

$$u_{z}(r, \vartheta, 0, \omega) = ZSS \cdot R_{ss} + ZDS \cdot R_{ds} + ZDD \cdot R_{dd}$$

$$u_{r}(r, \vartheta, 0, \omega) = RSS \cdot R_{ss} + RDS \cdot R_{ds} + RDD \cdot R_{dd}$$

$$u_{\vartheta}(r, \vartheta, 0, \omega) = TSS \cdot R_{ss}' + TDS \cdot R_{ds}'$$
(II-2-15)

where

$$R_{ss} = (f_{1}n_{1} - f_{2}n_{2}) \cos 2\vartheta + (f_{1}n_{2} + f_{2}n_{1}) \sin 2\vartheta$$

$$= \sin d \cos s \sin 2(\vartheta - \varphi) + \frac{1}{2} \sin 2d \sin s \cos 2(\vartheta - \varphi)$$

$$R_{ds} = (f_{1}n_{3} + f_{3}n_{1}) \cos \vartheta + (f_{2}n_{3} + f_{3}n_{2}) \sin \vartheta$$

$$= \cos 2d \sin s \sin (\vartheta - \varphi) - \cos d \cos s \cos (\vartheta - \varphi)$$

$$R_{dd} = f_{3}n_{3}$$

$$= \frac{1}{2} \sin s \sin 2d$$

$$R_{ss}' = (f_{1}n_{2} + f_{2}n_{1}) \cos 2\vartheta - (f_{1}n_{1} - f_{2}n_{2}) \sin 2\vartheta$$

$$= \sin d \cos s \cos 2(\vartheta - \varphi) - \frac{1}{2} \sin 2d \sin s \sin 2(\vartheta - \varphi)$$

$$R_{ds}' = (f_{2}n_{3} + f_{3}n_{2}) \cos \vartheta - (f_{1}n_{3} + f_{3}n_{1}) \sin \vartheta$$

$$= \cos d \cos s \sin (\vartheta - \varphi) + \cos 2d \sin s \cos (\vartheta - \varphi)$$

$$ZSS = \int_{0}^{\infty} (1/F_{R}) \{ S_{i}^{2} X |_{ij}^{12} Z_{j1} \} J_{2}(kr) dk$$

$$ZDS = \int_{0}^{\infty} (1/F_{R}) \{ S_{i}^{1} X |_{ij}^{12} Z_{j1} \} J_{1}(kr) dk$$

$$ZDD = \int_{0}^{\infty} (1/F_{R}) \{ S_{i}^{2} X |_{ij}^{12} Z_{j1} \} J_{0}(kr) dk$$

$$RSS = \int_{0}^{\infty} (1/F_{R}) \{ S_{i}^{2} X |_{ij}^{12} Z_{j2} \} J_{1}(kr) dk$$

$$-\int_{0}^{\infty} (1/F_{R}) \{ S_{i}^{2} X |_{ij}^{12} Z_{j2} \} 2J_{2}(kr)/kr \ dk$$

$$+\int_{0}^{\infty} (1/F_{L}) \{ S_{j}^{6} X_{5j} \} 2J_{2}(kr)/kr \ dk$$

$$RDS = \int_{0}^{\infty} (1/F_{R}) \{ S_{i}^{1} X |_{ij}^{12} Z_{j2} \} J_{0}(kr) \ dk$$

$$-\int_{0}^{\infty} (1/F_{R}) \{ S_{i}^{1} X |_{ij}^{12} Z_{j2} \} J_{1}(kr)/kr \ dk$$

$$+\int_{0}^{\infty} (1/F_{L}) \{ S_{j}^{5} X_{5j} \} J_{1}(kr)/kr \ dk$$

$$RDD = -\int_{0}^{\infty} (1/F_{R}) \{ S_{i}^{0} X |_{ij}^{12} Z_{j2} \} J_{1}(kr) \ dk$$

$$TSS = \int_{0}^{\infty} (1/F_{L}) \{ X_{5j} S_{j}^{6} \} J_{1}(kr) \ dk$$

$$-\int_{0}^{\infty} (1/F_{L}) \{ X_{5j} S_{j}^{6} \} 2J_{2}(kr)/kr \ dk$$

$$+\int_{0}^{\infty} (1/F_{R}) \{ S_{i}^{2} X |_{ij}^{12} Z_{j2} \} 2J_{2}(kr)/kr \ dk$$

$$TDS = \int_{0}^{\infty} (1/F_{L}) \{ X_{5j} S_{j}^{5} \} J_{0}(kr) \ dk$$

$$-\int_{0}^{\infty} (1/F_{L}) \{ X_{5j} S_{j}^{5} \} J_{1}(kr)/kr \ dk$$

$$+\int_{0}^{\infty} (1/F_{R}) \{ S_{i}^{1} X |_{ij}^{12} Z_{j2} \} J_{1}(kr)/kr \ dk$$

$$+\int_{0}^{\infty} (1/F_{R}) \{ S_{i}^{1} X |_{ij}^{12} Z_{j2} \} J_{1}(kr)/kr \ dk$$

with

$$4\pi \rho_m \omega^2 S_2^0 = 4k \ k_{\alpha_m}^2 \qquad 4\pi \omega^2 S_4^0 = 2k^2 \left[(2\beta_m / \alpha_m)^2 - 3 \right]$$

$$4\pi \rho_m \omega^2 S_1^1 = -2k \ k_{\beta_m}^2 \qquad 4\pi \omega^2 S_4^2 = -2k^2 \qquad (II-2-16)$$

$$4\pi \rho_m \beta_m^2 S_5^1 = -2k \qquad 4\pi S_6^2 = 2k^2 .$$

All other S_i^n equal zero. The solutions and source forms for other types of sources will be given in chapter V.

In the above equations, SS represents a strikeslip type of source, DS represents a dip-slip type of source, and DD represents a 45° dip-slip type of source. It is apparent that these three components are all that are necessary to represent the P-SV motion shear-dislocation source (Harkrider, 1976). SH motion only needs SS and DS components. RSS, RDS, and in equation (II-2-15) also include the nearfield terms. These terms decrease faster with r than the others, and thus, are important only at short distances or low frequencies. The near-field terms in and t directions have exactly the same forms. worthwhile to indicate that the near-field terms have a higher order Bessel function than their far-field counterparts. The calculation of near-field terms requires special care. This will be studied in section 2.4.

2.3 Comparison with Other Formalisms

Many approaches have been developed to treat wave propagation in isotropic and homogeneous layered media. this section, we will discuss three typical the system conapproaches from the standpoint of structed in this dissertation. It will be shown system is related to these other approaches with only a little modification. Examples describing merit of our approach will be given later.

The Eigenvector of ODE

We have already mentioned that the eigenfunction theory plays an important role in establishing our matrix system. The mathematical foundation of matrix theory was provided by Gilbert and Backus (1966) using the 'propagator matrix' description. The mathematical connection is enlarged in the following discussion.

The motion-stress vector satisfies the first order differential equation (II-1-11):

$$\frac{d}{dz}B = AB, \qquad (II-3-1)$$

where A is a matrix representing the material property. Within a given layer A is a constant, and the solution takes the form

$$B(z) = e^{(z-z_0)\cdot A} B(z_0),$$

where z_0 is the reference depth. The function $e^{(s-z_0)\cdot A}$, called the matricant or layer matrix for a homogeneous layer, is defined by Sylvester's theorem (Hildebrand, 1965, p.61). Compared to equation (II-1-17), the layer matrix defined in equation (II-1-18) must be

$$\alpha = e^{(z-z_0)\Lambda} . \tag{II-3-2}$$

Applying matrix multiplication, with E^{-1} and E from equations (II-1-13) and (II-1-19), equation (II-3-2) becomes

$$\mathbf{E}^{-1} \boldsymbol{\alpha} \mathbf{E} = \mathbf{E}^{-1} e^{(z-z_0) \cdot \mathbf{A}} \mathbf{E} = e^{(z-z_0) \mathbf{E}^{-1} \mathbf{A} \mathbf{E}}$$
.

Since $\Lambda(z) = E^{-1}\alpha E$ (equation II-1-18), we obtain an interesting form:

$$\mathbf{E}^{-1}\mathbf{A}\mathbf{E} = \begin{bmatrix} \nu_{\alpha} & 0 & 0 & 0 & 0 & 0 \\ 0 & \nu_{\beta} & 0 & 0 & 0 & 0 \\ 0 & 0 & -\nu_{\alpha} & 0 & 0 & 0 \\ 0 & 0 & 0 & -\nu_{\beta} & 0 & 0 \\ 0 & 0 & 0 & 0 & \nu_{\beta} & 0 \\ 0 & 0 & 0 & 0 & 0 & -\nu_{\beta} \end{bmatrix} \equiv \mathbf{V} .$$

It is found that matrix A is diagonalized in a way similar to that of the characteristic-value problem. According to eigenvalue theory, E must be constructed from the eigenvectors of A and $\pm \nu_a$, $\pm \nu_\beta$ are its eigenvalues. This fact is confirmed as we actually obtain the eigenvalues from matrix A (Aki and Richards, 1980, p.275). Therefore the eigenvectors e_k 's of A constitute the successive columns of the E matrix:

The corresponding eigenvalues λ 's are $\nu_{\alpha}, \nu_{\beta}, -\nu_{\alpha}, -\nu_{\beta}, \nu_{\beta}, -\nu_{\alpha}$. The e_k 's are mutually orthogonal.

With this view of the E matrix, matrix B in

equation (II-3-1) can also be expressed in terms of eigenvectors and eigenvalues. Project vector B into the direction of \mathbf{e}_1 eigenvector, and denote this component by \mathbf{M}_1 :

$$\mathbf{M}_1 = \mathbf{B} \cdot \mathbf{e}_1 = f_1(z) \cdot \mathbf{e}_1 .$$

Because of the linear independence of e_k 's, M_1 still satisfies the differential equation (II-3-1):

$$\begin{split} &\frac{d}{dz} (f_1(z) e_1) = A (f_1(z) e_1) \\ &(\frac{d}{dz} f_1(z)) e_1 = (A \cdot e_1) f_1(z) = (\lambda_1 e_1) f_1(z) , \end{split}$$

i.e.,

$$\frac{d}{dz} f_1(z) = \lambda_1 f_1(z) .$$

Hence in the e direction, we have the solution

$$\mathbf{M}_1 = e^{\lambda_1(z-z_0)} \mathbf{e}_1 .$$

Repeating the same procedure for the other components, we obtain the solution to the ordinary differential equation (II-3-1):

This form of solution is called a fundamental matrix.

The propagator matrix is defined from it (Gilbert and Backus, 1966, eq.2.4):

$$\mathbb{C}(z-z_0) = \mathbb{M}(z) \, \mathbb{M}^{-1}(z_0) = \mathbb{E} \, \Lambda(z) \, [\mathbb{E} \, \Lambda(z_0)]^{-1}$$
$$= \mathbb{E} \, \Lambda(z) \, \Lambda(-z_0) \, \mathbb{E}^{-1}$$
$$= \mathbb{E} \, \Lambda(z-z_0) \, \mathbb{E}^{-1} .$$

The most general solution of equation (II-3-1) is described by a linear combination of the fundamental matrix ${\tt M}$,

$$B = MK = E \wedge K$$

where K is any constant vector. It follows that the differential equation (II-3-1) can be written in the form:

$$\frac{d}{dz}(\mathbb{E} \Lambda(z) \mathbb{K}) = \mathbb{A} (\mathbb{E} \Lambda(z) \mathbb{K})$$

$$\frac{d}{dz}(\Lambda(z) \mathbb{K}) = (\mathbb{E}^{-1} \mathbb{A} \mathbb{E}) \Lambda(z) \mathbb{K}$$

$$= \mathbb{V} (\Lambda(z) \mathbb{K}) , \qquad (II-3-3)$$

where we have used the fact that E and K are independent of z. Since V is a diagonal matrix, it is said that the wave field has been 'decoupled' (Claerbout, 1976, p.169). It is easy to see that vector K is just the potential-constant vector defined before, which consists of upgoing and downgoing waves. These waves flow up and down in homogeneous regions without interacting with each other.

Compound Matrix for High Frequencies

The compound matrix $X|_{ij}^{12}$ expressed in equation (II-2-9) appears as a (1X6) matrix during compound-compound multiplication (Dunkin, 1965):

$$X \mid_{ij}^{12} = [X \mid_{12}^{12}, X \mid_{13}^{12}, X \mid_{14}^{12}, X \mid_{23}^{12}, X \mid_{24}^{12}, X \mid_{34}^{12}]^T$$
.

When $X|_{ij}^{12}$ is multiplied by a regular matrix as in equation (II-2-8), it can be viewed as a (4X4) antisymmetric matrix:

$$X \mid_{ij}^{12} = \begin{bmatrix} 0 & X \mid_{12}^{12} & X \mid_{13}^{12} & X \mid_{14}^{12} \\ -X \mid_{12}^{12} & 0 & X \mid_{23}^{12} & X \mid_{24}^{12} \\ -X \mid_{13}^{12} & -X \mid_{23}^{12} & 0 & X \mid_{34}^{12} \\ -X \mid_{14}^{12} & -X \mid_{24}^{12} & -X \mid_{34}^{12} & 0 \end{bmatrix}$$

This property provides an alternative method of formulation.

In recent papers, Abo-Zena (1979) and Menke (1979) decomposed the layer matrix multiplication into a form with recursive relations:

$$Y_m = A_m^T \cdot \cdot \cdot A_{N-1}^T \{ [EA]^T [EB] - [EB]^T [EA] \} A_{N-1} \cdot \cdot \cdot A_m$$
 (II-3-4)

(Abo-Zena, equation 40), where $\{[EA]^T [EB] - [EB]^T [EA]\}$ is an anti-symmetric matrix, which is actually equivalent to our $E^{-1}|_{El}^{12}$. After a tedious expansion, we find that

$$Y \mid_{ij}^{12} A \mid_{kl}^{il} = A_{ki}^T Y \mid_{ij}^{12} A_{jl}$$
.

The recursive procedure (II-3-4) is exactly the same as the compound matrix multiplication we used to calculate R in equation (II-2-9).

Since layer matrix α can be factored into $E \wedge E^{-1}$, Abo-Zena further removed the phase terms in Λ out of α in order to control the exponentially growing terms when the phase velocity is less than a wave velocity (Abo-Zena, equation 44). A similar technique can also be found in Kennett's (1974) 'phase-related' operation (ref. section 4.3). In such a reformulation, the square exponential terms are identically zero, as we have seen in deriving the compound matrix α . It is this procedure which makes the matrix method useful for very high frequencies.

Mantle Wave Simulation

A scheme by which the teleseismic body wave pulses from a seismic source are calculated, allowing for the effects of transmission through the mantle and crust, was given by Carpenter (1966) and Hudson (1969b). The calculation is divided into three parts: source crust response, mantle effect, and receiver crust response. The rays emitted from the source crust are allowed to travel through the homogeneous upper mantle, then enter the receiver crust and be recorded. Hudson (1969b) gave a detailed derivation providing both formalisms

and approximating solutions. Here we will extend his result by using the compound matrix. Although both Carpenter (1966) and Hudson (1969b) start from Haskell (1964), our derivation is relatively simple and direct.

In the source crust, the waves from the source are reflected and refracted and finally enter the bottom halfspace as mantle waves following the equation:

$$K_N = XS + RB_1$$
.

After expanding, setting $A_N^{"}=B_N^{"}=0$ and $T_{r_1}=T_{z_1}=0$, and substituting equation (II-2-7), we have

$$A_N = [R \mid \frac{12}{12} X_{3i} S_i + R \mid \frac{29}{12} X_{1i} S_i - R \mid \frac{19}{12} X_{2i} S_i] / R \mid \frac{12}{12}$$

for P waves. This equation is equivalent to equation (3.3) of Hudson (1969b). Another form can easily be found by using

$$R^{-1}|_{kl}^{ij} = q_N(-1)^{i+j+k+l}R|_{ij}^{kl}$$
.

The result is

$$A'_{N} = S_{i} R^{-1} \begin{vmatrix} 34 \\ j4 \end{vmatrix} X_{ji} / R^{-1} \begin{vmatrix} 34 \\ 34 \end{vmatrix}$$

or, by $R^{-1} = Z^{-1} X^{-1}$

$$A_N' = S_i Z^{-1} \begin{vmatrix} 34 & X_{i4}^{-1} \\ 34 \end{vmatrix} R^{-1} \begin{vmatrix} 34 \\ 34 \end{vmatrix}$$
.

Similarly for SV waves, we have

$$\begin{split} B_N' &= -S_i \; R^{-1} |_{j3}^{34} \; X_{ji} \; / \; R^{-1} |_{34}^{34} \\ &= -S_i \; Z^{-1} |_{ij}^{34} \; X_{j3}^{-1} \; / \; R^{-1} |_{34}^{34} \; \; . \end{split}$$

Inside the mantle, the waves as expressed in displacement form are $B_{N+1}=E_N\;\Lambda_N\;K_N$

Or.

$$\begin{bmatrix} U_r \\ U_z \end{bmatrix}_{N+1} = \begin{bmatrix} -\frac{k}{\rho} e^{-\nu_{\alpha}z} A_N^{i} + \frac{\nu_{\beta}k}{\rho} e^{-\nu_{\beta}z} B_N^{i} \\ -\frac{\nu_{\alpha}}{\rho} e^{-\nu_{\alpha}z} A_N^{i} + \frac{k^2}{\rho} e^{-\nu_{\beta}z} B_N^{i} \end{bmatrix} = \begin{bmatrix} U_r^P + U_r^{SV} \\ U_z^P + U_z^{SV} \end{bmatrix}_{N+1},$$
(II-3-5)

where z is measured from the deepest interface. The solutions are divided into P and S components corresponding to A_N and B_N respectively. The final solutions similar to equation (II-2-15) can easily be found following the procedures in section 2.2.

When waves pass through the mantle, the effect of the mantle can be simply included by adding geometric spreading factors for P and S motions respectively (Ben-Menahem and Singh, 1972). The wave fields expressed in equation (II-3-5) are decomposed into the compressional and shear components by using the angle ϑ which is the angle made by the downward vertical with the ray direction:

$$U^{P} = U_{r}^{P} \sin \vartheta_{P} + U_{z}^{P} \cos \vartheta_{P}$$

$$U^{SV} = U_{r}^{SV} \cos \vartheta_{S} - U_{z}^{SV} \sin \vartheta_{S}$$

where

$$\vartheta_P = \tan^{-1} \frac{k}{\nu_{\alpha_N}}$$

$$\vartheta_S = \tan^{-1} \frac{k}{\nu_{\beta_N}}$$

 U^P and U^{SV} are then multiplied by the spreading factors of P and SV waves, respectively, to take into account

the mantle effect. At the bottom of the receiver crust, the P and SV waves can be resolved back into the r and z directions as follows:

$$\begin{split} U_r' &= U^{P'} \sin \vartheta_P - U^{SV'} \cos \vartheta_S \\ U_z' &= -U^{P'} \cos \vartheta_P - U^{SV'} \sin \vartheta_S \;, \end{split}$$

where the prime indicates the mantle waves which reach the receiver crust.

The effect of the receiver crust is easily taken into account by using the Haskell matrix R from the whole layer stacking. For our system, we just need to change $(J_{32}-J_{42})$ in Hudson's equation (7.7) to R_{22} , $(J_{31}-J_{41})$ to R_{21} , and evaluate the period equation by compound forms.

In the above discussion, we have shown how easy it is to extend our system. Many theories with difficult derivations come out simply by using the relation of Haskell matrices developed in section 2.1. More examples will be given in the later chapters.

2.4 Numerical Integration

Contour Integration

All of the integrals in equation (II-2-15) have the general form

$$\int_{0}^{\infty} f(k,\omega) J_{n}(kr) dk . \qquad (II-4-1)$$

The evaluation of this integral is complicated because the function $f(k,\omega)$ has poles and branch points (Figure 2). The poles come from the zeros of the period equation, and the branch points from the radicals $u_{\alpha N}$ and ν_{BN} given in equation (II-1-5). To avoid dual these radicals, we introduce the branch cuts at the loci along which the real parts of $\nu_{lpha N}$ and ν_{BN} are equal to zero. For the limiting case of real frequency and real layer velocities the two branch cuts, corresponding to $\nu_{\alpha N}$ and $\nu_{\beta N}$ respectively, collapse to form one branch cut along the real k-axis from the S branch point to the origin, then along the imaginary k-axis to $-i\infty$. The number of poles is finite and all them are located on the real k-axis. The positions of poles determine the dispersion values on the $\omega-k$ domain.

The integration in equation (II-4-1) can be evaluated by applying Cauchy's theorem. This is a standard technique and is used, for example, by Ewing et al (1957), Harkrider (1964), Hudson (1969b), and Herrmann (1979) in their work. In order to deform the contour properly we need to apply a transformation which expresses the Bessel function in terms of the Hankel functions of the first and second kind:

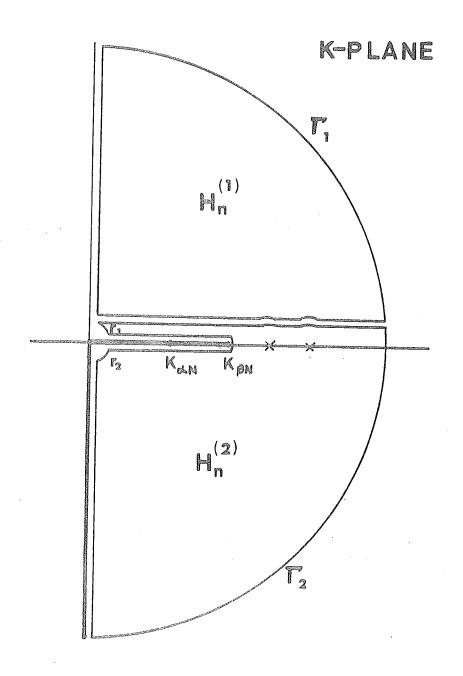


Figure 2. Contours in the complex k plane for evaluating the wavenumber integrals. The positions of the k_{α_N} and k_{β_N} branch points and the surface-wave poles ('X's') are indicated. Branch cuts are shown by thicker lines. Two circular arcs, γ_1 and γ_2 , sourround the possible Hankel function pole at k = 0.

$$J_n(kr) = \frac{1}{2} \left[H_n^{(1)}(kr) + H_n^{(2)}(kr) \right] .$$

The integral path can be deformed into the first and fourth quadrants, since $H_n^{(1)}$ and $H_n^{(2)}$ are analytic in the first and fourth quadrants, respectively, of the complex k plane. These two contours can further be combined into one as shown in Figure 2. Herrmann (1979) finally expressed the contour integration in the form:

$$\int_{0}^{\infty} f(k,\omega) J_{n}(kr) dk = -\pi i \sum Res \ f(k,\omega) H_{n}^{(2)}(kr)$$

$$+ i \int_{0}^{k} I_{m} f_{+}(k,\omega) H_{n}^{(2)}(kr) dk \qquad (II-4-2)$$

$$+ \left(\frac{1}{\pi}\right) \int_{0}^{\infty} \left[f_{+}(i\tau,\omega) \exp\left(-in\pi/2\right) \right] K_{n}(\tau r) d\tau ,$$

where the + or - subscripts indicate that ν_{α_N} , $\nu_{\beta_N} > \emptyset$ or < \emptyset , respectively, be used to evaluate the function. K_n is the modified Bessel function which decreases exponentially with increasing argument.

At each different frequency, when wavenumber k approaches zero, there might be some singularity introduced by the Hankel function. Wang and Herrmann (1980) realized that the solution (II-4-2) is correct only if $f(k,\omega) = O(k^{n+1})$ as k approaches zero where n is the order of the Bessel function. Otherwise the singularity of the Hankel function at k = 0 contributes to the

integral. After carefully examining the k terms in the response function, we find that such a Hankel singularity only occurs in the components of the near-field displacement of a double-couple source, for which $f(k,\omega) = O(k^0)$ for n = 1 and $f(k,\omega) = O(k^1)$ for n = 2.

Figure 2 shows the proper contour to be used for contour integration, where we have made a small circle to avoid the singularity at $k=\emptyset$. One cannot contract the small contour about the origin since the Hankel function singularity is still present at the lower limits of the branch line integrals. An interesting way to overcome this is to use a known integral whose behavior at $k=\emptyset$ is the same as that of the functions in equation (II-4-1). Using an integral from Harkrider (1976):

$$\int_{0}^{\omega} \frac{k_{v} J_{1}(kr) dk}{\nu_{v}} = [1 - exp(-ik_{v}r)] / ir ,$$

where $v_v^2(k,\omega) = k^2 - k_v^2$. The integral for n = 1 becomes

$$\begin{split} \int_{0}^{k} f(k,\omega) J_{1}(kr) \, dk &= -\pi i \sum Res \left[f(k,\omega) H_{1}^{(2)}(kr) \right] \\ &+ i \int_{0}^{k} I_{m} \left[f_{+}(k,\omega) + \frac{k_{v}}{\nu_{v}} I_{m} f_{+}(0,\omega) \right] H_{1}^{(2)}(kr) \, dk \\ &+ \frac{2}{\pi} \int_{0}^{\infty} I_{m} \left[f_{+}(i\tau,\omega) + \frac{k_{v}}{\nu_{v}} I_{m} f_{+}(0,\omega) \right] K_{1}(\tau r) \, d\tau \end{split}$$

$$+ \frac{1}{r} \operatorname{Re} f_{+}(0,\omega)$$

$$+ i \frac{I_{m} f_{+}(0,\omega)}{r} \left[1 - \exp(ik_{v}r)\right] .$$
(II-4-3)

To evaluate the near field integrals for n = 2, a careful examination of the integrands revealed that the integrand could easily be factored into the form $f(k,\omega)=kg(k,\omega) \quad . \quad \text{This contour integration used the other integral of Harkrider (1976):}$

$$\int_0^\infty \frac{k_v k J_2(kr)}{\nu_v} dk = -\frac{k_v}{r} exp(-ik_v r) + 2\left[1 - exp(ik_v r)\right] / ir^2.$$

Using this integral and performing the contour integration;

$$\begin{split} \int_{0}^{\cdot} f\left(k,\omega\right) J_{2}(kr) \; dk &= -\pi i \sum Res \; \left[k \; g\left(k,\omega\right) H_{2}^{(2)}(kr)\right] \\ &+ i \int_{0}^{\cdot} I_{m} \left[g_{+}(k,\omega) + \frac{k_{v}}{\nu_{v}} I_{m} g_{+}(0,\omega)\right] k \; H_{2}^{(2)}(kr) \; dk \\ &+ \frac{2}{\pi} \int_{0}^{\infty} I_{m} \left[g_{+}(i\tau,\omega) + \frac{k_{v}}{\nu_{v}(i\tau)} I_{m} g_{+}(0,\omega)\right] \tau K_{2}(\tau r) \; d\tau \\ &+ \frac{2}{\tau^{2}} \operatorname{Re} \; g_{+}(0,\omega) \\ &+ \frac{I_{m} g_{+}(0,\omega)}{\tau} \left[k_{v} \exp(-ik_{v} r) + i \frac{2}{\tau} (1 - \exp(-ik_{v} r))\right] \; . \end{split}$$

In both equations (II-4-3) and (II-4-4), the lower limits of integration equal zero because the functions $f(k,\omega)$ for n=1 and $g(k,\omega)$ for n=2 are even functions in k. Therefore the integrands in the branch line integrals are identically zero as the lower limit of

integration goes to zero. The choice of the parameter k_{ν} is somewhat arbitrary. For the SH terms, $k_{\nu}=k_{\beta_N}$ is used while $k_{\nu}=k_{\alpha}$ from equation (II-4-5) is used for the P-SV terms. This point will be further discussed later.

Numerical Integration

The solutions as shown in the integral form (II-4-2) consist of two contributions: the equation pole residue contribution and the branch line integrals. Generally speaking, the pole contribution gives rise to surface waves and the branch line integral yields most of the body waves. The pole and its residue will be discussed using normal mode theory the next chapter. Later in chapter IV, the branch line contribution will be examined using a leaky mode approach. Here we just give a brief discussion of these two contributions, and attempt to improve the numerical integration technique.

To evaluate the integral

$$\int_{0}^{k_{\beta_{N}}} \{ I_{m} f(k,\omega) \} H_{n}^{(2)}(kr) k^{m} dk ,$$

knowledge of the behavior of the function $I_m f(k,\omega)$ is essential. Figure 3 shows the variation of these functions along the real branch axis at a frequency of 1 Hz for a point source at a depth of 10 km in the central

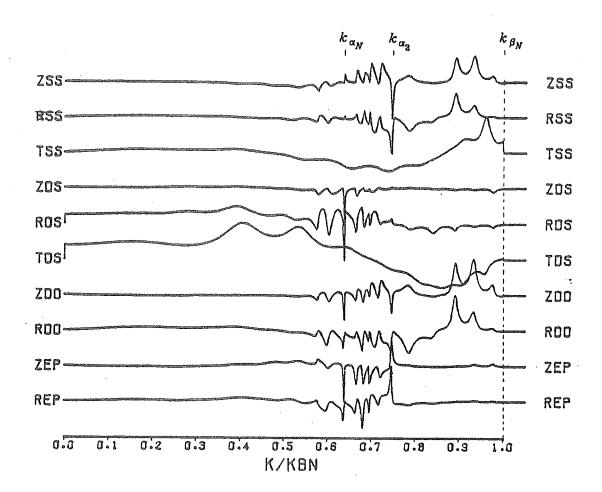


Figure 3. The responses of integrands in equation (II-2-15) along the real k-axis branch cut. The central United States (CUS) earth model, a source depth of 10 km and a frequency of 1.0 Hz, are used.

modelUnited States (CUS) (Nuttli et al, 1969; Herrmann, 1978a) of Table 1. The symbols indicate that the functions plotted arise from the corresponding integrands of far-field terms in equation (II-2-15). It can be noticed that the peaks and troughs of the excursions, corresponding to P-SV and SH respectively, are aligned, although the magnitude might vary. caused by some factors independent of the source type. Another feature seen in Figure 3 is that the tangential TSS and TDS, are substantially smoother components, than any of the P-SV function. The P-SV terms (Z and R components) oscillate rapidly near both $k_{\alpha N}$ and $k_{\beta N}$ branch points, indicating that they should be with care. Note also the behavior of the RDS and TDS plots as k goes to zero. These non-zero limits are the reason for the involved contour integration required in equations (II-4-3) and (II-4-4).

Figure 4 shows the variation of integrands for the ZSS component with frequencies between 0.10 and 10.0 Hz. The point source is at a depth of 10 km in the simple crust model (SCM) of Table 1. The requirements for adequate sampling of the integrand for proper numerical integration at high frequencies are obvious. Given the frequency f, we found it adequate to sample the region between $(0, k_{\beta_N})$ by 50+200*f points. This sampling represents one of the most time consuming aspects of our computations. Such calculations set practical

TABLE 1
Earth Models

			·
Thickness (km)	P vel (km/sec)	S vel (km/sec)	Density (g/cm³)
	Simple Crustal Model		
40	6.15	3·55	2.8
· elapse	8.09	4.67	3.3
	Central U.	Central U. S. Model	
1	5.00	2.89	2.5
9	6.10	3.52	2.7
10	6.40	3.70	2.9
20	6.70	3.87	3.0
Cozzo	8.15	4.70	3.4

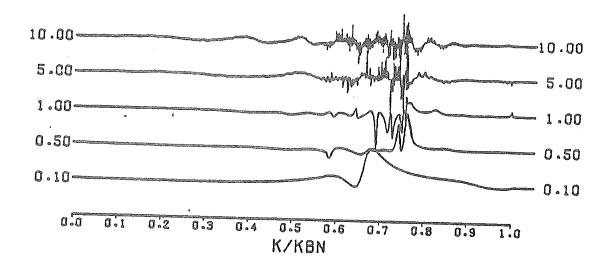


Figure 4. The real k-axis wavenumber responses of ZSS component along the real branch cut as a function of frequency between 0.1 and 10.0 Hz. The simple crust model (SCM) and a 10 km deep source are used.

limits to this form of numerical integration.

To alleviate any numerical problems near the P and S branch points, a change of variable of integration is introduced following Fuchs and Müller (1971):

$$k = \begin{cases} k_{\alpha} \sin \gamma & \gamma = [0, \frac{\pi}{2}] & \text{for } 0 < k \le k_{\alpha} \\ \frac{k_{\alpha} + k_{\beta_{N}}}{2} + \frac{k_{\alpha} - k_{\beta_{N}}}{2} \cos \gamma & \gamma = [0, \pi] & \text{for } k_{\alpha} < k \le k_{\beta_{N}} \end{cases}$$
(II-4-5)

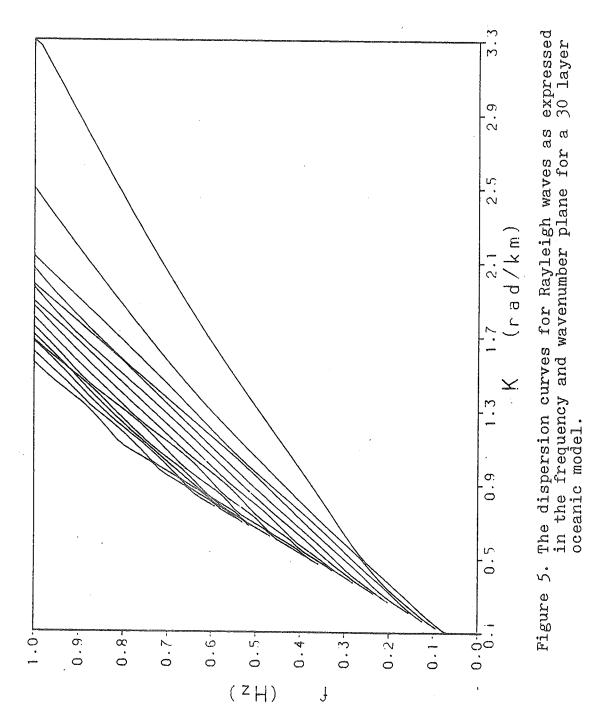
where k_{α} is chosen as an average of $k_{\alpha_n}*d_n$ over all upper layers. This formulation permits the solution of the halfspace problem as a subset of the layered halfspace problem by removing the possible singularities due to the branch points at $k_{\alpha N}$ and $k_{\beta N}$. For a layered medium, the transformation weights out the possibility of a surface wave pole at the $k_{\beta N}$ branch point interacting with the real branch line integral. The success of this particular transformation is evident from the quality of the synthetic seismograms presented later. For SH integrals, the transformation $k = k_{\beta N} \sin \gamma$ for $\gamma = (0, \pi/2)$ is used.

The residues due to the poles of the integrand occur at the zeroes of the period equation. The number of poles, or surface wave modes, increases with frequency. For example, the SCM model of Table 1 has one Rayleigh wave pole at the frequency of 0.01 Hz, and twenty poles at the frequency of 2 Hz. The number of

poles increases almost linearly with frequency. For the dispersion studies, the positions of poles need only be grossly located compared to the precision requirements for seismogram synthesis.

Figures 5 and 6 display the dispersion relations the frequency-wavenumber plane and phase velocityperiod plane, respectively, for a complex oceanic model. stair-like dispersion pattern is obvious, which usually causes difficulty for pole searching. overcome this, a pole searching technique was developed. Since the dispersion curves vary more in the frequency-wavenumber domain than in the phase velocity-period domain, the pole location will be traced in the $\omega - k$ plane.

We start with a very fine search between $k = k_{\beta_N}$ and $k = k_{\beta_{\min}}$ (for Rayleigh wave use $k = k_{\beta_{\min}}$ 0.88 to include the fundamental mode) at the two highest frequencies of interest. At this stage a dense search technique with an interval halving refinement is used to find zero crossings. The poles at the next lower frequency could, of course, be found by repeating the same procedures. However, we can save computation time by using the fact that for most earth models, the phase velocity of a given mode always increases with decreasing frequency. For a given mode we known that



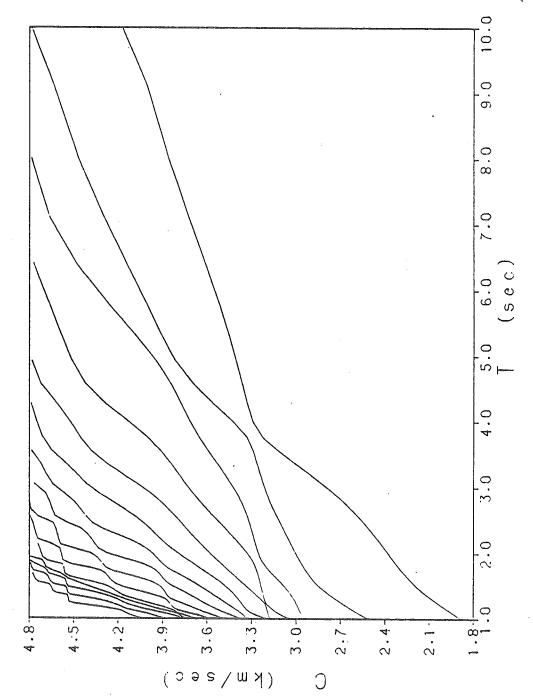


Figure 6. The dispersion curves for Rayleigh waves as expressed in the phase velocity and period plane. The same earth model as Figure 5 is used.

$$\omega = k c$$
,

and that

$$\Delta k = \frac{\Delta \omega}{c} - k \frac{\Delta c}{c} . \tag{II-4-6}$$

This can be used to estimate the pole position of a mode from its position at slightly higher frequencies. The previously calculated pole positions at two higher frequencies are used to give Δc . The search at the present frequency begins with the lowest order mode at $k_{new} = k_{old} + \frac{\Delta \omega}{c}$. This guarantees that a particular mode is not overshot.

Next, several fractions of the increment $k\frac{\Delta c}{c}$ are added to k_{nsw} to find the place where the sign of the period equation changes, at which time the interval-halving method is used to refine the zero. The computational efficiency as well as the results are found to be substantially improved, since the modes are now followed rather than searched at each frequency.

The numerical evaluation of the imaginary axis branch line integrals is the same as that given by Herrmann (1979). The Gauss-Laguerre integration rule was used since P-SV, SH functions vary harmonically and since the $K_n(\tau r)$ decayed exponentially;

$$\int_{0}^{\infty} f(i\tau,\omega) K_{n}(\tau r) d\tau = \sum_{k=1}^{m} w_{k} [f(i\frac{x_{k}}{\tau})] [x_{k} e^{x_{k}} K_{n}(x_{k})] \frac{1}{x_{k} \tau}$$

$$\int_{0}^{\infty} f(i\tau,\omega) K_{n}(\tau\tau) \tau d\tau = \sum_{k=1}^{m} w_{k} [f(i\frac{x_{k}}{\tau})] [x_{k} e^{x_{k}} K_{n}(x_{k})] \frac{1}{\tau^{2}},$$
 (II-4-7)

where w_k is the weight and $x=\tau r$. Because of the r term in equation (II-4-7), the imaginary branch line integral only becomes important at short distances. For computation, the integrand function equation (II-2-7) was used rather than compound matrix formulation (II-2-8) because there are no exponential terms involved. An m=100 order Gauss-Laguerre integration rule is used, which should be valid at radial distances as close as one-half source depth.

Synthetic Seismograms

After the values of $u_z(r,\vartheta,0,\omega)$, $u_r(r,\vartheta,0,\omega)$, and $u_\vartheta(r,\vartheta,0,\omega)$ are calculated at several discrete frequencies, we take the inverse Fourier transform with respect to frequency to form the time histories. The source time function with spectrum $s(\omega)$ should be taken into account at the same time:

$$u(r,\vartheta,0,t) = \int_{-\infty}^{\infty} s(\omega) u(r,\vartheta,0,\omega) \exp(i\omega t) d\omega/2\pi.$$

Since the fast Fourier transform (FFT) is used to approximate the Fourier integral, the source time function given by Herrmann (1979)

$$2\tau s(t) = \begin{cases} 0 & t \le 0 \\ \frac{1}{2} (\frac{t}{\tau})^2 & 0 < t \le \tau \\ -\frac{1}{2} (\frac{t}{\tau})^2 + 2(\frac{t}{\tau}) - 1 & \tau < t \le 3\tau \\ \frac{1}{2} (\frac{t}{\tau})^2 - 4(\frac{t}{\tau}) + 8 & 3\tau < t \le 4\tau \\ 0 & 4\tau < t \end{cases}$$
(II-4-8)

is used. In the context of dislocation theory, s(t) is proportional to the velocity of the dislocation, or equivalently the far-field displacement time history in an infinite medium. The time histories generated by such an impulse will be those of ground velocity. The Fourier amplitude spectrum of this pulse is enveloped by f^0 and f^{-3} asymptotes which intersect at a corner frequency of $f_c = 1/(4.36*\tau)$ (Herrmann and Wang, 1979). To avoid numerical noise problems with the FFT, τ is taken to be an integral of sampling interval, Δt .

The above numerical technique was tested by generating theoretical seismograms using the models listed in Table 1. Figures 7, 8, and 9 illustrate some high quality seismograms from a vertical dip-slip source in the SCM model for three different components, respectively. A source depth of 10 km, seismic moment of 1.0 E +20 dyne-cm, source-time function with $\tau = 0.4$ sec, and a frequency range from 0.0 to 1.25 Hz are used. The seismograms at distances less than 100 km include both the near— and far-field terms; beyond this dis-

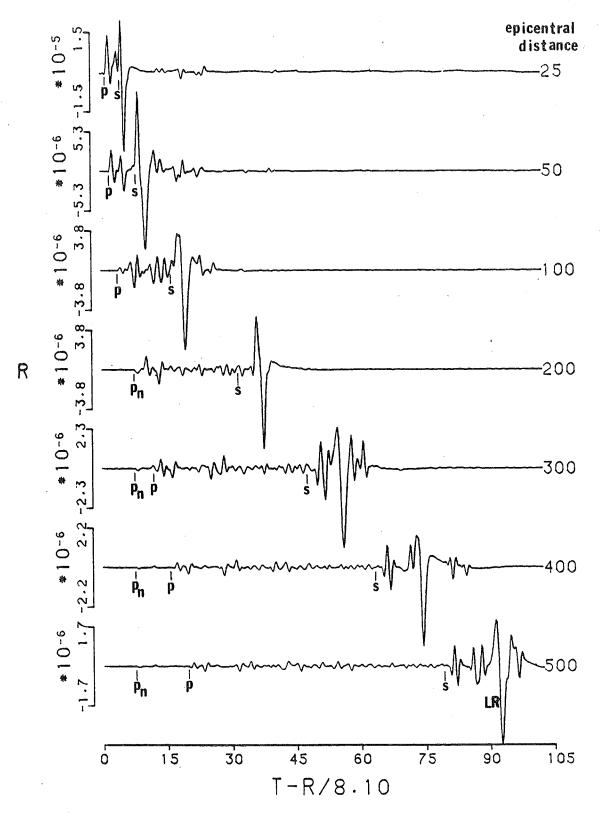


Figure 7. Radial component velocity time histories (RDS) due to a vertical dip-slip dislocation source at a depth of 10 km in SCM model. A source time function with τ =0.5 sec and seismic moment of 1.0E20 dyne-cm are used.

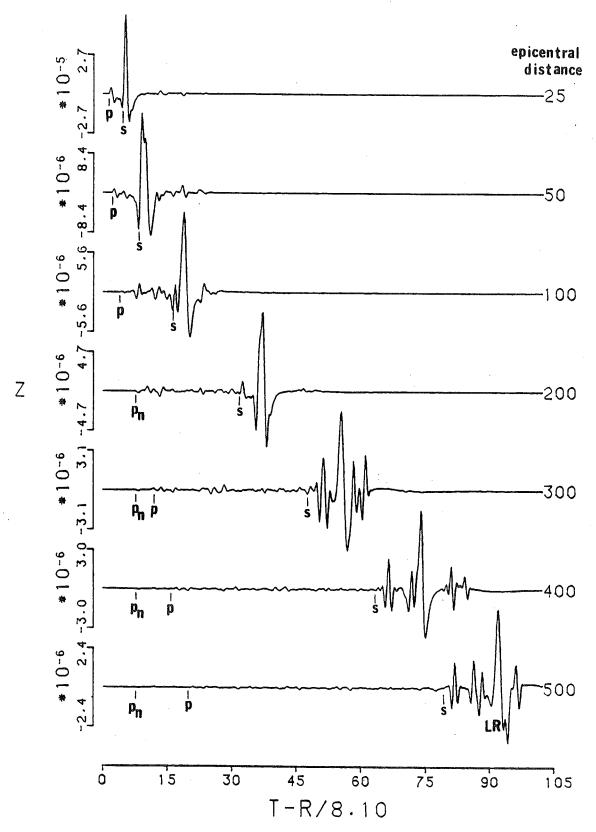


Figure 8. Results corresponding to Figure 7 but for the vertical component.

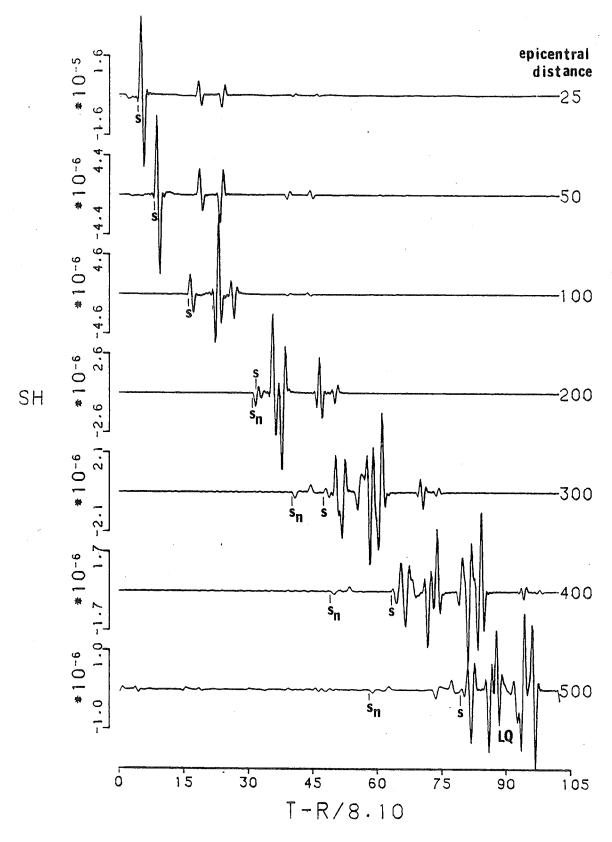


Figure 9. Results corresponding to Figure 7 but for the tangential component.

tance only the far-field terms are used. It is noted immediately that the body and surface phases are very clear. At large distances, there are multiple reflections following Pn, the first arrival, which last until the onset of the surface waves. Rayleigh-wave groups are well developed, even at short distances, in distinction to the Love-wave group which becomes apparent only at ranges larger than the S-wave critical distance (about 81.9 km). At the distance of 500 km, an interpolation of the pole contribution with one-half the frequency spacing was applied to double the time window (Kennett, 1980). At this distance, only the first half of the time series is plotted.

Figure 10 displays the radial component seimograms due to a 45° dip-slip source of 1 km depth in the CUS model of Table 1. The reverberation within the top layer provides a wave guide which generates large surface waves. The well-developed dispersed groups consist mostly of fundamental mode waves, which might have been greatly attenuated in the real earth structure. Some higher mode Lg wave arrivals prior to the dispersed wave train can easily be seen.

Figure 11 shows the effect of focal depth. The model used is the CUS model of Table 1, and the source type is vertical strike-slip. The epicentral distance is kept at 100 km. Numbers at the end of each seismo-

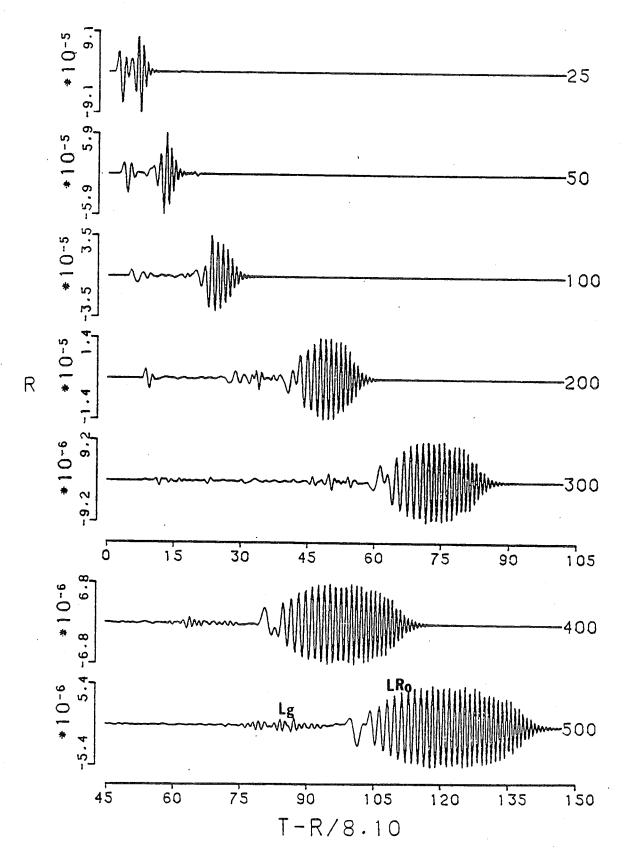


Figure 10. Radial component velocity time histories (RDD) due to a 45° dip-slip dislocation source at a depth of 10 km in CUS model. Other parameters are the same as in Figure 7.

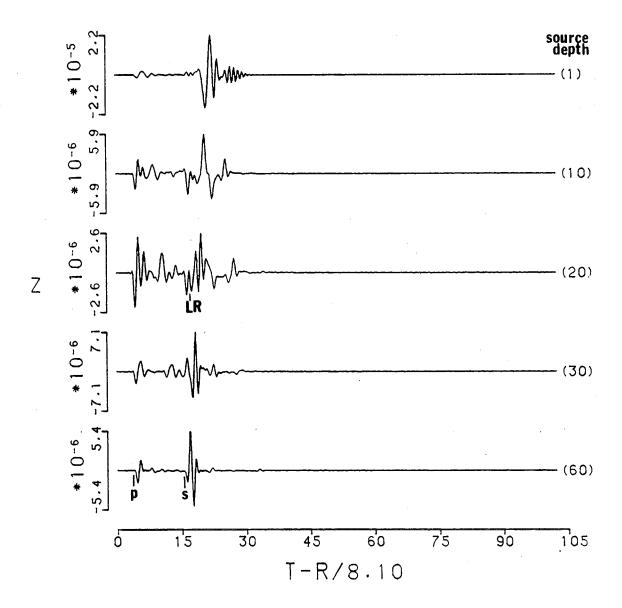


Figure 11. Vertical component velocity time histories (ZSS) due to a vertical strike-slip source in CUS model. The sources are buried at different depth as indicated at the end of each seismogram. Epicentral distance is kept at 100 km. Other parameters are the same as in Figure 7.

gram are the source depths. It seems that for shallow earthquakes the surface wave is dominant. When the source is deeper the seismogram becomes complex due to the increase of the body waves with respect to the surface waves. When the source is very deep the seismogram becomes simple again, because it consists mostly of body waves. Surface waves have not yet had time to develop. Note that the source at depth 60 km is already in the halfspace.

CHAPTER III

SURFACE WAVE - NORMAL MODE STUDY

The application of eigenfunction theory to the study of surface waves important step which is an incorporates modern mathematics into the investigation of wave propagation in the earth. Using Green's theory as a tool, the response functions in the transformed domain can be defined along mutually perpendicular directions of chosen coordinates. By the imposition boundary conditions, the values of response functions are discretized and yield specific modes of the surface wave. Because of this, the theory is also called normal mode theory. Orthogonality and normalization, two most useful procedures in eigenfunction theory, further produce the final solutions in a form which is easy to use. The reason for such a discrete excitation arises because surface waves are a type of trapped waves which reflect back and forth in the layers, and exponentially in the halfspace. A consequence of this system is the conservation of its total energy. Takeuchi and Saito (1972) used the calculus of variations to derive quantities concerning the derivatives of the conservative energy. These partial derivatives somewhat sensitive to the properties of the model and can be used to invert for an earth model from wave dispersion data.

The numerical application of eigenfunction theory discussed by Bolt and Dorman (1961) and Takeuchi and Saito (1972). They chose two independent solutions along two perpendicular directions as initial values and used the Runge-Kutta method to integrate over layers. The boundary condition at the free surface is imposed in order to regulate the variation to final solutions. Because of exponential growth of several terms, this method fails to fit the special case of very high frequencies. In this chapter, we will derive a new method to calculate the eigenfunctions. The main effort is directed toward relating the eigenfunction theory to the general layer matrix method. As a result of the symmetry properties of Haskell matrices revealed in the last chapter, the eigenfunctions as well as the integrals can be expressed in closed, analytic, energy forms. Some well-known formulae are revised and rigorously proved using the new framework. The overall objective of this chapter is the calculation of frequency signals in any plane layered model.

3.1 Haskell Matrices for the Layers

The eigenfuction problem for the generation of surface waves in plane layered media is an important topic in the theory of seismogram synthesis. After Lamb's famous study, many authors have explored this

problem using different approaches (Alterman et al, 1959; Keilis-Borok, 1960; Saito, 1967; Takeuchi and Saito, 1972; Harvey, 1981). Here we will start the solutions for displacements derived in the last chapter (equation II-2-8), and arrive at a formulation using eigenfunction values at the source depth. The reason we need the eigenfunctions rather than Haskell's matrices X or Z is because these eigenfunctions can be used to determine the energy which is conserved for surface waves traveling in the layers. Such an energy system is permitted to be perturbed to derive other properties. For example, quantities such as group velocity, dissipation functions, and amplitude factors, which arise from small perturbations of elastic parameters, can be obtained by the variation of total energy (Jeffreys, 1961; Harkrider and Anderson, 1966). In this section, a new formulation for evaluating the eigenfunction values at any depth, which was suggested by Dr. D. G. Harkrider (personal communication), will be presented. This formulation is not only computationally stable and precise, but is consistent with the system derived in the last chapter.

Now, let us discuss the eigenfunctions in the layers. The eigenfunction at any depth can be determined from the eigenfunctions at the free surface by the layer matrices in between

$$B_m = Z B_1 \tag{III-1-1}$$

where

$$\mathbf{Z} = \boldsymbol{a}_m \, \boldsymbol{a}_{m-1} \cdots \, \boldsymbol{a}_1 .$$

Ş

P-SV Eigenfunctions

As the first step, we consider the P-SV case. Expressing equation (III-1-1) in index form:

$$B_{m_k} = Z_{k1}U_{r_1} + Z_{k2}U_{z_1} , \qquad (III-1-2)$$

where we have used the surface boundary conditions which require the stresses to vanish at the free surface. For ease of expression, we will use the indices m, n, or N for the layers, and the indices i, j, k, or l for the components in the corresponding vector or matrix which have values l to 4 for P-SV waves and 5, 6 for SH waves. Normalizing by the surface z-component displacement U_{z_1} , equation (III-1-2) becomes

$$\overline{B}_{m_k} = B_{m_k} / U_{z_1} = Z_{k1} (U_{r_1} / U_{z_1}) + Z_{k2}$$
 (III-1-3)

 U_{r_1}/U_{z_1} is the ellipticity at the surface, which we already proved to be independent of the source and to possess several equivalent forms:

$$\varepsilon = \frac{U_{r_1}}{U_{z_1}} = -\frac{R_{12}}{R_{11}} = -\frac{R_{22}}{R_{21}}$$

$$= -\frac{R_{23}^{12}}{R_{13}^{12}} = -\frac{R_{24}^{12}}{R_{14}^{12}} = \frac{R^{-1}_{34}^{13}}{R^{-1}_{34}^{23}} = \frac{R^{-1}_{34}^{14}}{R^{-1}_{34}^{24}}.$$
(III-1-4)

Using the last of equation (III-1-4), $R^{-1}|_{34}^{34}=0$ (period equation), and R=XZ, equation (III-1-3) becomes

$$\overline{B}_{m_{k}} = Z_{k1} \left(R^{-1} \begin{vmatrix} \frac{14}{34} / R^{-1} \begin{vmatrix} \frac{24}{34} \end{vmatrix} + Z_{k2} \right)$$

$$= \left[Z_{k1} R^{-1} \begin{vmatrix} \frac{14}{34} + Z_{k2} R^{-1} \begin{vmatrix} \frac{24}{34} + Z_{k3} R^{-1} \begin{vmatrix} \frac{34}{34} + Z_{k4} R^{-1} \begin{vmatrix} \frac{44}{34} \end{vmatrix} / R^{-1} \begin{vmatrix} \frac{24}{34} \end{vmatrix} \right]$$

$$= \left[Z_{kp} R^{-1} \begin{vmatrix} \frac{p}{34} \end{vmatrix} / R^{-1} \begin{vmatrix} \frac{24}{34} \end{vmatrix} \right]$$

$$= \left[Z_{kp} (R_{p3}^{-1} R_{44}^{-1} - R_{p4}^{-1} R_{43}^{-1}) \right] / R^{-1} \begin{vmatrix} \frac{24}{34} \end{vmatrix}$$

$$= \left[(Z_{kp} Z_{pr}^{-1}) X_{r3}^{-1} Z_{4l}^{-1} X_{l4}^{-1} - (Z_{kp} Z_{ps}^{-1}) X_{s4}^{-1} Z_{4l}^{-1} X_{l3}^{-1} \right] / R^{-1} \begin{vmatrix} \frac{24}{34} \\ \frac{24}{8} - 1 \end{vmatrix} \frac{24}{84}$$

$$= \frac{Z_{4l}^{-1} X^{-1} \begin{vmatrix} \frac{kl}{34} \\ \frac{24}{8} - 1 \end{vmatrix} \frac{24}{84}}{R^{-1} \frac{24}{34}}$$
(III-1-5)

Equation (III-1-5) includes the inverses of Haskell's matrices, which can be related to normal matrices by using the symmetry properties of compound described in Appendix C:

$$\begin{split} Z_{jl}^{-1} &= (-1)^{j+l} \ Z_{5-l,5-j} \\ X^{-1} \begin{vmatrix} kl \\ 34 \end{vmatrix} &= q_N (-1)^{k+l+1} \ X \begin{vmatrix} 12 \\ 5-l,5-k \end{vmatrix} \\ R^{-1} \begin{vmatrix} 24 \\ 34 \end{vmatrix} &= q_N (-1) \ R \begin{vmatrix} 12 \\ 13 \end{vmatrix} \end{split} .$$

Using these, equation (III-1-5) can be another form:

$$\overline{B}_{m_k} = \frac{(-1)^{k+1} X \mid_{5-l, 6-k}^{12} Z_{5-l, 1}}{-R \mid_{13}^{12}}$$

by substituting 5-k=i 5-l=j,

$$\overline{B}_{m_{5-i}} = \frac{(-1)^i X \mid_{ij}^{12} Z_{j1}}{R \mid_{i3}^{12}}$$
 (III-1-6)

which is equivalent to

7 = and 1 ... or

$$X \mid_{ij}^{12} Z_{j1} = (-1)^i R \mid_{13}^{12} \overline{B}_{m_{5-i}}$$
 (III-1-7)

Equation (III-1-6) will be the form chosen for applica-

tion.

From equation (III-1-6), we find that eigenfunctions at any depth can be calculated by a formula similar to equation (II-2-8), which has already been shown to be computationally stable and accurate. This point is worthwhile emphasizing. Since all of the matrices, their compound forms, and all of the boundary conditions are involved, this formula has attached to it many numerical and theoretical advantages. For example, the squared exponentially growing terms are suppressed by compound matrices. The models used are not restricted by the number of layers or thicknesses. The frequencies are permitted to go very high, say 200 Hz, as long as care is taken in evaluating the compound matrices. At the same time the efficiency and accuracy of calculation are all improved.

If the eigenvalues have already been found from another calculation, equation (III-1-6) can be used in the following way. First, we insert the source depths, which might be several different values, into the velocity structure as interfaces. To find the eigenfunctions at these interfaces, the calculation is started from the bottom of the layer stack upward. At every interface, $X|_{ij}^{12}$ is calculated and stored. When reaching the surface, X is transformed into R, and $R|_{13}^{12}$ is obtained. Next, a downward procedure is taken to find

 Z_{j1} . After combining with the stored $X|_{ij}^{12}$ at that depth, we obtain the eigenfunction B by equation (III-1-6). It is obvious that in these procedures the layer matrix α , its compound forms $\alpha|_{kl}^{ij}$, and the compound matrix $E_N^{-1}|_{ij}^{12}$ are involved, hence the eigenfunction problem is said to be solved in an analytic way. Such an analytic form of calculation is useful especially for high frequencies and complex structures. One test has been made with Harkrider's oceanic model (Harkrider, 1970). The result is satisfactory even with frequency as high as 200 Hz.

If we start with different forms for ellipticity in equation (III-1-4), different formulas can be obtained following the same derivation procedures. They are

$$\varepsilon = -\frac{R_{12}}{R_{11}} \rightarrow \overline{B}_{m_i} = \frac{X_{1j} Z | j_2^i}{R_{11}}$$

$$\varepsilon = -\frac{R_{22}}{R_{21}} \rightarrow \overline{B}_{m_i} = \frac{X_{2j} Z | j_2^i}{R_{21}}$$

$$\varepsilon = -\frac{R | j_2^2}{R | j_2^2} \rightarrow \overline{B}_{m_{5-i}} = \frac{(-1)^i Z_{ji}^{-1} R | j_1^{12}}{R | j_2^2}$$

$$\varepsilon = -\frac{R | j_2^2}{R | j_2^2} \rightarrow \overline{B}_{m_{5-i}} = \frac{(-1)^i Z_{ji}^{-1} R | j_2^2}{R | j_2^2}$$

$$\varepsilon = \frac{R^{-1} | j_2^3}{R^{-1} | j_2^3} \rightarrow \overline{B}_{m_{5-i}} = \frac{(-1)^i X | j_2^2 Z_{j2}}{R | j_2^2}$$

$$\varepsilon = \frac{R^{-1} | j_2^3}{R^{-1} | j_2^3} \rightarrow \overline{B}_{m_{5-i}} = \frac{(-1)^i X | j_2^2 Z_{j2}}{R | j_2^2}$$

Except for the third one, these formulas are not very useful. The reason for this statement becomes clear when a surface water layer is introduced in the next

section.

Given the eigenfunctions at any depth, we can evaluate the energy integrals which are further used to find the quantities resulting from the perturbation of elastic properties. This technique is called the variational principle, which was introduced by Jeffreys (1961) and has been widely applied now (Takeuchi and Saito, 1972; Aki and Richards, 1980, chapter 7). The energy integrals needed to form the the system Lagrangian L_R for the Rayleigh wave,

$$L_R = \omega^2 I_0 - k^2 I_1 - 2k I_2 - I_3$$
 (III-1-9)

are

$$I_{0} = \int_{0}^{\infty} \rho \left[(\overline{U}_{r}^{2}) + \overline{(U}_{z})^{2} \right] dz$$

$$I_{1} = \int_{0}^{\infty} \left[(\lambda + 2\mu)(\overline{U}_{r})^{2} + \mu(\overline{U}_{z})^{2} \right] dz$$

$$I_{2} = \int_{0}^{\infty} \left[-\mu \overline{U}_{z} \frac{d\overline{U}_{r}}{dz} + \lambda \overline{U}_{r} \frac{d\overline{U}_{z}}{dz} \right] dz$$

$$I_{3} = \int_{0}^{\infty} \left[(\lambda + 2\mu)(\frac{d\overline{U}_{z}}{dz})^{2} + \mu(\frac{d\overline{U}_{r}}{dz})^{2} \right] dz .$$
(III-1-10)

The Lagrangian is defined as the difference between the kinetic and potential energies. Its value is required to be zero if the system is at the eigen state. Hence, this function can be used to diagnose the numerical results. It should be noted that the eigenfunctions used in the integrals (equation III-1-10) are all normalized by the surface z-component displacement. The

integrals in equation (III-1-10) can be expressed in terms of the integral I_{ij} ,

$$I_{ij} = \int \overline{B}_i \overline{B}_j \ dz$$

where.

$$\overline{\mathbf{B}} = [U_r, U_z, T_z, T_r]^T / U_{z_1}.$$

The eigenfunction or motion-stress vector \mathbf{B}_m in the m'th layer can be expressed as (equation II-1-14)

$$B_m(z) = E_m \Lambda(z) K_m$$

Taking the explicit forms of matrices Λ_m and K_m , we have

$$\mathbf{B}_{m}(z) = \mathbf{E}_{m} \cdot [A^{"}e^{\nu_{\alpha}z}, B^{"}e^{\nu_{\beta}z}, A^{'}e^{-\nu_{\alpha}z}, B^{'}e^{-\nu_{\beta}z}]^{T}.$$

Since the dependence on depth z arises solely from the diagonal exponential matrix Λ , the integration of B_iB_j over a given layer m with thickness d_m is easily found to be

$$\int_{z_{m}}^{z_{m+1}} B_{i} B_{j} dz = \left[E_{i1} E_{j1} \left(A^{"} e^{\nu_{\alpha} d} \right)^{2} + E_{i3} E_{j3} \left(A^{'} \right)^{2} \right] \frac{1}{-2\nu_{\alpha}} \left(e^{-2\nu_{\alpha} d} - 1 \right)$$

$$+ \left[E_{i2} E_{j2} \left(B^{"} e^{\nu_{\alpha} d} \right)^{2} + E_{i4} E_{j4} \left(B^{'} \right)^{2} \right] \frac{1}{-2\nu_{\beta}} \left(e^{-2\nu_{\beta} d} - 1 \right)$$

$$+ \left[\left(E_{i1} E_{j2} + E_{i2} E_{j1} \right) \left(A^{"} e^{\nu_{\alpha} d} \right) \left(B^{"} e^{\nu_{\alpha} d} \right)$$

$$+ \left(E_{i3} E_{j4} + E_{i4} E_{j3} \right) \left(A^{'} \right) \left(B^{'} \right) \right] \frac{1}{-(\nu_{\alpha} + \nu_{\beta})} \left(e^{-(\nu_{\alpha} + \nu_{\beta}) d} - 1 \right)$$

$$+ \left[\left(E_{i1} E_{j4} + E_{i4} E_{j1} \right) \left(A^{"} e^{\nu_{\alpha} d} \right) \left(B^{"} \right) \right]$$

$$\left(\text{III-1-11} \right)$$

$$+ (E_{i2}E_{j3} + E_{i3}E_{j2}) (B''e^{\nu_{\beta}d}) (A')] \frac{1}{-(\nu_{\alpha} - \nu_{\beta})} (e^{-\nu_{\alpha}d} - e^{-\nu_{\beta}d})$$

$$+ [(E_{i1}E_{j3} + E_{i3}E_{j1}) (A''e^{\nu_{\alpha}d}) (A')] d e^{-\nu_{\alpha}d}$$

$$+ [(E_{i2}E_{j4} + E_{i4}E_{j2}) (B''e^{\nu_{\beta}d}) (B')] d e^{-\nu_{\beta}d} ,$$

where the layer index m has been omitted. In equation (III-1-11), A' and B' are part of elements of the vector K_m (or $\Lambda_m(0)$ K_m) which possess the values of potential constants at the top of layer, and $A''e^{\nu_a d_m}$ and $B''e^{\nu_\beta d_m}$ are part of components of the vector $\Lambda_m(d_m)$ K_m which is at the bottom of layer. These quantities can be determined from equation (II-1-15) and equation (II-1-21):

$$\mathbf{K}_{m} = [A'', B'', A', B']_{m}^{T} = \mathbf{E}_{m}^{-1} \mathbf{B}_{m}$$

$$\Lambda_{m} \mathbf{K}_{m} = [A''e^{\nu_{\alpha}d}, B''e^{\nu_{\beta}d}, A'e^{-\nu_{\alpha}d}, B'e^{-\nu_{\beta}d}]_{m}^{T} = \mathbf{E}_{m}^{-1} \mathbf{B}_{m+1},$$

i.e., A' and B' are determined from eigenfunctions at the top of the layer and $A''e^{\nu_\alpha d_m}$ and $B''e^{\nu_\alpha d_m}$ from eigenfunctions at the bottom. If it were not done this way, the A', A'', B', and B'' could be very inaccurate by just using $\mathbf{E}_m^{-1}\mathbf{B}_m$. Suppose that the eigenfunctions have been determined precisely, the energy integrals as expressed in the analytic forms as equation (III-1-11) can be calculated with sufficient accuracy.

Equation (III-1-11) is applied to the layers lying above the half-space. In the half-space, the integral

has the simpler form

$$\int_{z_N}^{\infty} B_i B_j dz = E_{i3} E_{j3} (A_N^{'})^2 \frac{1}{2\nu_{\alpha_N}}$$

$$+ E_{i4} E_{j4} (B_N^{'})^2 \frac{1}{2\nu_{\beta_N}}$$

$$+ (E_{i3} E_{j4} + E_{i4} E_{j3}) (A_N^{'}) (B_N^{'}) \frac{1}{\nu_{\alpha_N} + \nu_{\beta_N}}$$

where A_N and B_N are the elements of vector K_N , and can be determined from $K_N = E_N^{-1} \; B_N$

Some of the integrals in equation (III-1-10) require the derivatives of U_r and U_z with respect to the depth. From the differential equation (II-1-11), these derivatives can be expressed in terms of the eigenfunctions U_r , U_z , T_z , and T_r as

$$\frac{dU_r}{dz} = kU_z - \frac{\omega^2}{\mu} T_r$$

$$\frac{dU_z}{dz} = \frac{1}{\lambda + 2\mu} (-k \lambda U_r + \omega^2 T_z) .$$

Therefore, the integrals in equation (III-1-10) as expressed in the form of I_{ij} 's are

$$I_{0} = \sum_{n=1}^{N} \rho_{n} (I_{11} + I_{22})_{n}$$

$$I_{1} = \sum_{n=1}^{N} (\lambda + 2\mu)_{n} (I_{11})_{n} + \mu_{n} (I_{22})_{n}$$

$$I_{2} = \sum_{n=1}^{N} \omega^{2} [\sigma_{n} (I_{13})_{n} + (I_{24})_{n}] - k[(\lambda\sigma)_{n} (I_{11})_{n}] + \mu_{n} (I_{22})_{n}]$$

$$I_{3} = \sum_{n=1}^{N} k^{2} [(\lambda\sigma)_{n} (I_{11})_{n} + \mu_{n} (I_{22})_{n}] - 2k \omega^{2} [\sigma_{n} (I_{13})_{n} + (I_{24})_{n}]$$

$$+ \omega^{4} [(\sigma/\lambda)_{n} (I_{33})_{n} + (1/\mu)_{n} (I_{44})_{n}]$$
(III-1-12)

where

$$\sigma = \frac{\lambda}{\lambda + 2\mu} \ .$$

It is interesting to note that the Lagrangian L_R defined in equation (III-1-9) has the following form as expressed in I_{ij} 's:

$$L_{R} = \sum_{n=1}^{N} \left[(\rho - \frac{k^{2}}{\omega^{2}} \xi)_{n} (I_{11})_{n} + \rho_{n} (I_{22})_{n} - (\frac{\omega^{2}}{\lambda + 2\mu})_{n} (I_{33})_{n} - (\frac{\omega^{2}}{\mu})_{n} (I_{44})_{n} \right] \omega^{2}.$$
 (III-1-13)

The coefficients before the I_{ij} 's in this equation are just those of skew-diagonal elements in the differential equation (II-1-11). This property will be used in Appendix D for finding the amplitude factors from energy integrals.

SH Eigenfunctions

The above derivation was for the P-SV, or Rayleigh waves. We have found the analytic solutions of the eigenfunction at any depth, and the analytic forms for taking energy integrals. Similarly, for the Love waves we have the following properties:

- (1) period equation $R_{65} = 0$.
- (2) symmetry of Haskell matrices: $Z_{ij} = (-1)^{i+j} Z_{11-j,11-i}$ and $X_{ij}^{-1} = (-1)^{i+j} q_N X_{11-j,11-i}$ with $q_N = -2k^2 \mu_N \gamma_{\beta_N} / \rho_N^2$, and i, j=5,6.
- (3) since $R_{5i} \cdot R_{i5}^{-1} = 1$, therefore $R_{56} \cdot R_{65}^{-1} = 1$

Now the eigenfunctions normalized with the surface tangential displacement become

$$\overline{B}_{m_k} = Z_{k \, 5}
= X_{kp}^{-1} R_{p \, 5}
= X_{k \, 6}^{-1} R_{66}
= X_{k \, 6}^{-1} / R_{66}^{-1}
= \frac{(-1)^{k+1} X_{5,11-k}}{R_{56}} \qquad k=5.6 ,$$
(III-1-14)

or equivalently

$$X_{5i} = (-1)^i R_{56} \overline{B}_{m_{11-i}}$$
 $i=5.6$ (III-1-15)

The Lagrangian for Love waves is

$$L_L = \omega^2 I_0 - k^2 I_1 - I_2 \tag{III-1-16}$$

with

$$I_0 = \int_0^\infty \rho \ \overline{U}_{\vartheta}^2 \ dz$$

$$I_1 = \int_0^\infty \mu \ \overline{U}_{\vartheta}^2 \ dz$$

$$I_2 = \int_0^\infty \mu \ (\frac{d\overline{U}_{\vartheta}}{dz})^2 \ dz$$

The integration over one of the upper layers is

$$\int_{z_{m}}^{z_{m+1}} B_{i} B_{j} dz = \left[E_{i5} E_{j5} \left(C'' e^{\nu \beta^{d}} \right)^{2} + E_{i6} E_{j6} \left(C' \right)^{2} \right] \frac{1}{-2\nu_{\beta}} \left(e^{-2\nu_{\beta}^{d}} - 1 \right) + \left[\left(E_{i5} E_{j6} + E_{i6} E_{j5} \right) \left(C' \right) \left(C'' e^{\nu_{\beta}^{d}} \right) \right] d e^{-\nu_{\beta}^{d}}$$

and over bottom half-space is

$$\int_{z_N}^{\infty} B_i B_j \, dz = E_{i6} \, E_{j6} \, (C')^2 \, \frac{1}{2\nu_{\beta}} \quad .$$

The energy integrals as expressed in terms of I_{ii} 's take the forms:

$$I_{0} = \sum_{n=1}^{N} \rho_{n} (I_{55})_{n}$$

$$I_{1} = \sum_{n=1}^{N} \mu_{n} (I_{65})_{n}$$

$$I_{2} = \sum_{n=1}^{N} \frac{1}{\mu_{n}} (I_{66})_{n}$$
(III-1-17)

It is noted that there are some extra ω 's present in equation (III-1-12) but not in equation (III-1-17). The reason arises from the somewhat arbitrary definitions for stress eigenfunctions which we chose in equation (II-1-7). The P-SV stresses used here should be multiplied by ω^2 before comparing with other formalisms such as that of Takeuchi and Saito (1972).

3.2 Normal Mode Theory

Normal mode theory is a method which uses the boundary value problem technique to deal with the waves propagating within layers. A normal mode defines a preferred frequency of vibration for the system. The surface wave, which is the most prominent phase on the seismogram, comes from the summation of the contributions from various preferred vibrations, or modes, of the system. There have been a number of investigations of this theory (Haskell, 1964; Ben-Menahem and Harkrider, 1964; Vlaar, 1966; Saito, 1967; and Levshin and

Yanson, 1971). Their results are principally equivalent (Tsai and Aki, 1970), although some expressional differences exist, especially for the source functions. In this section, a derivation will be performed using the results constructed in the last chapter, in which Saito's (1967) important extension to evaluate the residual contributions by means of variational principles will be intensively discussed.

The Fourier-Hankel transformed displacements at the free surface expressed in equation (II-2-8) are

$$\begin{split} U_{z_1}(\omega,k,n) &= S_i X |_{ij}^{12} Z_{j1} / R |_{i2}^{12} \\ U_{r_1}(\omega,k,n) &= -S_i X |_{ij}^{12} Z_{j2} / R |_{i2}^{12} \qquad i,j=1,4 \\ U_{v_1}(\omega,k,n) &= -S_j X_{5j} / R_{55} \qquad j=5,6 . \end{split}$$

From the discussion in section 2.2, the displacements at any depth for a double-couple dislocation source can be written as:

$$\begin{split} u_z \left(r \,, \vartheta , z \,, \omega \right) &= \int U_z^{(2)} J_2(kr) dk \cdot R_{ss} + \int U_z^{(1)} J_1(kr) dk \cdot R_{ds} + \int U_z^{(0)} J_0(kr) dk \cdot R_{dd} \\ u_r \left(r \,, \vartheta , z \,, \omega \right) &= \int - U_r^{(2)} J_1(kr) dk \cdot R_{ss} + \int - U_r^{(1)} J_0(kr) dk \cdot R_{ds} + \int U_r^{(0)} J_1(kr) dk \cdot R_{dd} \\ u_{\vartheta} \left(r \,, \vartheta , z \,, \omega \right) &= \int - U_{\vartheta}^{(2)} J_1(kr) dk \cdot R_{ss} + \int - U_{\vartheta}^{(1)} J_0(kr) dk \cdot R_{ds} \quad , \end{split} \quad \text{(III-2-1)}$$

where only the far-field terms are retained, and the R's describing the radiation patterns are given in equation (II-2-15). The superscripts representing the azimuthal mode number are parenthesized for clarity. Since the eigenfunctions found in the last section are all nor-

malized with respect to U_{z_1} (P-SV) or U_{ϑ_1} (SH), the integral to be evaluated in equation (III-2-1) actually is

$$\int_{0}^{\infty} B^{(n)}(z) J_{m}(kr) dk = \int_{0}^{\infty} \overline{B}(z) U_{z_{1}}^{(n)} J_{m}(kr) dk$$

where \overline{B} is the normalized eigenfunction at the depth z, and $U_{z_1}^{(n)}$ is the surface z-component displacement. Of course, for Love waves we use U_{ϑ_1} instead of U_{z_1} .

With the eigenfunctions available at different depths, we are ready to derive the surface wave fields. After laborious substitution and expansion, final results, similar to those of Saito (1967), Levshin and Yanson (1971), or Harvey (1981) will be found. Here we just show a component, RSS, for instance. With the aid of equation (III-1-7), the RSS component can be written as

$$\begin{split} \widehat{RSS} &= \int_{0}^{\infty} -\overline{U}_{r}(z) \ U_{z_{1}}^{(2)} J_{1}(kr) \ dk \\ &= \int_{0}^{\infty} -\overline{U}_{r}(z) \ \frac{S_{i}^{(2)} X |_{ij}^{12} Z_{j1}}{R |_{12}^{12}} \ J_{1}(kr) \ dk \\ &= \int_{0}^{\infty} -\overline{U}_{r}(z) \ (-1)^{i} \ \frac{R |_{12}^{12}}{R |_{12}^{12}} \ \overline{B}_{m_{5-i}} \ S_{i}^{(2)} J_{1}(kr) \ dk \end{split}$$

where the index m indicates the source layer. Setting the period equation $R|_{12}^{12}=0$, and applying the residue theorem, we can find the surface wave displacements from the pole contributions:

$$RSS = -\pi i \sum \left[-\overline{U}_r(z) \right] \frac{R \mid_{13}^{12}}{\frac{\partial}{\partial k} R \mid_{12}^{12}} (-1)^l \overline{B}_{m_{5-l}} S_l^{(2)} H_1^{(2)}(kr) ,$$

where the summation consists of each of the normal mode contribution at a given frequency. For far-field, i.e. r very large, the Hankel function $H_n^{(2)}(kr)$ has the approximation

$$H_n^{(2)} = \left[\frac{2}{\pi kr}\right]^{\frac{1}{2}} e^{-ikr + i\frac{\pi}{4}(1+2n)}.$$

By substituting the source function $4\pi\omega^2S^{(2)}_4=-2k^2\ ,$ the RSS component surface wave at large distance becomes

$$RSS = \sum \left[-\overline{U}_r(z) \right] A_R \frac{1}{2\pi} k \overline{U}_{r_m} \left[\frac{2\pi}{kr} \right]^{\frac{1}{2}} e^{-ikr - i3\pi/4}$$

where

$$A_R = \frac{R \mid \frac{12}{13}}{\frac{\partial}{\partial k} R \mid \frac{12}{12}} \frac{k}{\omega^2}.$$

The amplitude factor or amplitude response A_R , defined by Harkrider (1964), can be evaluated in terms of phase velocity c, group velocity U, and the energy integral I_0 according to the following formula:

$$A_R = \frac{1}{2 c U I_0} \tag{III-2-2}$$

(Harkrider and Anderson, 1966; Levshin and Yanson, 1971; Takeuchi and Saito, 1972). The verification of this relation was first provided by Saito (1967). He employed the variational technique and decomposed the eigenfunction into two parts, one of

which is continuous along the z-coordinate and the other is discontinuous by the interruption of the source at the source depth (Aki and Richards, 1980, p.310). However, such a splitting of the eigenfunction is not necessary if it is only desired to express the pole residue in terms of energy integrals. In Appendix D, we apply similar techniques from the variational principle, but use only the forms of the eigenfunction defined before. The derivation is relatively straightforward, but involved. An extra k and an extra k, which were presented in section 2.3, automatically appear in our formulation.

Applying this result to the partial derivative of the period equation, we will obtain the excitation of surface waves in terms of the eigenfunction at the source depth and the energy integrals:

$$RSS = \sum \left[-\overline{U}_r(z) \right] \frac{1}{2cUI_0} \frac{1}{2\pi} k \overline{U}_{r_m} \left[\frac{2\pi}{kr} \right]^{\frac{1}{2}} e^{-ikr - i3\pi/4}$$

The same procedures can be applied to all other components with the results:

$$u_{z}(r,\vartheta,z,\omega) = \sum D_{kR} \frac{\overline{U}_{z}(z)}{2cUI_{0}} \left[\frac{2\pi}{kr} \right]^{\frac{1}{2}} e^{-ikr - i\frac{\pi}{4}}$$

$$u_{r}(r,\vartheta,z,\omega) = \sum D_{kR} \frac{-\overline{U}_{r}(z)}{2cUI_{0}} \left[\frac{2\pi}{kr} \right]^{\frac{1}{2}} e^{-ikr - i\frac{3\pi}{4}}$$

$$u_{v}(r,\vartheta,z,\omega) = \sum D_{kL} \frac{\overline{U}_{v}(z)}{2cUI_{0}} \left[\frac{2\pi}{kr} \right]^{\frac{1}{2}} e^{-ikr + i\frac{\pi}{4}}$$
(III-2-3)

where

$$\begin{split} D_{kR} &= \frac{1}{2\pi} \left[(k\overline{U}_{r_m}) R_{ss} + \{ 2 \frac{k_{\alpha_m}^2}{\rho_m} \overline{T}_{z_m} - k [3 - (\frac{2\beta_m}{\alpha_m})^2] \overline{U}_{r_m} \} R_{dd} \right. \\ &+ i \left. (\frac{k_{\beta_m}^2}{\rho_m} \overline{T}_{r_m}) R_{ds} \right] \end{split}$$

$$D_{kL} = \frac{1}{2\pi} \left[k \overline{U}_{\vartheta_m} R_{ss}^{\prime} - i \frac{1}{\mu_m} \overline{T}_{\vartheta_m} R_{ds}^{\prime} \right] .$$

These solutions are the same as others (Saito, 1967; Herrmann, 1974; etc.). It is noted that at the free surface $\overline{U}_{r_1} = \varepsilon$ and $\overline{U}_{z_1}(0) = 1$. Hence the z- and r-component displacements at a particular frequency have the ratio $i\varepsilon$ (Haskell, 1953), i.e., a 90 degrees phase shift and ε maximum amplitude ratio.

Water Laver

If there is a water layer on the top of the solid layer stack, as in oceanic models, the situation does not become overly complicated. Only a simple modification of the formula derived above is needed. Denote the surface water layer by index 'Ø'. The calculation of eigenfunctions in the water layer is just a case of acoustic wave propagation which is widely used in applied physics. The equations to be solved are

$$\nabla^2 \varphi = \frac{1}{\alpha^2} \frac{\partial^2 \varphi}{\partial t^2}$$

$$u_r = \frac{\partial \varphi}{\partial r}$$

$$u_z = \frac{\partial \varphi}{\partial z}$$

$$T_z = \lambda \nabla^2 \varphi .$$

The boundary conditions become

$$\varphi = 0 \qquad at \quad z = 0$$

$$u_{z_1} = u_{z_0}$$

$$T_{z_1} = T_{z_0}$$

$$at \quad z = d_0$$

(Ewing et al, 1957, p.158). It is noticed that the u_r component is no longer continuous across the solid-liquid boundary, and $T_r=0$ in the water layer, for which only compressional potential φ is still needed. The solutions similar to those of equation (II-1-4), (II-1-6), and (II-1-7) reduce to

$$\varphi(r,\vartheta,z,\omega) = \frac{\cos n\,\vartheta}{\sin n\,\vartheta} \; J_n(kr) \bigg\} \; Z_1(z) \bigg\}$$
 and
$$Z_1(z) = A' \, e^{-\nu_\alpha z} + A'' \, e^{\nu_\alpha z} = -2A' \sinh \nu_\alpha z \; ,$$

where we have used the free surface boundary condition, i.e., $Z_1(0)=0$. The eigenfunctions are defined in the same way as before;

$$\begin{split} &4\pi\rho_0 u_r(r,\vartheta,z,\omega) = \cos\vartheta\left[-\{-kZ_1^c\}\frac{dJ_n(kr)}{dkr}\right] + \sin\vartheta\left[c \to s\right] \\ &4\pi\rho_0 u_z(r,\vartheta,z,\omega) = \cos\vartheta\left[\left\{\frac{dZ_1^c}{dz}\right\}J_n(kr)\right] + \sin\vartheta\left[c \to s\right] \end{aligned} \quad \text{(III-2-4)} \\ &4\pi T_z(r,\vartheta,z,\omega) = \cos\vartheta\left[\omega^2\{-Z_1\}J_n(kr)\right] + \sin\vartheta\left[c \to s\right] ,$$

i.e.,

$$U_r = -\frac{k}{\rho_0} Z_1 = \frac{2A'}{\rho_0} k \sinh \nu_{\alpha} z$$

$$U_z = \frac{1}{\rho_0} \frac{dZ_1}{dz} = -\frac{2A'}{\rho_0} \nu_{\alpha} \cosh \nu_{\alpha} z$$

$$T_z = -Z_1 = 2A' \sinh \nu_{\alpha} z$$

After normalization by U_{z_1} at $z = d_0$, we have

$$\begin{split} \overline{U}_r &= -k \sinh \nu_{\alpha} z / (\nu_{\alpha} \cosh \nu_{\alpha} d_0) \\ \overline{U}_z &= \cosh \nu_{\alpha} z / \cosh \nu_{\alpha} d_0 \\ \overline{T}_z &= -\rho_0 \sinh \nu_{\alpha} z / (\nu_{\alpha} \cosh \nu_{\alpha} d_0) \end{split}$$
 (III-2-5)

Since there is the same Bessel function dependence for T_z and U_z components as in equation (III-2-4), we find that

$$\overline{T}_{z_1} = \left(-\frac{\rho_0 \sinh \nu_{\alpha} d_0}{\nu_{\alpha} \cosh \nu_{\alpha} d_0}\right) \overline{U}_{z_1} = T \overline{U}_{z_1}$$

at the solid-liquid boundary. T was also derived by Harkrider (1964) from the ratio a_{32}/a_{22} for the water layer $\beta_0=0$.

Now let us check the effect of the water layer on the results derived before. The surface displacements to be found are now at the top of the solid layers. First, the period equation is altered to appear as

$$\begin{bmatrix} 0 \\ 0 \\ A_{N} \\ B_{N} \end{bmatrix} = \mathbb{R} \begin{bmatrix} U_{r_{1}} \\ U_{z_{1}} \\ T_{z_{1}} \\ 0 \end{bmatrix} = \mathbb{R} \begin{bmatrix} U_{r_{1}} \\ U_{z_{1}} \\ TU_{z_{1}} \\ 0 \end{bmatrix}$$
(III-2-6)

where R includes the matrices for solid layers only. The first two rows give the new period equation:

$$R \mid_{12}^{12} + T \mid_{13}^{12} = 0$$
.

It is noted that because the period equation changes, so do the dispersion values. The form of the ellipticity relation at the solid-liquid boundary is not affected if we use the following derivation:

$$\varepsilon = \frac{U_{r_1}}{U_{z_1}} = -\frac{R_{12} + TR_{13}}{R_{11}} = -\frac{R_{22} + TR_{23}}{R_{21}}$$

$$= -\frac{R_{12}R_{23} + TR_{13}R_{23}}{R_{11}R_{23}} = -\frac{R_{22}R_{13} + TR_{23}R_{13}}{R_{21}R_{13}}$$

$$= -\frac{R_{12}R_{23} - R_{22}R_{13} + T(R_{13}R_{23} - R_{23}R_{13})}{R_{11}R_{23} - R_{21}R_{13}} = -\frac{R_{12}R_{13}}{R_{13}}, \quad (III-2-7)$$

or equivalently,

$$\varepsilon = \frac{R^{-1}|_{34}^{14}}{R^{-1}|_{34}^{24}} \tag{III-2-8}$$

However, other forms, such as those in equation (III-1-4), can no longer be used. For example, there is yet another form:

$$\varepsilon = -\frac{R \mid_{24}^{12} + T \mid_{34}^{12}}{R \mid_{14}^{12}}$$

As a consequence, the eigenfunction formula (equations III-1-6 and III-1-8), if expressed in the corresponding forms with equations (III-2-7) and (III-2-8), will stay the same, but this is not true for the other forms. This point should be noticed when a water layer is imposed.

The energy trapped in the water layer is easy to obtain by direct integration from equation (III-2-5). Here we just list the results needed to calculate the Lagrangian:

$$\int_{0}^{d_{0}} \overline{U_{r}} \overline{U_{r}} dz = \frac{k^{2}}{2A \nu_{\alpha}^{2}} [B - d_{0}]$$

$$\int_{0}^{d_{0}} \overline{U}_{z} \overline{U}_{z} dz = \frac{1}{2A} [B + d_{0}]$$

$$\int_{0}^{d_{0}} \overline{U}_{r} \frac{d\overline{U}_{z}}{dz} dz = \frac{-k}{2A} [B - d_{0}]$$

$$\int_{0}^{d_{0}} \frac{d\overline{U}_{z}}{dz} \frac{d\overline{U}_{z}}{dz} dz = \frac{\nu_{\alpha}^{2}}{2A} [B - d_{0}]$$

where

$$A = \cosh^2 \nu_{\alpha} d_0 \qquad \qquad B = \frac{\sinh 2\nu_{\alpha} d_0}{2\nu_{\alpha}}$$

The Lagrangian integral components I_i have the following forms, obtained by setting $\mu_0=0$ in equation (III-1-10),

$$\begin{split} I_0 &= \int \, \rho_0 \, (\, \overline{U}_r^{\, 2} + \overline{U}_z^{\, 2} \,) \, dz \\ I_1 &= \int \, \lambda_0 \, \overline{U}_r^{\, 2} \, dz \\ I_2 &= \int \, \lambda_0 \, \overline{U}_r \, \, \frac{d\overline{U}_z}{dz} \, dz \\ I_3 &= \int \, \lambda_0 \, \frac{d\overline{U}_z}{dz} \, \frac{d\overline{U}_z}{dz} \, dz \end{split} \; .$$

These integrals over the water layer should be added to the corresponding integrals for solid layers derived in section 3.1. Although the displacement fields calculated are at the top of solid layers, we find part of the wave energy is trapped in the first water layer, especially for high frequency signals. When using the energy viewpoint to approach the eigenfunction problem, the contribution from the water layer cannot be simply ignored. The presence of the water layer does not affect the corresponding SH solution.

Synthetic High Frequency Seismograms

In the above discussion, we have set up a complete system to calculate eigenfunctions and the related parameters. A FORTRAN program was designed to compute the surface wave time history or its spectrum for relatively high frequencies and moderately complex models. An explanation of this program is given in Appendix E.

Figures 12, 13, and 14 show examples obtained from these programs for three different components. In each figure the upper five seismograms are obtained by using symmetric triangular source time function with base of one second, and the bottom seismogram from a source time function. The frequencies cover the range from 0 to 10 Hz, although the triangular source has the corner frequency at 0.7 Hz. The CUS model in Table 1 is used. Source parameters are dip = 50° , slip = 180° , strike = 40° , and depth = 14 km. The receivers are located to the north of source, and the epicentral distances are indicated at the end of each seismogram. Furthermore, a Q-model with $Q_{\beta} = 250$ for the top 24 depth followed by $Q_{\beta} = 2000$ for the halfspace is also assumed. In these figures an apparent low frequency fundamental mode signal arives after some high frequency Lg (Airy) phases. These synthetic seismograms seem to be of quite high quality. Even the step-source generated seismograms, which contain 10 Hz frequen-

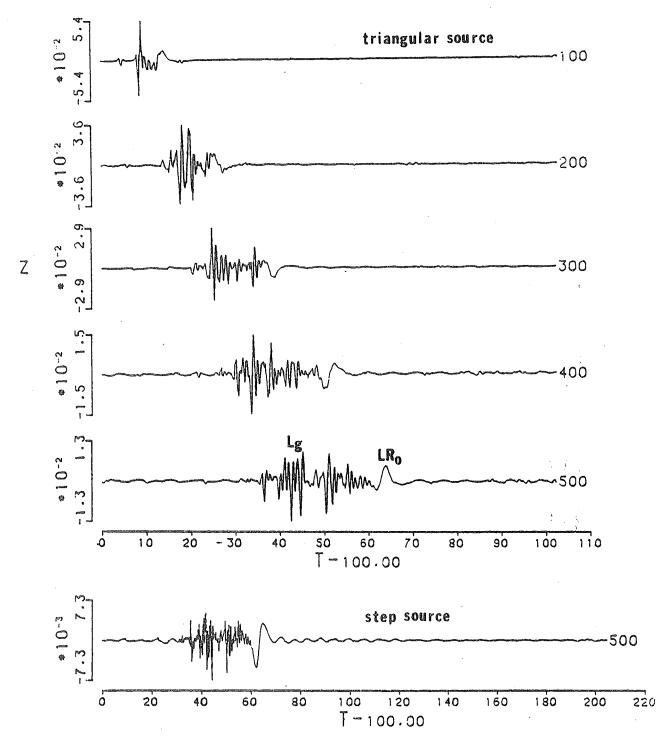


Figure 12. Theoretical seismograms generated by eigenfunction programs. The upper five seismograms are due to a dislocation source with a triangular source time function buried at the depth of 14 km in CUS model. The frequencies used cover the range from 0 to 10 Hz. The bottom seismogram is due to the same dislocation source but with a step source time function. A Q-model with $Q_{\theta} = 250$ for top 24 km and $Q_{\theta} = 2000$ for other layers is used. The number at the end of each seismogram indicates the epicentral distance.

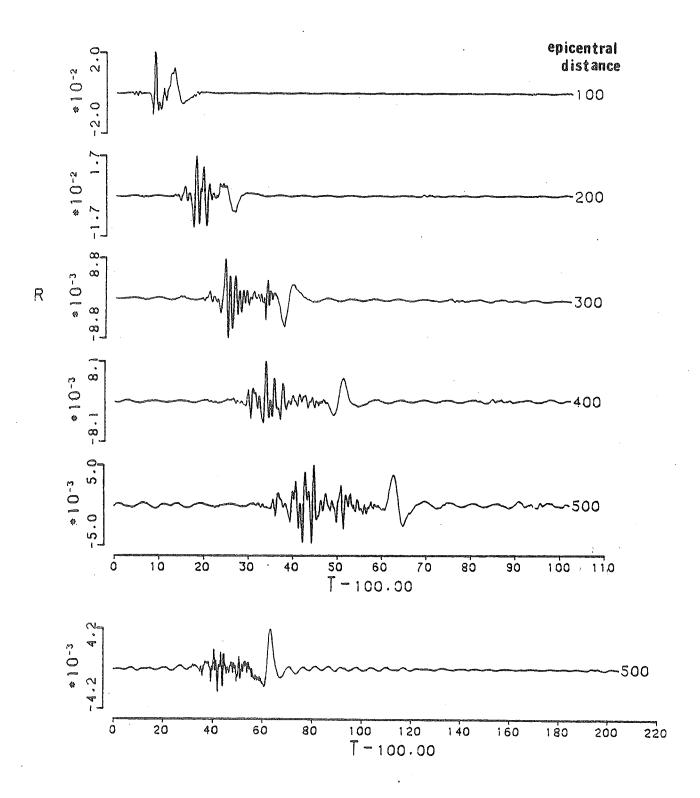


Figure 13. Results corresponding to Figure 12 but for the radial component.

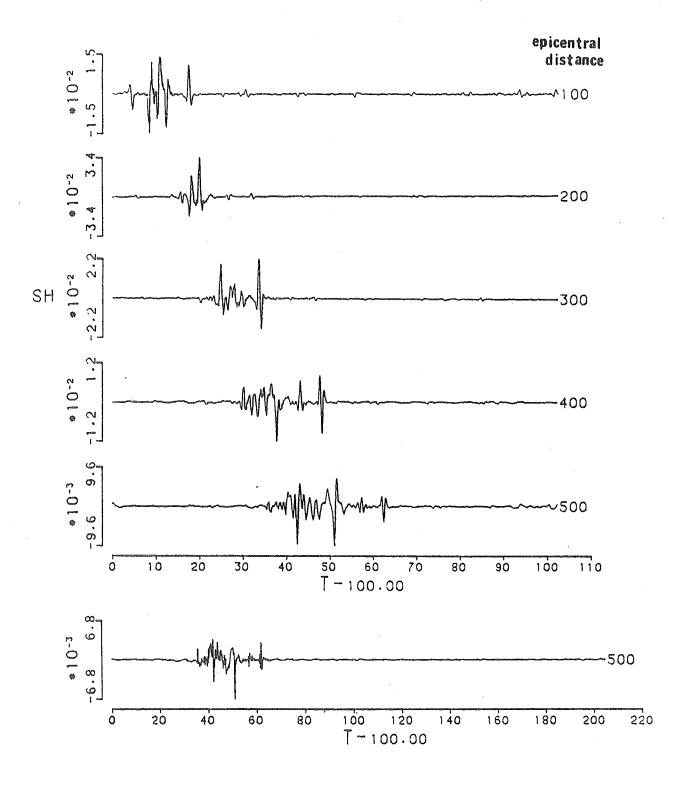


Figure 14. Results corresponding to Figure 12 but for the tangential component.

cies, exhibit well developed and fairly noiseless waveforms.

CHAPTER IV

BODY WAVE - LEAKY MODE STUDY

In the previous chapter, the normal mode theory of surface waves was considered. The objective of this chapter is to present a detailed discussion about generation of body waves. This part of the waveform comes from, at least for most P-SV cases, the branch line integral discussed in section 2.4. Figure 15a shows the contributions of the pole residue (middle), branch line integral (bottom), and the total seismogram (top) of the REP component with source depth = and r = 100 km for the SCM model of Table 1. Figure 15b gives a similar display for the RDS component with source depth = 1 km and r = 25 km for the CUS model. It is obvious that there are two definitely different signals arising from the pole contribution and the branch line integral, which constitute the This fact can further be seen from Figures seismogram. 16 and 17. However for the SH case, the situation totally different. Figure 18 indicates that, for the SH case, the pole contribution constitutes most of signal, and the branch line integral is only required the total seismogram 'causal' (Herrmann, 1978a). Thus, what is the role of the branch line integral in constructing a seismogram? This question will be discussed in this chapter using the leaky mode

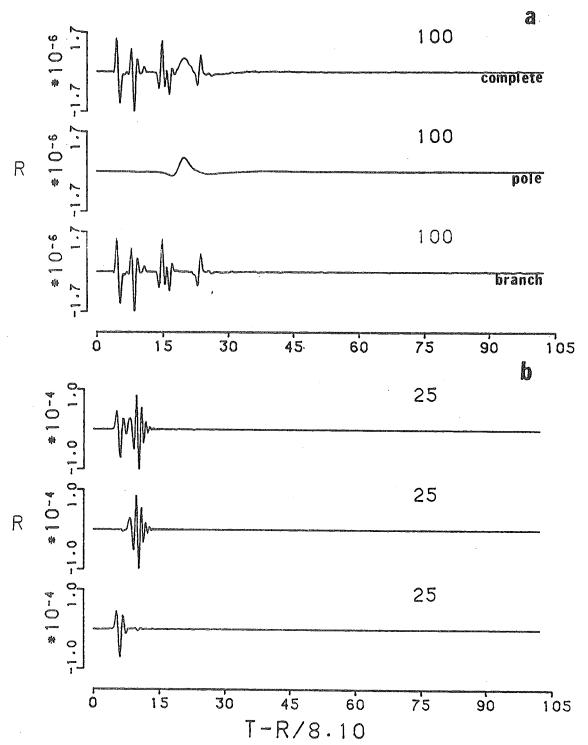


Figure 15. Study of contribution of various components of contour integration. (a) is the set of vertical component seismograms due to an explosive source at 100 km away and buried at 10 km depth in SCM model. (b) is the set of radial component seismograms due to a 45° dip-slip source at 25 km away and buried at 1 km depth in CUS model. In each set of seismograms, the top one is the complete solution, the middle is the pole contribution, and the bottom is the contribution from branch line integral.

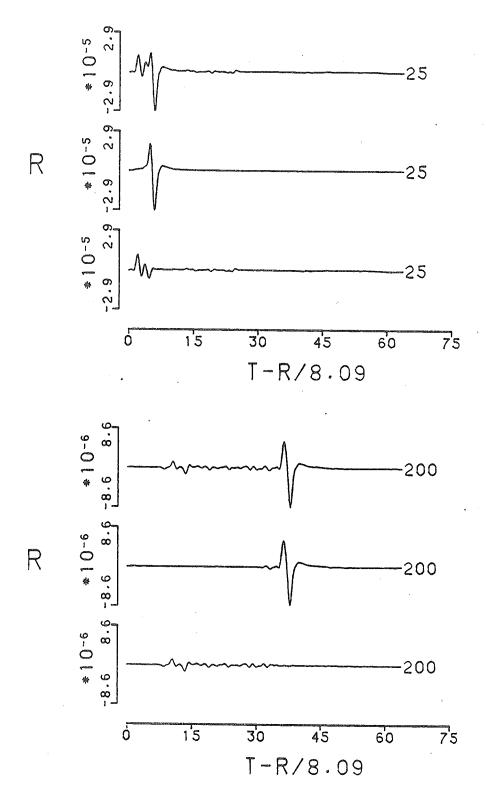


Figure 16. Same comparison as for Figure 15. The radial component seismograms due to a dip-slip source buried at 10 km depth in SCM model are displayed. Two sets of seismograms correspond to epicentral distances at 25 km and 200 km, respectively.

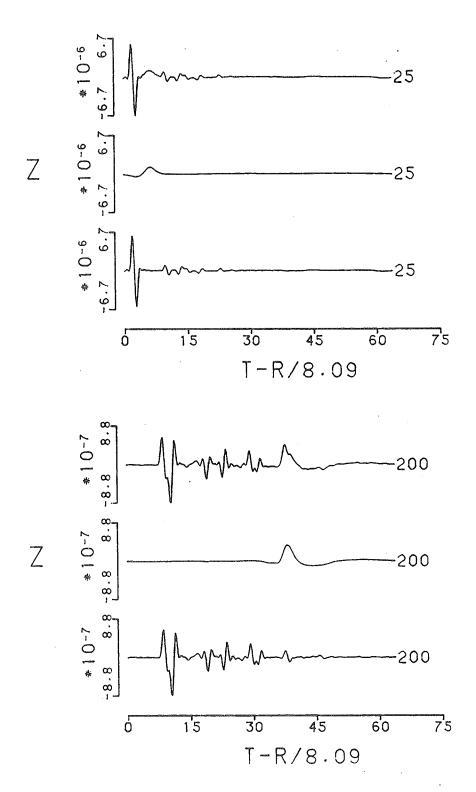


Figure 17. Same comparison as Figure 16, but for the vertical component.

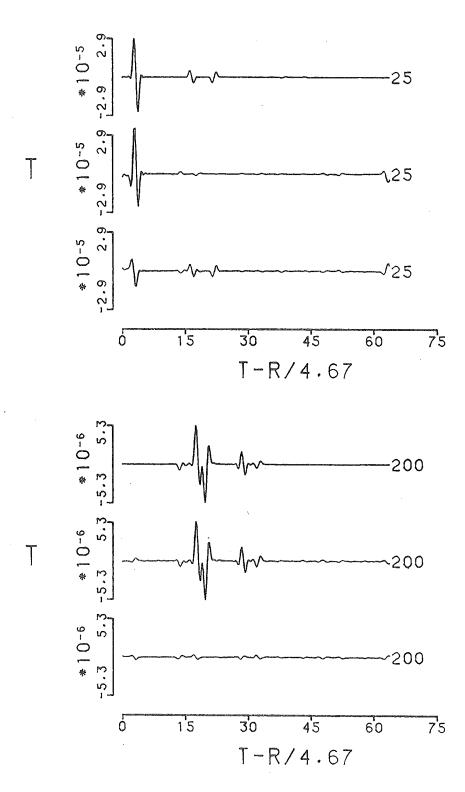


Figure 18. Same comparison as Figure 16, but for the tangential component.

approach.

Harvey (1981) imposed a deep and rigid cap at the bottom of the structure to lock the energy in the upper layers. It was found that this process is an approximation which ignores the leakage of waves and forces the leaky mode to appear on the real wavenumber axis as pseudo-normal modes. Since the normal modes are usually easy to handle, this approach provides a good way to simulate the body waves.

The reflectivity method of Fuchs and Müller (1971) is a widely used method for calculating body waves. Using a theory from Kennett (1974), we are able to generalize the reflectivity method. Most importantly, such a modification is made by simply adjusting our system constructed in chapter II. Hence, the reflectivity method is only a special case of our system.

4.1 The Influence of Leaky Modes on Body Waves

In section 2.4, we constructed the full wave field after combining two different types of solutions: the pole contribution and the branch line integral. The poles which describe the vibration modes of layered wave guides exhibit dispersion and contribute mainly to the surface wave. This point was further confirmed by the normal mode approach as presented in chapter III.

It is natural to ask whether the branch cut integral naturally gives rise to the body wave. In this section, we will extend the normal mode theory to include the leaky mode (or leaking mode), and investigate the role of branch line integration by means of the leaky mode approach.

The study of leaky modes began when the computer in its infancy. The direct evaluation of a line integral along the branch cuts is difficult. overcome this, two methods were proposed. One was the use of the steepest descent method to find the approximate solutions for large source-receiver distances (Lapwood, 1949; Fuchs, 1971). The other was an attempt deform the integration path into other Riemann sheets and evaluate the pole residue there (Rosenbaum, 1960; Phinney, 1961). The branch cut and Riemann sheets introduced in order to were make the vertical wavenumber single valued. There exist four sheets for P-SV with respect to four combinations of $\operatorname{Re}\left(\left.
u_{eta}
ight)$ being positive or negative, and two sheets for SH. We will use (+,-) for positive $Re(\nu_{\alpha})$ and negative $\operatorname{Re}(\nu_{\beta})$, respectively, and similar sign symbols for other combinations. The poles in the lower Riemann sheets are called leaky modes, as distinct from normal modes. Since these poles are situated off the real in the complex plane, imaginary parts cause the attenuation as waves propagate with distance or

Physically, the leaky modes describe traveling disturbances as a leaking S wave or P and S waves, respectively, into the substratum. At moderate distances, such a mode might be seen as it sometimes results in large amplitude dispersed oscillatory trains appearing in the interval between the P and S arrivals, which are called PL waves.

Many authors have explored leaky modes by extending the normal (or trapped) mode theory (Phinney, 1961; Gilbert, 1964; Laster et al, 1965; Haskell, Abramovici, 1968; Alsop, 1970; Cochran et al, 1970; Dainty, 1971; Watson, 1972). Among these works, that of Gilbert (1964) is of fundamental importance. Using real k-complex f as independent variables, Gilbert classified the modes into two types: Lamb's roots and organ pipe roots, when k approaches zero. Lamb's roots are associated largely with the properties of the half-space. One of these roots is the fundamental mode of the Rayleigh wave which occupies (+,+) position over the whole range of k-f. But the other Lamb's behave like organ pipe modes as k or f become large. An organ pipe mode represents the standing wave trapped in the layer. It is the mode associated with the energy reverberated almost vertically inside the layers. forms of organ pipe mode exist: π modes which possess P-wave properties and originate from the (-,+)and Σ modes which have S-wave properties and come from

the (+,-) sheet.

Following the work of Gilbert (1964), investigation principally concentrated on the search for modes, i.e., defining the dispersion curve and its attenuation Laster et al (1965) provided a theoretical relation. basis for the dispersion study. They also generated theoretical seismograms from leaky mode contributions and indicated the importance of the leaky mode early part of the seismogram by comparison to experimental data. Cochran et al (1970) displayed the dispersion curves for π and Σ modes in a multiple elastic wave guide, and found lattice dispersion patterns which are characterized by π -pseudo and Σ -pseudo modes. The π -pseudo modes were shown to make up the oscillatory part of the seismogram between P and S, which depends solely on the P velocities. Abramovici (1968) the compound matrix technique to the leaky mode calculation, and defined a transfer function to study the contribution of individual modes. Except for Laster et al (1965), the real k-complex f approach of Gilbert (1964)has been employed to compute dispersion curves by all other investigators. Watson (1972)made an interesting study using real f-complex k analysis. The most significant contribution of his work was the iden-PL and OP modes. His approach is espetification of cially suitable for our present study. It will be is these two modes which affect the response

along the branch cut and enter into the final contribution of seismogram synthesis.

Following Watson (1972) we take the frequency to be real and allow the poles to wander on the complex wavenumber plane. The branch cuts are kept along the real and imaginary axes, as before. This differs from the cuts used by other studies (Gilbert, 1964;, Laster et al , 1965; Watson, 1972), in which the branch cuts were usually made directly in the first or fourth quadrant to expose the lower Riemann surfaces. Our object is to investigate the influence of leaky mode poles on the response function along the real branch cut.

It is already known (Gilbert, 1964) that the poles in the complex plane occur in sets of four, that is, if k is a pole, so are -k, k° , and $-k^{\circ}$. Here we will restrict ourselves to the fourth quadrant only. Keeping the branch cut along the real and imaginary axis, the radical ν_{α} has values in the fourth quadrant given by

$$\nu_{\alpha} = \pm \left[\sqrt{\frac{r+x}{2}} - i \sqrt{\frac{r-x}{2}} \right]$$
 (IV-1-1)

where

$$x = k_R^2 - k_I^2 - k_\alpha^2$$

$$y = 2 k_R k_I$$

$$\tau = \sqrt{x^2 + y^2}$$

A similar definition for $\nu_{m{\beta}}$ can be obtained by changing α to $m{\beta}$.

Poles were searched for on the lower complex-k sheets corresponding to different frequencies. When frequency changes, the leaky poles also transit along some paths, and sometimes even pass across the branch cut and appear on other Riemann sheets. Since our branch line integrations are carried out on the (+,+) sheet, we will consider only those poles which have the ability to enter this top sheet. When poles transit through the branch cut between $\;k_{lpha N}\;$ and $k_{eta N}\;$, only $\;
u_{lpha}\;$ changes sign, but through the cut between \emptyset and $k_{\sigma N}$ both signs of the radicals change. Hence the candidates are those poles just below the $k_{\alpha N} - k_{\beta N}$ cut on the (+,-) sheet and those near the cut between the origin and $k_{lpha N}$ on the (-,-) sheet. These two regions are called region I and region II by Watson (1972) The poles on the (-,+) sheet have no Pilant (1979). effect on the integration, however they might transit into (+,-) or (-,-) sheets and become important.

A computer program was set up to find the curves in a finite region of interest in the complex-k plane, such that the real or imaginary parts of the period equation are zero. The intersections of these curves are taken as the roots of the leaky modes. Since we are only interested in their properties, no refinement of

pole position is necessary. Figure 19 shows the curves of the null real part of the period equation (denoted by '+' sign), and curves with the null imaginary part (denoted by 'x') at frequencies $0.32~\mathrm{Hz}$ and $0.33~\mathrm{Hz}$. The model used is the SCM model listed in Table 1. These curves exhibit a particular pattern, which will help us to identify a specified pole as the frequency varies. The pattern obviously shows that two kinds of poles exist, the OP modes which always remain on the (+,-) sheet, follow an exponential type of path from $-i\infty$, and the PL modes which wander in the vicinity of the $k_{\alpha N} - k_{\beta N}$ branch cut.

Using these, we made a close search for poles by varying the frequency from 0.25 Hz to 0.40 Hz in steps of 0.01 Hz. The results are shown in Figures 20 to 22. Now the roles of the OP and PL modes are clear. In Figure 20 the OP poles migrate from the region with large imaginary k, i.e. high attenuation, and approach the point k_{α} . The PL modes, on the other hand, emerge from the (-,+) sheet by crossing the cut around k_{α_1} , and shift slowly in the region very close to the branch cut. These poles will contribute to the integration along the cut, which under some particular conditions significant enough to generate a dispersive are wavetrain between P and S, called PL waves. The PL OP modes collide at the place just beneath $\,k_{lpha_1}\,$. After this, two kinds of modes mingle together and generate

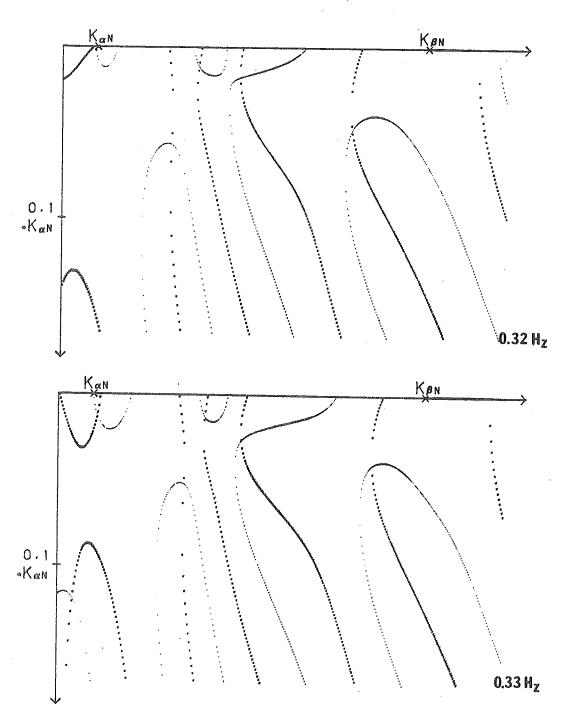


Figure 19. Curves of the null real part of Rayleigh wave period equation (denoted by '+' sign) and the null imaginary part (denoted by 'x' sign) in the fourth quadrant of (+,-) sheet of complex k plane. k_{α_N} and k_{β_N} are branch points. The top figure is obtained at 0.32 Hz and the bottom at 0.33 Hz. The SCM model is used.



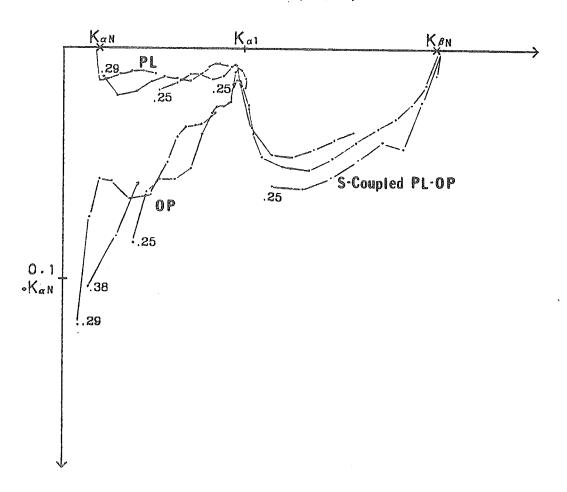


Figure 20. Paths of leaky modes of SCM model through the fourth quadrant of (+,-) sheet of complex k plane. The frequencies change from 0.25 Hz to 0.40 Hz. Two kinds of modes, namely PL and OP modes, exist before k_{α_1} the wavenumber corresponding to the first layer P velocity. After this point, two modes mingle together and form the shear-coupled PL-OP mode. The numbers at the beginning of each path indicate the starting frequencies.

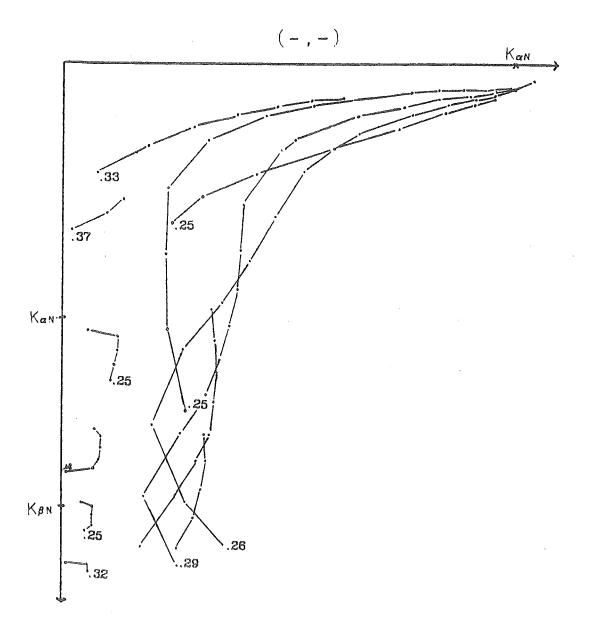


Figure 21. Same as Figure 20, but for leaky modes on (-,-) sheet of complex k plane.

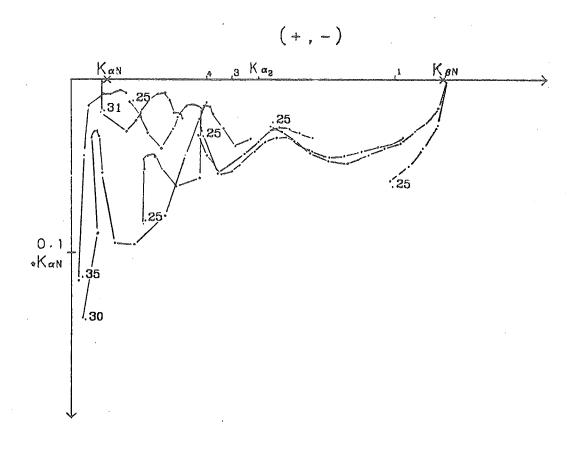


Figure 22. Same as Figure 20, but for five-layer CUS model.

the shear-coupled PL wave.

As the frequencies keep increasing, the poles will approach the shear branch point, then cross the cut, and enter the (+,+) sheet to become the normal modes. It is thus possible that, at some frequencies, the shear branch point is also a pole. One might suspect that the PL poles, when crossing the cut from (-,+) to (+,-) sheet, will give a singularity on our integration path. It is fortunate that this will not happen, because the integration is carried out on the (+,+) sheet and the cut which the PL modes cross is not the one we are integrating.

Figure 21 shows the leaky poles on the (-,-) sheet. Except around the P branch point, these poles have relatively large imaginary components. Hence we can expect that their contribution will be small. This is easy to understand since the waves from this part have large phase velocity, or equivalently travel nearly vertically. Hence the leakage of waves into the halfspace will carry away most of the energy. These poles will be named 'weak' leaky poles.

When the model becomes complex, the pole shifting also becomes complicated. Figure 22 shows the variation on the (+,-) sheet for the CUS model of Table 1. In this frequency range (0.25 to 0.4 Hz), the effect of the first weathered layer cannot be seen. The

variation between k_{α_2} and k_{α_N} is rapid; however, after k_{α_2} the shear-coupled PL modes retain their simple shifting pattern.

From the above discussion, we find that at least three kinds of poles which affect the values of the integrand along the real branch cut: leaky poles between 0 and k_{α_N} , (2) PL-OP poles between $k_{\alpha y}$ and k_{α_0} (3) shear-coupled PL between k_{α_2} and k_{β_N} . Figures 23 and 24 strongly support this conclusion. In Figure 23 pole positions responses integrand like those of Figure 3 displayed, with four plots corresponding to 0.25, and 1.0 Hz. Different source types are arbitrarily chosen. Figure 24 shows the same variations with different source type and corresponding leaky poles for the CUS model at 0.25 and Hz, respectively. As expected, the correspondence between the pole location and integrand variation apparent. It is obvious that the PL-OP poles are the most significant contributors to the integration, which give rise to the main part of the body waves with dominant P characteristics. The shear-coupled PL poles important only for several components such as ZSS. are These shear-coupled poles have an apparent tendency to associated with normal poles, which can be said to possess properties which are between 'pure' surface and body waves. The causality appearing in Figures

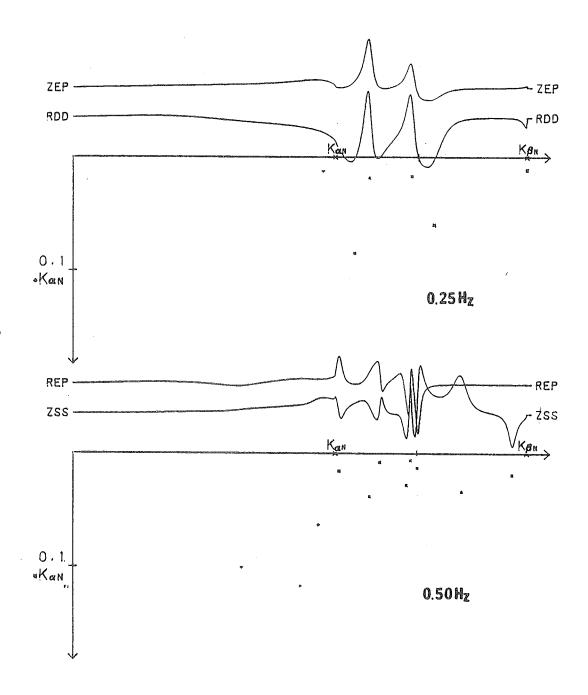


Figure 23. The effect of leaky modes on the variations of integrands along the real branch cut. 'x' denotes the modes on the (+,-) sheet and '+' denotes the modes on the (-,-) sheet. The response curves of integrands are obtained using the SCM model with the source at 10 km depth. The names of the integrand responses are the same as those in Figure 3. Four plots (two in the next page) correspond to the frequencies at 0.25, 0.50, 0.75, and 1.00 Hz, respectively.

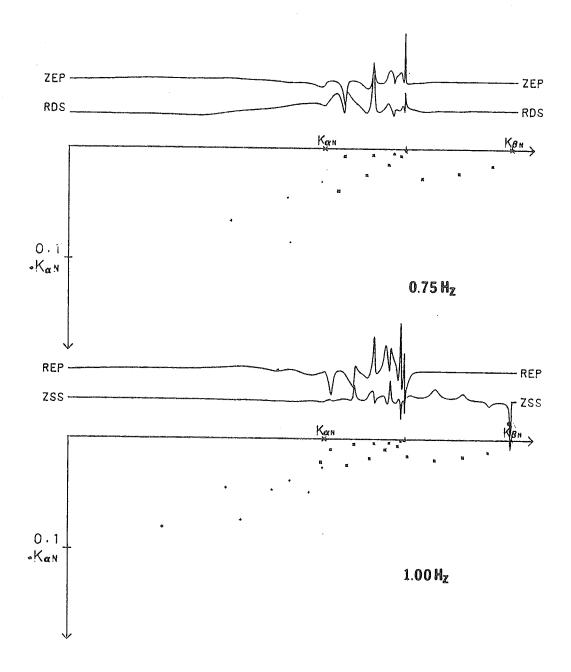


Figure 23. (cont'd)

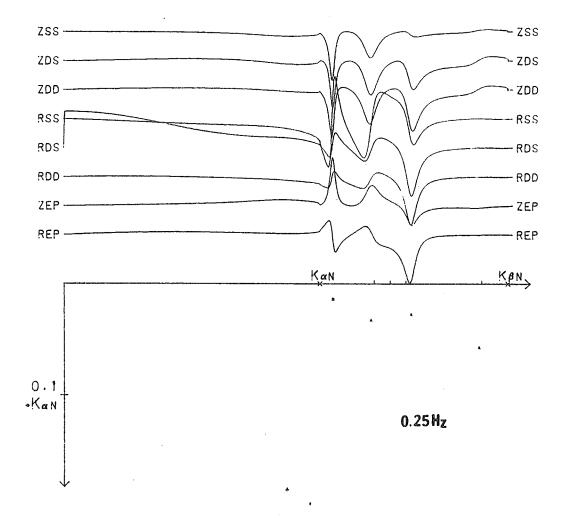


Figure 24. Results corresponding to Figure 23, but for the CUS model at the frequencies 0.25 and 0.75 Hz.

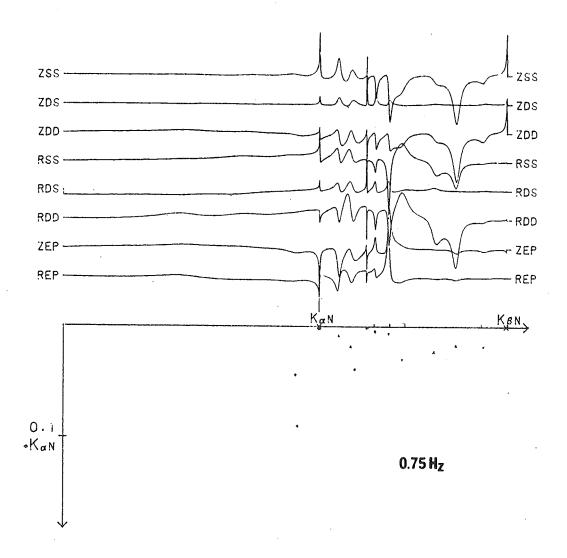


Figure 24. (cont'd)

16 to 18 arises from the contributions of these poles which compensate the Gibb's effect of abrupt truncation of normal modes at the S branch point. The contribution of weak leaky poles is relatively small, even though they might be very near the branch cut. This leads to the argument that the (-,-) cut is a stronger 'barrier' than the (+,-) cut.

For SH waves, the situation changes since there are no more PL poles. Figure 25 shows the migration of SH-OP modes, which have the properties of weak poles shear poles in the P-SV case. Their contributions are only in compensating the noncausality from normal In this sense, we might be able to say that the shear or shear-coupled leaky poles and the normal modes near the S branch point are the factors which constitute the S type body waves. In this range, it possible to define an absolutely distinguishing point for surface and body waves. A similar conclusion was obtained from ray expansion theory which supposes that body waves come from a finite number of rays, but when infinite rays are included they form the surface wave (Kennett, 1974). There is not a definite separation between body and surface waves.

After the study of the influence of leaky modes on the integrand, it is now clear how to perform branch line integration. The variable transform in equation

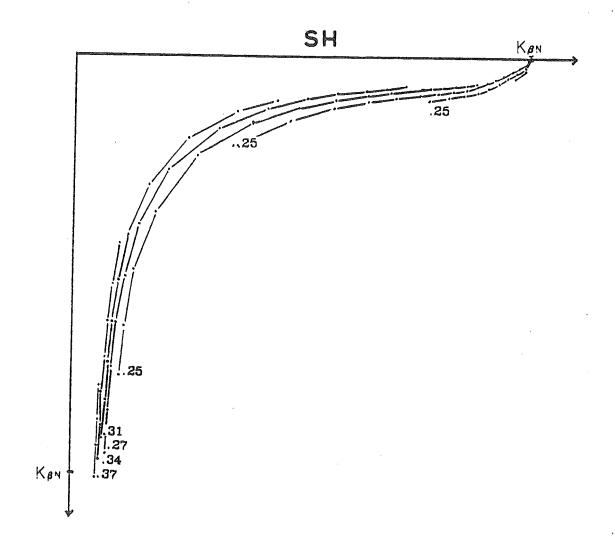


Figure 25. Paths of leaky modes of SCM model for the SH case. Other parameters are the same as Figure 20.

(II-4-5)just makes the sampling closer in the region where PL-OP modes are important, and thus improves the computational accuracy. Finally, the solution in equation (II-4-1) is determined not only from the integrand discussed above, but from the inner product of this integrand and the Hankel function. The Hankel function oscillatory function which depends on kr, i.e., is wavenumber times epicentral distance. The final contribution to synthesizing seismograms can be looked upon as a cross-correlation between these two oscillatory functions along the real wavenumber axis. Because of the relatively smooth variation of the SH component (Figure 3), the branch line integration contribution to the tangential component seismogram is significant only at short distances. However, for P-SV waves, the effect will persist to much greater distances, as seen from Figures 16 and 17.

4.2 Locked Mode Approximation

In the previous section we have discussed the relationship between branch line integrals and leaky modes as a way to improve the evaluation of body waves. An interesting method to treat this contribution, called the locked mode approximation, was introduced by Harvey (1981). He simply added a deep rigid cap to the bottom of the structure to 'catch' leaky modes. Since

the cutoff wavenumber for the locked modes is controlled by the half-space S-wave branch point, the branch cut shrinks and opens a place for more higher order 'locked' modes as the S-wave velocity in the bottom layer increases. These 'locked mode contributions' can easily be evaluated by normal mode superposition (Harkrider, 1964), which has been successfully used in synthesizing surface waves. Thus the method provides an easy-to-solve solution for the branch line integral or leaky mode contributions.

Another novel approach to the problem of line integration was proposed by Bouchon and Aki (1977) and Bouchon (1979, 1981). They made an attempt to discretize the wavenumber responses by presenting an infinite number of 'extra' sources at evenly grid points or rings. Because of the interference of waves from these sources, the integrands are quantified and form an exact solution to be evaluated at discrete wavenumber points. This technique is beyond the of this dissertation. Here we just simply discuss the method of locked mode approximation which, in addition, serves test for our eigenfunction solutions of as chapter III.

With the same methodology as used before, the integration is still made using real wavenumber and real frequency so that the search for poles is easier.

Nevertheless an anomalously high velocity cap layer is placed at the bottom of the structure. Such a process just serves to 'squeeze' the leaky modes in the lower Riemann sheets into the real 'trapped' mode position. The cap layer not only traps most of the seismic energy in the upper layers, but also brings in those pulses not existing in the real structure. If the cap layer is situated at a depth so large that the energy reaching it is small compared to the energy in the near surface layers, these false phases can be isolated and filtered by a wavenumber window (Embree et al , 1963). However for the test of normal mode theory of chapter III, we will not consider this filtering, so as to pursue the normal mode superposition method faithfully.

When using the locked mode approximation, several requirements should be fulfilled beforehand:

(1) The calculation of locked mode positions must be precise. Figure 26 shows the pole positions along the real wavenumber axis for a cap layer structure. The model is SCM listed in Table 1 with a cap layer at 240 km depth and P velocity 20 km/sec, S velocity 10 km/sec, and density 6 gm/cm³. As the regular normal modes are still kept at the same places, the surface wave contribution is not affected. However, the newly generated 'locked' poles are numerous and very close to each other. For example, 15 modes for the original

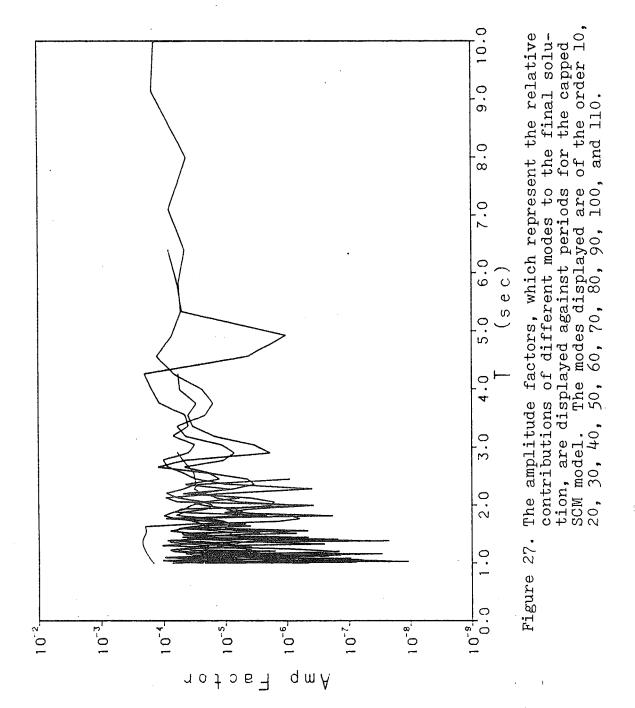
Kan	KbN	x x	0.25 Hz
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NI NICHERON PRINCES NEED IN NICK NICK NICK NICK NICK NICK NICK N	W H W R N M N M N AMERICANSAGEMENTERSECTIONS A A A A	x x x x <u>,</u> x 499	0.75 H _Z
- HEREN BERNES BERNES MEN M.	SR H N.A. MICCOSCORDEGICOSCORDEGICOSCORDIS		1.0 H _z

Figure 26. The positions of poles along the real k-axis at the frequencies 0.25, 0.50, 0.75, and 1.00 Hz. k_{α_N} and k_{β_N} are branch points for SCM model without the cap layer. When the cap layer is added, the leaky modes are forced to migrate into the normal mode positions as those shown to the left of $k_{\beta N}$. These created 'locked' modes are numerous and are difficult to locate. The positions of regular normal modes are essentially not affected by the present of the cap layer.

model at 1 Hz become 119 modes for the capped layer model. The treatment of these pseudo-normal modes requires special care.

- (2) The pole contribution arising from the partial derivative of the period equation (or equivalently the amplitude factor defined in equation III-2-2) is not easy to evaluate since these poles have very large phase velocities and cause numerical difficulty. Thus we use equation (III-2-2) instead of the alternate expression for A_R .
- (3) Because of the large number of poles involved, it is important to pay attention to the computational efficiency. In the previous section, we have found a different importance for each kind of leaky mode. Is there any difference of contribution from these locked poles? One test is shown in Figure 27, where the amplitude factors for mode order 10, 20, 30, 40, 50, 60, 70, 80, 90, 100, and 110 are displayed. This figure reveals that each mode has about the same level of contribution. Hence, unfortunately, all of the locked modes must be taken into account.

To overcome these difficulties, a new technique other than the usual normal mode method (Harkrider, 1964; Saito, 1972) is required. Harvey (1981) used Abo-Zena's (1980) formulation to calculate displacement eigenfunctions, which in turn give the stress



eigenfunctions, by a constraint derived from the relation between the displacements and stresses. Similarly after the study of normal mode theory in chapter III, we also raised the eigenfunction theory to a very powerful stage. This formulation will be used to consider the idea of locked mode approximation. On the other hand, the locked mode approximation may provide a good test of the stability of our method. Figures 28 to 30 present the results.

Figure 28 illustrates the results from methods chapters II and III, as compared to the locked mode approximation. In this figure, (a) is the locked mode the branch line integration and approximation, (b) pole, (c) the normal pole contribution, and (d) branch line integral. The display is the ZSS component due to a point source at 10 km depth in the SCM model with highest frequency 1 Hz and source time function (equation II-4-8) $\tau = \emptyset.5$ sec. The agreement (b) is excellent except for some artificial early arrivals due to the fact that the cap effectively introduces a sharp wavenumber cutoff in the solution. This means that the methods we developed III can be trusted. Figure 29 is the chapters II and RDS component for the same case, and Figure 30 TDDcomponent. The match of results from the locked mode approximation to the complete seismogram is obvious. However the computation times are quite

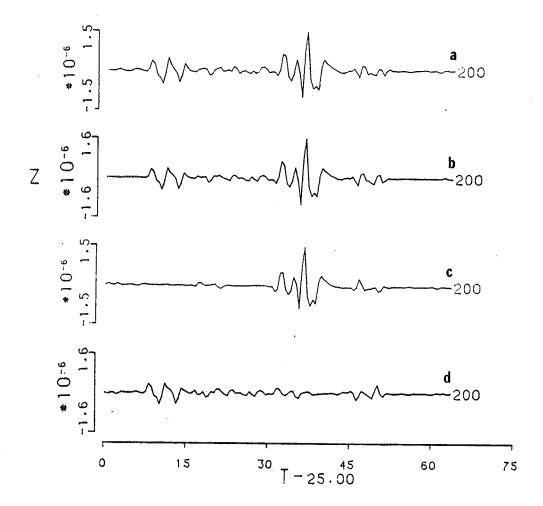


Figure 28. Comparison of locked mode approximation to the complete solution using the method of chapter 2.

(a) the locked mode approximation solution; (b) the wave integral complete solution; (c) the pole contribution; (d) the branch line integral contribution. The SCM model with a strike-slip dislocation source at the depth of 10 km is used. The cap layer is located at 200 km deep and has a P velocity of 20 km/sec, S velocity of 10 km/sec, and density of 6 gm/cm³.

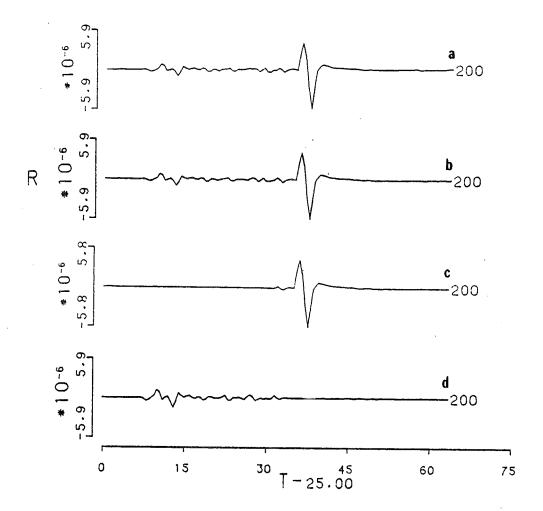


Figure 29. Results corresponding to Figure 28, but for the radial component and a dip-slip dislocation source.

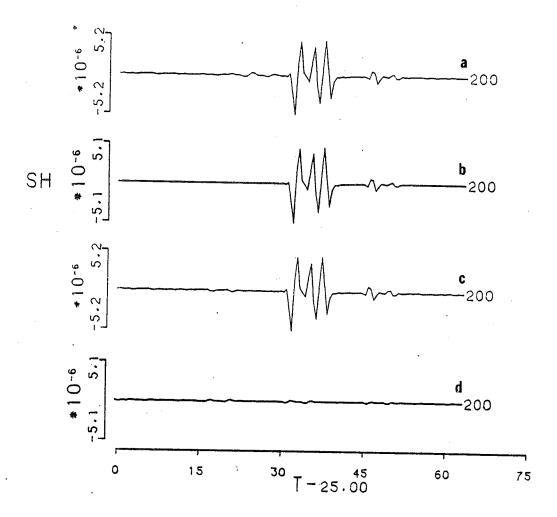


Figure 30. Results corresponding to Figure 28, but for the tangential component and a 450 dip-slip source.

different. Using the locked mode approximation the calculation takes about 2 hours on a DEC PDP 11/70 minicomputer, while using the direct integration method requires only half an hour. Since these two methods use the same wave theory, any time consuming aspect, such as the case of high frequencies, will be required for both methods. The advantage of the locked mode approach is that once the dispersion curves are found for a given model, they do not have to be recomputed.

4.3 Reflection Method and Reflectivity Method

In the previous chapters, we built up the whole field wave by summing the layer responses over wavenumber and frequency. The calculation includes all details of energy possibly excited in the layered medium. Sometimes, for the study of particular portions of the structure, we desire to suppress the reflections from certain interfaces. This is especially useful for body wave studies. This section will develop a method to 'eliminate' the reflections from some layer boun-The significance of the pole contribution to the generation of surface waves becomes apparent if the reflections from the free surface, the strongest reflector, are artificially suppressed. Two theories closely related to this section are worth reviewing. One is Kennett's reflection and transmission coefficient method, which is referred to as the 'reflection method' here (Kennett and Kerry, 1979), and the other is Fuchs's reflectivity method (Fuchs and Müller, 1971). In this section, these methods will be stated using the terminology of this dissertation.

Reflection Method

The reflection method introduced by Dr. Kennett and his colleagues (Kennett, 1974; Kennett et al, 1978; Kennett and Kerry, 1979; Kennett, 1980; and Kerry, 1981) possesses properties of both ray theory and wave theory. They established a connection between conventional matrix methods and the reflection and transmission properties of a single interface. This approach lends itself to a ray interpretation, however, from a view point of gross reflection and transmission response of layers. Starting from the response of the entire stack of layers, the Haskell matrix R in equation (II-2-6):

$$R = E_N^{-1} \alpha_{N-1} \cdots \alpha_1.$$

 α 's, the layer matrices, can be expressed in terms of the fundamental matrix E and phase matrix Λ (equation II-1-18),

$$\alpha_n = \mathbf{E}_n \ \Lambda_n \ \mathbf{E}_n^{-1} \ .$$

Hence,

$$R = (E_{N}^{-1} E_{N-1}) \Lambda_{N-1} (E_{N-1}^{-1} E_{N-2}) \cdots \cdots$$

$$\cdots \Lambda_{n} (E_{n}^{-1} E_{n-1}) \Lambda_{n-1} \cdots (E_{2}^{-1} E_{1}) \Lambda_{1} E_{1}^{-1} . \qquad (IV-3-1)$$

The reflection and transmission can only occur at velocity discontinuities; therefore only the terms within the parenthesis of equation (IV-3-1) contain the information of wave-interface interaction. The matrix Λ is only used to 'phase relate' the waves through the layer. Consider a single interface. The potential-constant vector K, which represents the waves inside the layer, can be rewritten as

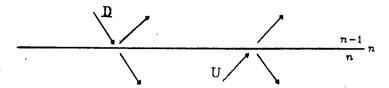
$$\mathbf{K}_{n} = \begin{bmatrix} A \\ B \\ A \\ B \end{bmatrix}_{n} = \begin{bmatrix} \mathbf{U} \\ \mathbf{D} \end{bmatrix}_{n}$$

where U and D are 2xl matrices for the P-SV case containing upgoing P,SV and downgoing P,SV potentials, respectively. For the SH case, U and D are scalars. The waves just above and below the interface satisfy

$$\begin{bmatrix} \mathbf{U} \\ \mathbf{D} \end{bmatrix}_{n} = \mathbf{E}_{n}^{-1} \mathbf{E}_{n-1} \begin{bmatrix} \mathbf{U} \\ \mathbf{D} \end{bmatrix}_{n-1} = \mathbf{F}_{n} \begin{bmatrix} \mathbf{U} \\ \mathbf{D} \end{bmatrix}_{n-1}.$$

The underscoring denotes the quantities at the bottom of the n-1 layer. F_n is called the reflection-and-transmission matrix. It is noted that F_n does not contain the phase terms. These terms, as discussed in section 2.2, might be exponentially growing and cause an accuracy problem.

Consider a downgoing wave incident at the boun-dary,



we define

$$\mathbf{U}_{n-1} = \mathbf{r}_D \ \mathbf{D}_{n-1}$$

$$\mathbf{D}_n = \mathbf{t}_D \ \mathbf{D}_{n-1}$$

 \mathbf{r}_D and \mathbf{t}_D are matrices of the reflection and transmission coefficients for both P and SV waves in downward propagation,

$$\mathbf{r}_D = \begin{bmatrix} \tau^D_{pp} & \tau^D_{sp} \\ \tau^D_{ps} & \tau^D_{ss} \end{bmatrix} \qquad \mathbf{t}_D = \begin{bmatrix} t^D_{pp} & t^D_{sp} \\ t^D_{ps} & t^D_{ss} \end{bmatrix}$$

Similarly, an upgoing wave incident at the boundary defines \mathbf{r}_U , \mathbf{t}_U . Following the discussion of Kennett (1974), or more clearly Frasier (1970), we have

$$\mathbf{F}_{n} = \mathbf{E}_{n}^{-1} \mathbf{E}_{n-1} = \begin{bmatrix} \mathbf{t}_{\bar{U}}^{1} & -\mathbf{t}_{\bar{U}}^{1} \mathbf{r}_{D} \\ & & \\ \mathbf{r}_{U} \mathbf{t}_{\bar{U}}^{1} & \mathbf{t}_{D} - \mathbf{r}_{U} \mathbf{t}_{\bar{U}}^{1} \mathbf{r}_{D} \end{bmatrix}_{n}$$
(IV-3-2)

and

$$\mathbf{r}_{D} = -\mathbf{F}_{11}^{-1}\mathbf{F}_{12}$$

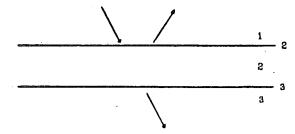
$$\mathbf{t}_{D} = \mathbf{F}_{22} - \mathbf{F}_{21}\mathbf{F}_{11}^{-1}\mathbf{F}_{12}$$

$$\mathbf{r}_{U} = \mathbf{F}_{21}\mathbf{F}_{11}^{-1}$$

$$\mathbf{t}_{U} = \mathbf{F}_{11}^{-1}$$
(IV-3-3)

Note that $det(F_{11})$ is the Stoneley function.

If the waves pass through a layer from the bottom of layer 1 to the top of layer 3;



the response for potential-constant vectors through the layer becomes

(
$$E_3^{-1}E_2$$
) Λ_2 ($E_2^{-1}E_1$) .

The situation does not become complex, since an iterative form can be found (Kennett, 1974):

$$\begin{split} \mathbf{r}_{D_{31}} &= \mathbf{r}_{D_2} + \mathbf{t}_{U_2} (\mathbf{P}^{-1} \mathbf{r}_{D_3} \mathbf{P}^{-1}) \left[1 - \mathbf{r}_{U_2} (\mathbf{P}^{-1} \mathbf{r}_{D_3} \mathbf{P}^{-1}) \right]^{-1} \mathbf{t}_{D_2} \\ \mathbf{t}_{D_{31}} &= \left(\mathbf{t}_{D_3} \mathbf{P}^{-1} \right) \left[1 - \mathbf{r}_{U_2} (\mathbf{P}^{-1} \mathbf{r}_{D_3} \mathbf{P}^{-1}) \right]^{-1} \mathbf{t}_{D_2} \\ \mathbf{r}_{U_{3\underline{1}}} &= \mathbf{r}_{U_3} + (\mathbf{t}_{D_3} \mathbf{P}^{-1}) \mathbf{r}_{U_2} \left[1 - (\mathbf{P}^{-1} \mathbf{r}_{D_3} \mathbf{P}^{-1}) \mathbf{r}_{U_2} \right]^{-1} (\mathbf{P}^{-1} \mathbf{t}_{U_3}) \\ \mathbf{t}_{U_{3\underline{1}}} &= \mathbf{t}_{U_2} \left[1 - (\mathbf{P}^{-1} \mathbf{r}_{D_3} \mathbf{P}^{-1}) \mathbf{r}_{U_2} \right]^{-1} (\mathbf{P}^{-1} \mathbf{t}_{U_3}) \end{split}$$

where

$$\mathbf{P} = \begin{bmatrix} e^{\nu_{\alpha} d_z} & 0 \\ 0 & e^{\nu_{\beta} d_z} \end{bmatrix} ,$$

and $3\underline{1}$ represents the stack between the top of layer 3 and the bottom of layer 1. In equation (IV-3-4) we have used P to phase relate the reflection and transmission coefficients at interface 3 to interface

- 2. However, note that r_{U_3} need not be 'phase related'. Equation (IV-3-4) reveals some interesting features:
 - (1) The term $[1-\mathbf{r}_U(\mathbf{P}^{-1}\mathbf{r}_D\mathbf{P}^{-1})]^{-1}$ can be expanded as a power series;

 $[1-\mathbf{r}_{U}(\mathbf{P}^{-1}\mathbf{r}_{D}\mathbf{P}^{-1})]^{-1}=1+\mathbf{r}_{U}(\mathbf{P}^{-1}\mathbf{r}_{D}\mathbf{P}^{-1})+\mathbf{r}_{U}(\mathbf{P}^{-1}\mathbf{r}_{D}\mathbf{P}^{-1})\mathbf{r}_{U}(\mathbf{P}^{-1}\mathbf{r}_{D}\mathbf{P}^{-1})+\cdots$

which describes the ray propagation with different internal reflections and transmissions in the layer. Kennett (1980) has provided a detailed explanation about this ray viewpoint of wave theory. Hence, the total response of layers is just a superposition of multiple reflections from boundaries.

- (2) The phase term which might grow exponentially is excluded from the calculation of reflection and transmission coefficient at layer boundaries. Hence the problem of loss of precision can be well controlled.
- (3) To calculate R, we need only start at the base of the layers, progress upward, iteratively adding a layer to the stack, one at a time. This iterative approach represents the most important step of Kennett's method.

Suppose that the stacking from the half-space to the top of first layer (just beneath the surface) has been accomplished. The (1,1) component of 2x2 partition of matrix R can be determined from

$$\mathbf{R} = \begin{bmatrix} \mathbf{t}_{\bar{U}^1} & -\mathbf{t}_{\bar{U}^1} \mathbf{r}_D \\ \vdots & \ddots & \vdots \end{bmatrix}_{N_1} \begin{bmatrix} \overline{\mathbf{E}}_{11} \\ \overline{\mathbf{E}}_{21} \end{bmatrix}_1,$$

i.e.,

$$\mathbf{R}_{11} = \mathbf{t}_{\bar{U}_{N1}}^{1} \left(\overline{\mathbf{E}}_{11} - \mathbf{r}_{\bar{D}} \overline{\mathbf{E}}_{21} \right)$$

where \overline{E} represents the inverse of the E matrix of the first layer. Going back to our formulas for the displacement at the free surface (equation II-2-5),

$$\begin{bmatrix} 0 \\ 0 \end{bmatrix}_{N} = X S + \begin{bmatrix} R_{11} & R_{12} \\ \vdots & \vdots & \vdots \\ R_{21} & R_{22} \end{bmatrix} \begin{bmatrix} W_{1} \\ \vdots \\ 0 \end{bmatrix}_{1}$$

or

$$\mathbf{W}_{1} \equiv \begin{bmatrix} U_{r_{1}} \\ U_{x_{1}} \end{bmatrix} = -\mathbf{R}_{11}^{-1} \mathbf{X}_{1i} \mathbf{S}_{i} . \tag{IV-3-5}$$

If we use the discontinuity in the constant - potential vector to describe the source (their explicit form will be listed in section 5.1) rather than a discontinuity in displacement and stress,

$$\mathbf{X} \mathbf{S} = \mathbf{E}_{N}^{-1} \mathbf{E}_{N-1} \Lambda_{N-1} \mathbf{E}_{N-1}^{-1} \cdots \Lambda_{m} (\mathbf{E}_{m}^{-1} \mathbf{S})$$

$$= (\mathbf{E}_{N}^{-1} \mathbf{E}_{N-1} \Lambda_{N-1} \cdots \Lambda_{m}) \Sigma_{m}$$

$$= \begin{bmatrix} \mathbf{t}_{\bar{U}}^{-1} & -\mathbf{t}_{\bar{U}}^{-1} \mathbf{r}_{D} \\ \vdots & \ddots & \vdots \\ \Sigma_{D} \end{bmatrix}_{m}$$

and equation (IV-3-5) becomes

$$\mathbf{W}_{1} = -\left(\overline{\mathbf{E}}_{11} - \mathbf{r}_{D_{N1}}\overline{\mathbf{E}}_{21}\right)^{-1} \mathbf{t}_{U_{N1}} \mathbf{t}_{\bar{U}_{NS}}^{1} \left(\Sigma_{U} - \mathbf{r}_{D_{NS}}\Sigma_{D}\right). \tag{IV-3-6}$$

 $det(\overline{\mathbb{E}}_{11}-r_{D_{N1}}\overline{\mathbb{E}}_{21})$ is the period equation. Note that, except for the source depth, any partition of layer stacking at other places such as receiver depth has been avoided. There exist several different forms of the period equation in the derivation (Kennett, 1974; Kennett, 1980; Kennett and Kerry, 1979; and Kerry, 1981). After lengthly discussion, Kerry (1981) finally chose this form to calculate dispersion values. However, by using a different partition of layer stacking, he was able to find an interesting screen effect of a low velocity zone at depth.

We will not expand equation (IV-3-6) to calculate seismograms. Only the concept of reflection and transmission coefficients or equation (IV-3-2) will be used in the following development for the reflection suppression technique.

Reflectivity Method

The reflectivity method is an important application of the Haskell matrix (Fuchs, 1968; Fuchs and Müller, 1971; Kind and Müller, 1975; Müller and Kind, 1976; and Kind, 1976, 1978, 1979), hitherto mostly used for body wave computation for an explosive type of source. This method involves a double integration with respect to real wavenumber and real frequency, but over a small range of interest such as those with real

number of incident angle. Such a numerical integration corresponds only to our real branch line integral. The integration could cover the whole range of response including the surface-wave poles by adding a small complex part to the velocity in each layer, accounting for anelasticity, which would move poles off the positive k-axis (Kind, 1979). However, one has to sample the integrands very closely over a much larger range of angle of incidence, and the computation time increases substantially. The reason for dense sampling is similar to that which we have discussed in section 4.1, i.e., the influence of leaky or normal modes on the integration along the real wavenumber axis. In principle, the reflectivity method is not suitable for the computation of normal modes.

Consider the potential-constant vector K just beneath the free surface. If there is no reflection at the free surface, we can simply make $D_1 = \emptyset$, and the free surface displacement is just $W_1 = E_{11}U_1$, where E_{11} is, for P-SV, the (1,1) component of the 2 by 2 partition of the Haskell E matrix of the top layer, and for SH, a scalar. Fuchs (1968) used this form to treat the free surface in the original theory of the reflectivity method. However, if we take into account the free surface effect, but do not allow the wave to be reflected back into the earth, then from equation (II-1-15),

$$\begin{bmatrix} \mathbf{W}_1 \\ \mathbf{0} \end{bmatrix} = \begin{bmatrix} \mathbf{E}_{11} & \mathbf{E}_{12} \\ \mathbf{E}_{21} & \mathbf{E}_{22} \end{bmatrix}_{\mathbf{1}} \begin{bmatrix} \mathbf{U} \\ \mathbf{D} \end{bmatrix}_{\mathbf{1}},$$

and the solution is

$$W_1 = (E_{11} - E_{12} E_{22}^{-1} E_{21}) U_1 = F_1 U_1 . (IV-3-7)$$

This is the form Fuchs and Müller (1971) used. For SH such a modification just doubles the amplitude, and for P-SV some conversions between P and SV take place but without much effect.

For a reflecting and transmitting interface, we can simply set \mathbf{r}_U and \mathbf{r}_D equal to zero to suppress its 'reflectivity'. From equation (IV-3-2) the reflection-and-transmission matrix F becomes

$$\mathbf{F}_{n} = \mathbf{E}_{n}^{-1} \mathbf{E}_{n-1} = \begin{bmatrix} \mathbf{t}_{\bar{U}}^{1} & 0 \\ 0 & \mathbf{t}_{\bar{D}} \end{bmatrix}_{n}$$

Furthermore under this condition, (IV-3-3) provides that

$$\mathbf{t}_{II}^{-1} = \mathbf{F}_{11} .$$

Because of the reciprocity relation, Kennett <u>et al</u> (1978) and Fraiser (1970) found that

$$\mathbf{t}_{D} = \mathbf{t}_{U}^{T} \times (\text{normalization factors})$$
.

In our system, the normalization factors, as discussed in deriving equation (II-1-12), enter the result as

$$\mathbf{t}_D = \begin{bmatrix} t_{pp}^D & t_{sp}^D \\ t_{ps}^D & t_{ss}^D \end{bmatrix} = \begin{bmatrix} t_{pp}^U \alpha & t_{ps}^U b \\ t_{sp}^U c & t_{ss}^U d \end{bmatrix}$$

with

$$a = \frac{v_{\alpha}/\rho}{v_{\alpha}/\rho} \quad b = \frac{v_{\beta}k^2/\rho}{v_{\alpha}/\rho} \quad c = \frac{v_{\alpha}/\rho}{v_{\beta}k^2/\rho} \quad d = \frac{v_{\beta}/\rho}{v_{\beta}/\rho},$$

where primes denote the quantities just above the interface and the unprimed quantities are just beneath the interface. The matrix F takes the form:

$$\mathbf{F} = \begin{bmatrix} F_{11} & F_{12} & 0 & 0 \\ F_{21} & F_{22} & 0 & 0 \\ 0 & 0 & \frac{F_{22}}{\Delta} \alpha & -\frac{F_{21}}{\Delta} b \\ 0 & 0 & -\frac{F_{12}}{\Delta} c & \frac{F_{11}}{\Delta} d \end{bmatrix}$$
 (IV-3-8)

where

$$\Delta = F|_{12}^{12}$$

Now let us stack a whole set of layers together. Use equation (II-2-5)

$$K_N = X S + R B_1$$

= $X S + (R E_1) (E_1^{-1}B_1) = X S + Q K_1$

where

$$\mathbf{Q} = \mathbf{R} \mathbf{E}_1 = \mathbf{X} \mathbf{Z} \mathbf{E}_1 = \mathbf{X} \mathbf{P}$$

$$= \mathbf{E}_{N}^{-1} \mathbf{E}_{N-1} \Lambda_{N-1} \mathbf{E}_{N-1}^{-1} \cdots \Lambda_1$$

$$\mathbf{P} = \mathbf{E}_{m} \Lambda_{m} \mathbf{E}_{m}^{-1} \cdots \Lambda_1$$

Following the same procedures between equation (II-2-6) and (II-2-8); we have

$$U_{1} = \begin{bmatrix} -S_{i}X \mid_{ij}^{12} P_{j2} \\ \\ S_{i}X \mid_{ij}^{12} P_{j1} \end{bmatrix} / Q \mid_{12}^{12}$$

and the free surface displacements are

$$\begin{bmatrix} U_{r_1} \\ U_{z_1} \end{bmatrix} = \mathbf{F}_1 \begin{bmatrix} -S_i X \mid_{ij}^{12} P_{j2} \\ S_i X \mid_{ij}^{12} P_{j1} \end{bmatrix} / Q \mid_{12}^{12}$$
(IV-3-9)

where

$$\mathbf{F}_{1} = \mathbf{E}_{11} - \mathbf{E}_{12} \, \mathbf{E}_{22}^{-1} \, \mathbf{E}_{21} = \begin{bmatrix} -\frac{2\gamma\nu_{\alpha}\nu_{\beta}}{k} & -2k\,\nu_{\beta}(\gamma-1) \\ \\ 2\frac{\nu_{\alpha}}{\rho}(\gamma-1) & \frac{2\gamma\nu_{\alpha}\nu_{\beta}}{\rho} \end{bmatrix}_{1} / \rho_{1}(\frac{\gamma\nu_{\alpha}\nu_{\beta}}{k^{2}} - (\gamma-1)^{2})_{1}.$$

To calculate Q, the form (IV-3-8) is used if the reflection is to be suppressed. If it is not to be suppressed, the layer matrix α can be used. In the transmission zone of the reflectivity method, all of the reflections are suppressed, and in the reflection zone, the normal case is retained. Hence, Q in the reflectivity method is

$$Q = E_N^{-1} \alpha_{N-1} \cdots \alpha_n \Lambda_{n-1} F_{n-1} \cdots \Lambda_2 F_2 \Lambda_1$$

Some aspects of our system for the reflectivity method are interesting to point out:

- (1) It is just a simplified case for our system but is a generalization of the reflectivity method.
- (2) The conversion of waves during passage through the interface is involved. If such a conversion is not of interest, we need just set the off-diagonal components in the matrix F (equation IV-3-8) to zero. This is the approach used in the usual re-

flectivity method.

- (3) Any kind of point source, namely explosion, double couple, or other, is permitted, since the source term S in equation (IV-3-9) is left in a general form.
- (4) All benefits we have discussed for our system are also applied to this new reflectivity method.

Synthetic Seismogram

Several synthetic seismograms were made to illustrate the layer reflection suppression technique by applying the integration method of chapter II. 31 shows a simple test using the SCM model and an explosive source at a depth 10 km. Figure 31a radial component complete seismogram which is used to compare to Figure 31b, in which the free-surface reflection is 'suppressed'. As expected, Figure 31b just exhibits some easily identified arrivals such Pn, Sn, P, PP, PS which are reflected or refracted back from the crust-mantle boundary. All other multiple phases as well as surface waves are eliminated. Figure 32 shows the similar comparison for the z component. Since a compressional type of source is used, there are no obvious direct shear signals.

For the test of reflection suppression in lower boundaries, an intermediate layer with $\alpha=6.70$ km/sec $\beta=3.87$ km/sec, and $\rho=3.0$ gm/cm³ is inserted between

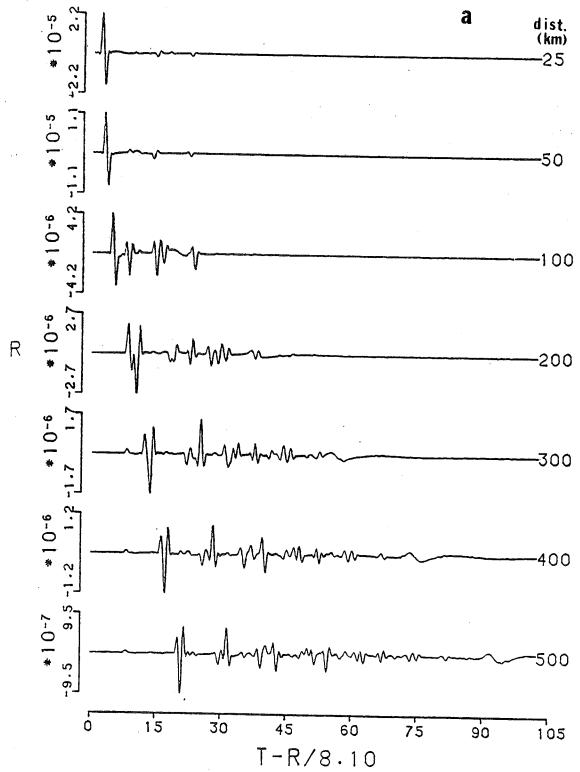


Figure 31a. Radial component complete seismograms. The SCM model and an explosive source at 10 km depth are used. A distance range of 25-500 km is presented. Multiple reflections and surface waves are well-developed, especially for large distances.

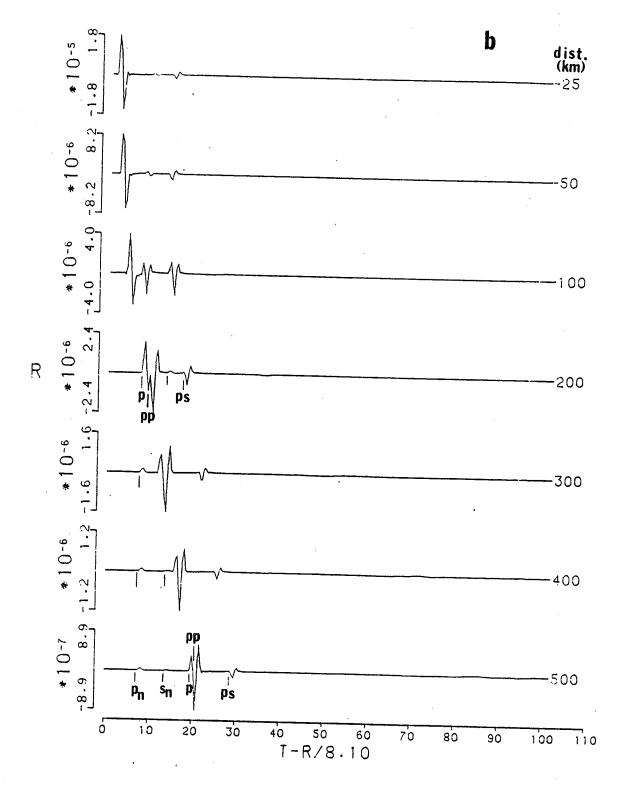


Figure 31b. Results using the same source and model of Figure 31a, but the reflectivity of the free surface is suppressed. Some identified arrivals from the crust-mantle boundary are indicated.

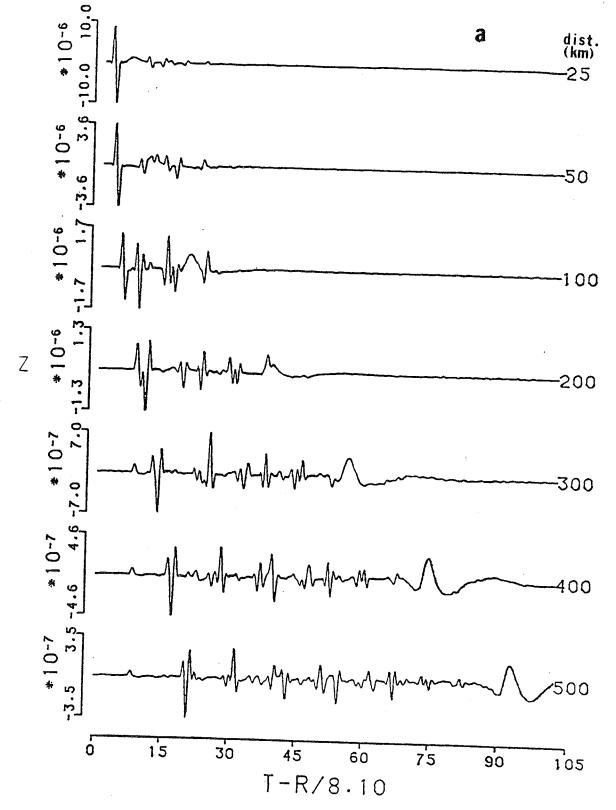


Figure 32. Results for the same source and model of Figure 31, but for the vertical component.

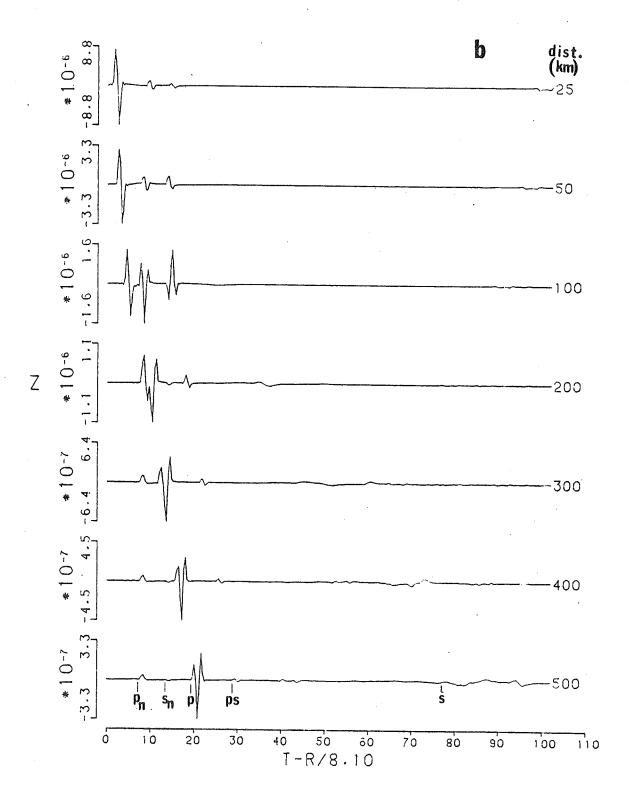


Figure 32. (cont'd)

and 40 Km in SCM model. Several seismodepths $2\emptyset$ km grams at epicentral distances of 200 km and 300 km, due strike-slip source, are displayed in Figure 33, where (a) is for the SCM model, (b) is for the modified and (c) is for the modified SCM model with reflection-deprived second layer boundary. Receivers located at an azimuth of the Ø degrees from the strike of fault. Because of the small velocity contrast across the reflection-deprived boundary, there is not much difference between (b) and (c) for the R com-However for the SH component (Figure 34), the effect is much more apparent. This means that the shear wave is more sensitive to changes in transverse reflection and transmission response than compression and SV waves. This finding might be significant for the development of shear wave exploration techniques. Seismograms (a) (c) in Figures 33 and 34 are not and comparable except for arrival time. Hence the choice of correct model in seismogram simulation is important.

If the free surface reflection of the modified SCM model is further suppressed, the seismogram from a dislocation source will be deprived of most of its wave energy, and numerical noise becomes apparent. We do not plot this result; instead, an explosive source is used for the example. Figure 35 shows the seismograms at 100, 200, 300 km epicentral distances for the modified SCM model. The top traces are complete seismograms and

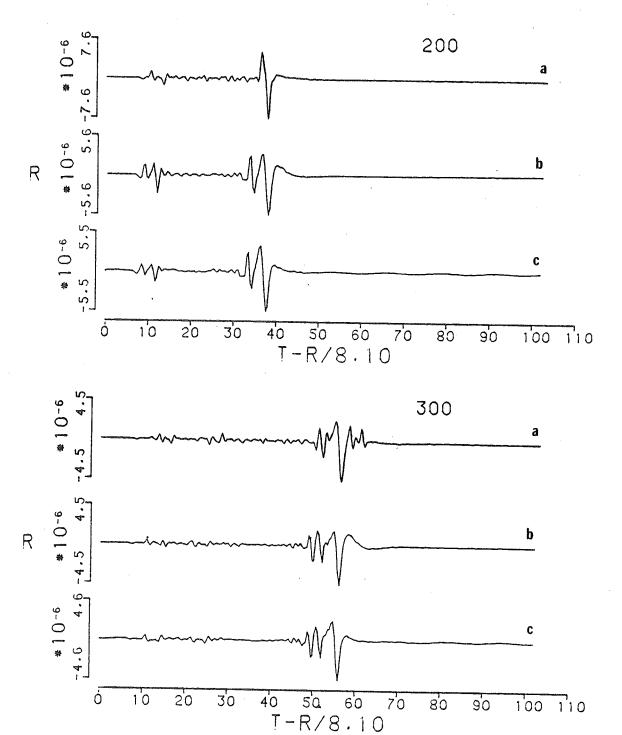


Figure 33. The effect of reflection suppression. The displays are for the radial component due to a strike-slip source buried at 10 km depth, and for stations at distances of 200 and 300 km. In each set, (a) is for the two-layer SCM model; (b) is for the three-layer modified SCM model; and (c) is for the modified SCM model with the reflections from the secondary interface suppressed.

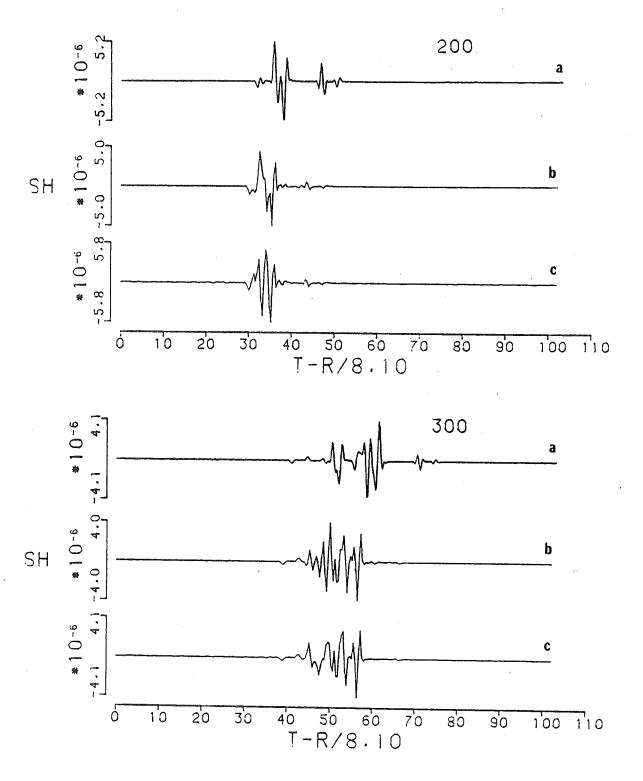


Figure 34. Results using the same source and model of Figure 33, but for the tangential component.

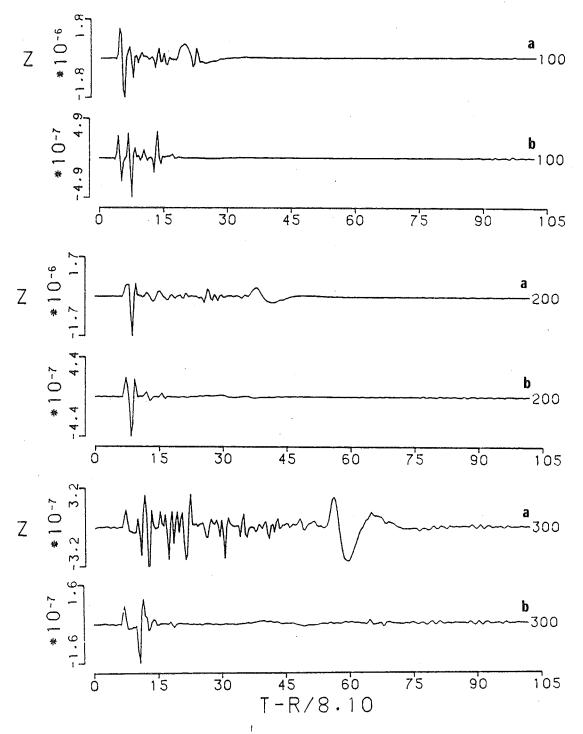


Figure 35. (a) Vertical component complete seismograms due to an explosive source buried at 10 km depth for the modified SCM model. (b) is the seismogram for the same model but with the reflections from the free surface and secondary interface suppressed. Compare (b) with Figure 32b.

the bottom traces are for reflections only from the crust-mantle boundary. The surface waves are all cancelled. Only those phases shown in Figure 32b remain.

CHAPTER V

SOURCE AND INSTRUMENT

For the completeness of the system developed in this dissertation, two factors, namely source and receiver, should also be included. The literature discussing sources is vast, and we will touch on only the major points needed to synthesize seismograms. For modeling the instrument, a method from Mitchell and Landisman (1969) is revised to more efficiently retrieve the instrument response parameters. These parameters are included in the final stage to make up the synthetic seismograms.

5.1 Source Considerations

The source factor can be included in the solutions by simply adding a singularity to the wave equation. Since the equivalence between body force expression and the discontinuity of displacement and strain at the source location was demonstrated (Burridge and Knopoff, 1964), the development of source theory has become much easier by the use of Green's functions. A Green's function is a solution of a differential equation involving the source singularity, which is independent of the form of the source function and is determined

only by the differential equation and the boundary conditions. One useful application of the Green's function is that the solutions from an arbitrary source, for example a single force f, can be set up by convolving the Green's function and this source function:

$$u_i = G_{ij} * f_j.$$

If the source becomes more complicated, such as second order sources (couple or dipole), more derivatives with respect to the direction are involved:

$$u_i = -u_{i,k} n_k ,$$

where n_k represents the direction. These extra directional dependences can be properly described by a moment tensor M;

$$u_i = G_{ij,k} * M_{jk}$$

A detailed discussion of the concept of a moment tensor is given in Aki and Richards (1980, chapters 3 and 4). In a cylindrical coordinate system, the Green's function of the wave equation includes the kernels in three directions, i.e., r, z, and v. Hence the Green's function itself also contains the directivity terms v. Usually the Green's function is not easy to find. For the purpose of extracting the source dependent terms, a whole space solution from Stoke's formula (Aki and Richards 1980, equation 4.23), which has been extensively discussed in Love (1944), is expanded to a form

in which it is easy to separate the Green's function and moment tensor.

One of the merits of wave integral theory is solutions can be represented in a suitable canonical form. That is, in a form in which the parameters of interest are taken into account at discrete points, so that the parameter input at the end of the computation chain can be varied without having to repeat the previous computations. Hence, it is desirable to factor the directivity dependence terms, or equivalently the radiation pattern, from the source representation and calculate only some fundamental source type solutions. These solutions are combined with the directional in the last stage to form the final soludependences Equation (II-2-15) is such an example where three fundamental types of solutions, i.e., SS, DS, and DD, are prepared before any orientation of double couple source is considered.

Since it is not easy to isolate directivity dependences from source expressions in the forms of the moment tensor, we choose the other approach from Haskell (1963). Using a direct and clear style, Haskell (1963, 1964) listed five source types, i.e., single force, single dipole, single couple, double couple, and compression source, in terms of a discontinuity in his vector $\mathbf{K} = \begin{bmatrix} A' + A'' & A' - A'' & B' - B'' & B' + B'' \end{bmatrix}^T$. A matrix \mathbf{F} of Wang

(1980) has been used and Herrmann to transform vector to the form $[A'', B'', A', B']^T$ and in Haskell's K section 2.2 we further used a matrix E to convert this form to a vector representing the discontinuity in the motion-stress vector (equation II-2-2). Using Σ^{H} represent the terms from Haskell (1964), represent a discontinuity in potential-constant vector, and S to represent a discontinuity in motion-stress vector, there exists the following relations:

$$\mathbf{S} = \mathbf{E} \, \Sigma$$
$$\Sigma = \mathbf{F} \, \Sigma^H .$$

where

$$\mathbf{F} = \frac{1}{2} \begin{bmatrix} -1 & -1 & 0 & 0 \\ 0 & 0 & -1 & 1 \\ -1 & -1 & 0 & 0 \\ 0 & 0 & 1 & 1 \end{bmatrix}.$$

These relations can be used to relate the different source expressions listed below.

Choose the coordinate system (x,y,z) as (1) x = north (2) y = east (3) z = downward into the earth, and f for the force acting at the source and f the normal to the fault plane, as defined in Haskell (1963). For an order one source, the only form is the single force. The solution can be expressed by

$$\mathbf{u} = f_3 S^0 Y_0 + (f_1 \cos \vartheta + f_2 \sin \vartheta) S^1 Y_1$$

$$4\pi\omega^2 \mathbf{S}^0 = [0, 0, -2k, 0]^T$$
$$4\pi\omega^2 \mathbf{S}^1 = [0, 0, 0, -2k]^T$$

where $f = [f_1, f_2, f_3]$ and ϑ is the receiver azimuth measured from the north direction clockwise. Y_n represents $U_z J_n$ for z component, and $U_{r,\vartheta} J_{n-1}$ for r, ϑ components as in equation (II-2-14).

For sources of order two, we find five fundamental forms are necessary, by examining the Haskell's (1964) source functions. They are

$$4\pi\omega^{2} S^{0'} = \begin{bmatrix} 0, & 0, & 0, & k^{2} \end{bmatrix}^{T}$$

$$4\pi\omega^{2} S^{0} = \begin{bmatrix} 0, & 2kk_{\alpha}^{2}/\rho, & 0, & 4k^{2}\beta^{2}/\alpha^{2} \end{bmatrix}^{T}$$

$$4\pi\omega^{2} S^{1'} = \begin{bmatrix} 0, & 0, & 2k^{2}, & 0 \end{bmatrix}^{T}$$

$$4\pi\omega^{2} S^{1} = \begin{bmatrix} -2kk_{\beta}^{2}/\rho, & 0, & 0, & 0 \end{bmatrix}^{T}$$

$$4\pi\omega^{2} S^{2} = \begin{bmatrix} 0, & 0, & 0, & k^{2} \end{bmatrix}^{T}.$$

The solutions for different source types are obtained after combining part or all of these fundamental forms. Such as,

dipole without moment:

$$\begin{aligned} \mathbf{u} &= \left[\; \left(1 - 3 f_{\,3}^{\,2} \right) \; \mathbf{S}^{0'} + f_{\,3}^{\,2} \; \mathbf{S}^{0} \; \right] \; \mathbf{Y}_{0} \\ &+ \left[\; f_{\,1} f_{\,3} \cos \vartheta + f_{\,2} f_{\,3} \sin \vartheta \; \right] \; \mathbf{S}^{1} \; \mathbf{Y}_{1} \\ &+ \left[\; \left(f_{\,2}^{\,2} - f_{\,1}^{\,2} \right) \cos 2\vartheta - 2 f_{\,1} f_{\,2} \sin 2\vartheta \; \right] \; \mathbf{S}^{2} \; \mathbf{Y}_{2} \end{aligned}$$

single couple:

$$\mathbf{u} = [-3f_3n_3 \mathbf{S}^{0'} + f_3n_3 \mathbf{S}^{0}] \mathbf{Y}_0$$

$$+ [((f_1n_3 - f_3n_1)\cos\vartheta + (f_2n_3 - f_3n_2)\sin\vartheta) \mathbf{S}^{1'}$$

$$+ (f_1n_3\cos\vartheta + f_2n_3\sin\vartheta) \mathbf{S}^{1}] \mathbf{Y}_1$$

+ [$(f_2n_2-f_1n_1)\cos 2\vartheta - (f_1n_2+f_2n_1)\sin 2\vartheta$] S^2 Y_2 double couple without moment:

$$\mathbf{u} = [-6f_{3}n_{3} S^{0'} + 2f_{3}n_{3} S^{0}] Y_{0}$$

$$+ [(f_{1}n_{3} + f_{3}n_{1}) \cos \vartheta + (f_{2}n_{3} + f_{3}n_{2}) \sin \vartheta] S^{1} Y_{1}$$

$$+ [2(f_{2}n_{2} - f_{1}n_{1}) \cos 2\vartheta - 2(f_{1}n_{2} + f_{2}n_{1}) \sin 2\vartheta] S^{2} Y_{2}$$

center of compression:

$$\mathbf{u} = \mathbf{S}^0 \mathbf{Y}_0 .$$

Note that for far-field solutions, S^0 Y_0 and S^2 Y_2 differ only by a constant phase term. Hence, for far-field double-couple dislocation sources, only three independent solutions are required, and the explosion solution is just one of them.

The fundamental types as expressed in terms of $\,\Sigma\,$ have the forms:

$$4\pi\omega^{2} \Sigma^{0'} = \begin{bmatrix} k^{3}/2\nu_{\alpha}, & -k/2, & -k^{3}/2\nu_{\alpha}, & -k/2 \end{bmatrix}^{T}$$

$$4\pi\omega^{2} \Sigma^{0} = \begin{bmatrix} kk_{\alpha}^{2}/\nu_{\alpha}, & 0, & -kk_{\alpha}^{2}/\nu_{\alpha}, & 0 \end{bmatrix}^{T}$$

$$4\pi\omega^{2} \Sigma^{1'} = \begin{bmatrix} -k^{2}, & k^{2}/\nu_{\beta}, & -k^{2}, & -k^{2}/\nu_{\beta} \end{bmatrix}^{T}$$

$$4\pi\omega^{2} \Sigma^{1} = \begin{bmatrix} 2k^{2}, & -(\nu_{\beta} + k^{2}/\nu_{\beta}), & 2k^{2}, & (\nu_{\beta} + k^{2}/\nu_{\beta}) \end{bmatrix}^{T}$$

$$4\pi\omega^{2} \Sigma^{2} = \begin{bmatrix} k^{3}/2\nu_{\alpha}, & -k/2, & -k^{3}/2\nu_{\alpha}, & -k/2 \end{bmatrix}^{T}$$

for second order sources, and

$$4\pi\omega^2 \Sigma^{0'} = \begin{bmatrix} k, -k/\nu_{\beta}, & k, k/\nu_{\beta} \end{bmatrix}^T$$

$$4\pi\omega^2 \Sigma^0 = \begin{bmatrix} -k^2/\nu_{\alpha}, & 1, k^2/\nu_{\alpha}, & 1 \end{bmatrix}^T$$

for a single-force type of source.

Similar relations can be derived for SH components. The results are

$$\mathbf{u} = \begin{bmatrix} 2(f_1 \sin \vartheta - f_2 \cos \vartheta) \end{bmatrix} \mathbf{S}^1 \mathbf{Y}_1$$
$$4\pi\omega^2 \mathbf{S}^1 = \begin{bmatrix} 0, k\omega^2 \end{bmatrix}^T$$

for a single-force source. For second order sources, the following are defined:

dipole without moment:

$$\mathbf{u} = [(f_2 f_3 \cos \vartheta - f_1 f_3 \sin \vartheta] S^1 Y_1$$

$$+ [(f_1^2 - f_2^2) \sin 2\vartheta - 2f_1 f_2 \cos 2\vartheta] S^2 Y_2$$

single couple:

$$\mathbf{u} = [f_1 n_2 - f_2 n_1] \mathbf{S}^0 \mathbf{Y}_0$$

$$+ [f_2 n_3 \cos \vartheta - f_1 n_3 \sin \vartheta] \mathbf{S}^1 \mathbf{Y}_1$$

$$+ [(f_1 n_1 - f_2 n_2) \sin 2\vartheta - (f_1 n_2 + f_2 n_1) \cos 2\vartheta] \mathbf{S}^2 \mathbf{Y}_2$$

double-couple without moment:

$$\mathbf{u} = [(f_{2}n_{3} + f_{3}n_{2})\cos\vartheta - (f_{1}n_{3} + f_{3}n_{1})\sin\vartheta] \mathbf{S}^{1} \mathbf{Y}_{1}$$

$$+ 2[(f_{1}n_{1} - f_{2}n_{2})\sin 2\vartheta - (f_{1}n_{2} + f_{2}n_{1})\cos 2\vartheta] \mathbf{S}^{2} \mathbf{Y}_{2}$$

where

$$4\pi\omega^{2} S^{0} = \begin{bmatrix} 0, k^{2}\omega^{2} \end{bmatrix}^{T}$$

$$4\pi\omega^{2} S^{1} = \begin{bmatrix} 2kk_{\beta}^{2}/\rho, & 0 \end{bmatrix}^{T}$$

$$4\pi\omega^{2} S^{2} = \begin{bmatrix} 0, k^{2}\omega^{2} \end{bmatrix}^{T}$$

5.2 Instrument Response Parameters by Least-Square Inversion

Mitchell and Landisman (1969) described a method for determining the instrumental constants of an

electromagnetic seismograph from a digitized calibration pulse. They applied a least-square technique to fit the pulse by varying four instrumental constants: period of seismometer (Ts) and galvanometer (Tg), and damping of seismometer (hs) and galvanometer (hq). A Fourier transformation was used to calculate the synthetic pulse from a known relationship between the constants in the frequency domain before comparison to the observed calibration pulse. Jarosch and Curtis simplified this method by deriving explicit equations for the pulse in the time domain, and included the scale factor as a free parameter to be determined. Mitronovas (1976) added another parameter, the original time, in the least-square fitting procedures to improve the error involved in specifying the onset time in pulse digitization.

This section will develop a similar method, but now, fitting the response parameters (to be defined in equation (V-2-1)) rather than the instrument constants Ts, hs, etc. It is important to realize that these parameters are the ones which determine the effects of the seismograph on the seismogram, not the instrument constants, although, theoretically both are mutually related. The seismograph is considered to be a 'black box'. Only linear response of the system is assumed; nothing about the instrument itself need be known. Such a method makes it possible to fit the calibration

pulse of any instrument in the time domain, especially those for which suitable formula to relate the instrument constants and the system response do not exist. There are several factors influencing the results of the time domain inversion method, such as the relation of the trial values to the actual, precision of the reference pulse, and errors in digitization. If these factors are highly controlled, the new method will be of practical use especially for the short-period instruments.

Most instruments in the current seismic detecting stations usually can be considered as a linear and stable system. A well-known response function approximating this system is: (Luh, 1977)

$$bF(s) = \frac{b \ s^m}{\sum_{j=1}^{n+1} a_j s^{j-1}},$$
 (V-2-1)

where s is the Laplace transform variable, b the scale factor, and the a's are assigned as the instrument response parameters. b and a's are all real positive coefficients with $a_{n+1}=1$, $a_1 \neq 0$. These parameters represent the system, and are those to be determined by the least-square inversion. m, n are chosen as integers representing the slopes near both ends of the amplitude response bands. The function to be fitted is given by

$$z(t) = bf(t,a_n,a_{n-1},\cdots,a_1),$$

where $f\left(t
ight)$ is the time domain counterpart of F(s) . Perturbing this function, the error will be

$$e_i = bf_i + b\sum_{j=1}^n \frac{\partial f_i}{\partial a_j} \delta a_j + f_i \delta b - z_i ,$$

where i indicates the data point on the time axis, and $f_i = f(t_i)$. For a least-squares fit, we want the sum of the squares of the errors or residuals (SSR) to be a minimum, i.e., $E = \sum_i e_i^2$ to be minimum and, hence,

$$\frac{\partial E}{\partial (\delta b)} = \frac{\partial E}{\partial (\delta a_j)} = 0 j = 1, \dots, n.$$

Writing

$$\varepsilon_i = z_i/b - f_i$$

the normal equations are

$$\begin{bmatrix} \sum_{i} (\frac{\partial f_{i}}{\partial a_{1}})^{2} & \cdots & \sum_{i} \frac{\partial f_{i}}{\partial a_{1}} \frac{\partial f_{i}}{\partial a_{n}} & \sum_{i} \frac{\partial f_{i}}{\partial a_{1}} f_{i} \\ \vdots & \vdots & \vdots \\ \sum_{i} \sum_{i} \frac{\partial f_{i}}{\partial a_{n}} \frac{\partial f_{i}}{\partial a_{1}} & \cdots & \sum_{i} (\frac{\partial f_{i}}{\partial a_{n}})^{2} & \sum_{i} \frac{\partial f_{i}}{\partial a_{n}} f_{i} \\ \sum_{i} \frac{\partial f_{i}}{\partial a_{1}} f_{i} & \cdots & \sum_{i} f_{i} \frac{\partial f_{i}}{\partial a_{n}} & \sum_{i} f_{i}^{2} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \sum_{i} \varepsilon_{i} \frac{\partial f_{i}}{\partial a_{1}} \\ \vdots \\ \sum_{i} \varepsilon_{i} \frac{\partial f_{i}}{\partial a_{n}} \\ \sum_{i} \varepsilon_{i} f_{i} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \sum_{i} \varepsilon_{i} \frac{\partial f_{i}}{\partial a_{n}} \\ \sum_{i} \varepsilon_{i} f_{i} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta b_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n} \end{bmatrix} \begin{bmatrix} \delta a_{1} \\ \vdots \\ \delta a_{n}$$

Following Jarosch and Curtis (1973), we take ($\alpha+\delta\alpha$) as the new value α' , but use

$$b' = b \exp(\delta b / b) \tag{V-2-3}$$

for the new scale factor, to force its fast convergence to exact value. However in practice, because of the linear dependence of b on the a's in the low frequency range (note that b and a's are terms of the numerator and denominator, respectively, in the response function V-2-1), the inclusion of b gives a constraint between b and a's, which sometimes causes the divergence of this method. Hence, an independent determination of b by other methods such as amplitude response analysis is recommended. The least-squares inversion is used to determine the a's only. In such a case the required modification of the above matrix equation is just to remove the n+1'th row and column.

Using the Laplace transformation pairs:

$$\frac{1}{s+a} \rightarrow e^{-at} \qquad \frac{1}{(s+a)^2} \rightarrow t e^{-at} , \qquad (V-2-4)$$

it is easy to find f(t) by factoring the response function F(s) and taking the inverse Laplace transform:

$$F(s) = \frac{s^{m}}{s^{n} + a_{n} s^{n-1} + \dots + a_{2} s + a_{1}}$$

$$= \frac{c_{1}}{s - d_{1}} + \frac{c_{2}}{s - d_{2}} + \dots + \frac{c_{n}}{s - d_{n}}$$

$$\Rightarrow f(t) = c_{1} e^{d_{1} t} + c_{2} e^{d_{2} t} + \dots + c_{n} e^{d_{n} t} = \sum_{i=1}^{n} c_{i} e^{d_{i} t}$$
where

 $C_i = \frac{d_i^m}{\prod\limits_{i=1}^n (d_i - d_k)}$

The roots d_i and coefficients c_i might be complex numbers. It is noted that in this expansion no double roots are permitted. These double roots might happen when parts of the instrument are set at critical damping. However, we find that a small coupling $\sigma^2 > 0.0001$ will push the imaginary part of the roots to be nonzero. Since the roots are always in conjugate pairs, the values of the c's are determined by $Re(d_i)$ and $|d_i|$. Hence, a small disturbance in the imaginary part, which can be added arbitrarily or with the inclusion of coupling factors, will not affect the result much, but will prevent the problem of a double root.

The time functions, $\frac{\partial f(t)}{\partial a_j}$ are more difficult to find. To avoid using numerical differencing, an analytic form can be derived as follows:

$$\frac{\partial F(s)}{\partial a_{j}} = \frac{(-1) s^{j-1+m}}{(s^{n} + a_{n} s^{n-1} + \dots + a_{2} s + a_{1})^{2}}
= \frac{(-1) s^{j-1+m}}{(s-d_{1})^{2} (s-d_{2})^{2} \dots (s-d_{n})^{2}}
= -\left[\frac{p_{1j}}{(s-d_{1})^{2}} + \frac{q_{1j}}{s-d_{1}} + \frac{p_{2j}}{(s-d_{2})^{2}} + \frac{q_{2j}}{s-d_{2}} + \dots + \frac{p_{nj}}{(s-d_{n})^{2}} + \frac{q_{nj}}{s-d_{n}}\right]$$

where

$$p_{ij} = \frac{d_i^{j-1+m}}{\prod\limits_{k=1}^{n} (d_i - d_k)^2} = d_i^{j-1-m} c_i^2 \qquad i,j = 1,...,n.$$

The expression for q_{ij} needs more derivations. Set

$$H(s) = \prod_{k=1}^{n} (s - d_k)^2$$

 q_{ij} can be evaluated by

$$q_{ij} = \left[\frac{s^{j-1+m}}{(s-d_i)H(s)} - \frac{p_{ij}}{(s-d_i)} \right]|_{s \to d_i}$$

$$= \frac{1}{(s-d_i)} \left[\frac{s^{j-1+m}}{H(s)} - \frac{d_i^{j-1+m}}{H(d_i)} \right]|_{s \to d_i}$$

$$= \frac{\partial}{\partial s} \left[\frac{s^{j-1+m}}{H(s)} \right]|_{s \to d_i}$$

$$= (j-1+m) \frac{s^{j-2+m}}{H(s)} - 2 \frac{s^{j-1+m}}{H(s)} \left(\sum_{k=1}^{n} \frac{1}{k \neq i} \frac{1}{s-d_k} \right)|_{s = d_i}$$

$$= p_{ij} \left[(j-1+m) d_i^{-1} - 2 \left(\sum_{k=1}^{n} \frac{1}{k \neq i} \frac{1}{d_i - d_k} \right) \right]$$

Hence from the Laplace transform (V-2-4), the time domain derivative $\frac{\partial f(t)}{\partial a_i}$ has the form:

$$\frac{\partial f(t)}{\partial a_j} = -\sum_{i=1}^n \left(p_{ij} t e^{d_i t} + q_{ij} e^{d_i t} \right)$$

where

$$p_{ij} = d_i^{j-1-m} c_i^2$$

$$q_{ij} = p_{ij} \left[(j-1+m) d_i^{-1} - 2 \sum_{k=1}^n \frac{1}{d_i - d_k} \right]$$

$$d_i^{-1} = 0$$

We find that the derivatives of f(t) have the close forms in the time domain. This is difficult to attain by varying the instrumental constants, as in common least-square calibration pulse inversion methods.

Since the differences in order of values of elements in the inversion matrix (equation V-2-2) are usually about ten to twenty, a scaling constant sometimes

is needed to prevent numerical overflow or underflow. It is made by simply changing the transformation factor s to $\rho s'$ and the parameters found will be scale adjusted $b' = \rho^{m-n}b$, $a_j' = \rho^{j-n}a_j$. The time axis in this case is reduced by $1/\rho$ because of the corresponding scale change in the frequency.

To test the new method, an example from Mitronovas (1976) is chosen. The results are shown in Table 2. We first compute a synthetic pulse by the equation (e) Jarosch and Curtis (1973) for the instrument with seismometer overdamped and galvanometer underdamped, then normalize to the maximum values indicated, and round off to nearest integer values to simulate differences in the digitization accuracy. A set of a values of the instrument parameters are then obtained using the new inversion technique. Table 2 also lists the values using the method of Mitchell and Landisman (1969) as a comparison. There does not seem to be any significant difference between these methods, although the new method seems a little better in accuracy for such a synthetic case. This test confirms the of the new method.

Figures 36 and 37 show the application to real data for the LPZ WWSSN instrument at FVM station. Figure 36 displays the digitized calibration pulse and its amplitude spectrum. Using this pulse as an input, we

TABLE 2

Comparison of Parameters for Synthetic Pulses

a ₁ (10 ¹) a ₂ 1066312 .2915167 1066272 .2915026
1060013
1035897
1031482
.1034220 .1034212

of Mitchell and Landisman (1969), and the lower values are solu-* In each case the upper values are solutions based on the method tions based on the present method.

 $^{ ilde{ heta}}$ Sum of squares of residuals between the reference and the calculated pulse.

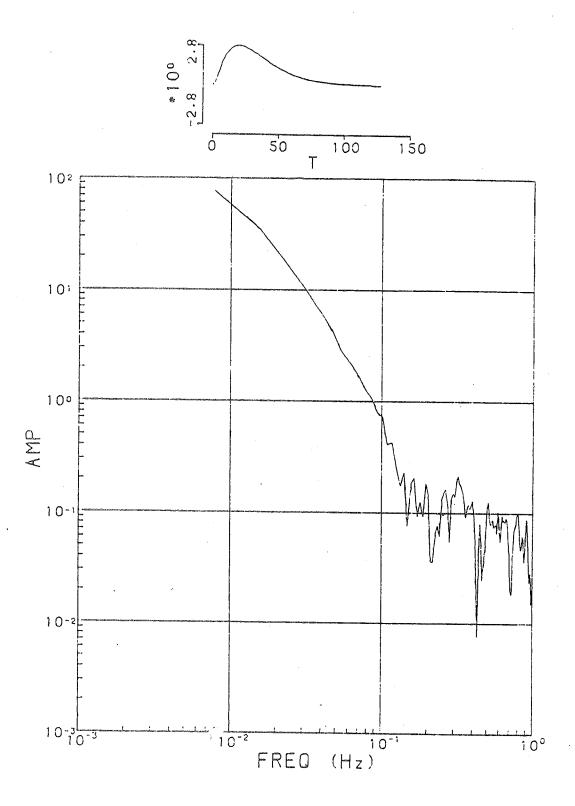


Figure 36. The calibration pulse and its amplitude spectrum for the WWSSN LPZ instrument at the station FVM.

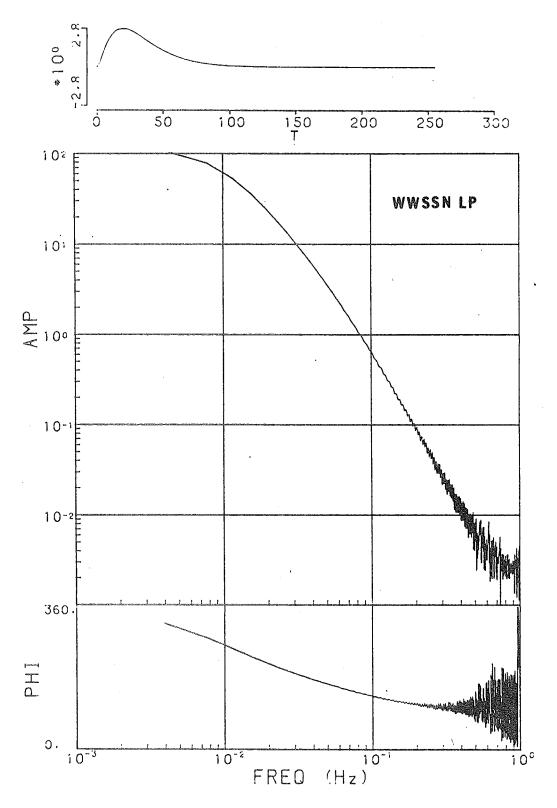


Figure 37. The impulse-response pulse and its amplitude and phase spectra of a simulated instrument obtained by applying the inversion technique to the calibration pulse of Figure 36.

determine the instrument response parameters by the present inversion method. Figure 37 shows the impulse response pulse of this simulating instrument, and its amplitude and phase responses. It is found that, except for high frequency noise, the new method can adequately model an instrument even if we lack knowledge of the instrument constants. Furthermore, the α parameters obtained can be easily incorporated in the program to describe the instrument effect. In the next section we will discuss such an application in the time domain.

5.3 Simulating the Instrument by IIR

With the response parameters determined, instrument response (V-2-1) upon the input the data by direct multiplication in the frequency domain, and then take an inverse FFT to synthesize the time series. However, if the generated seismogram is already history, a time domain operator representing the instrument would be more convenient to In this section an infinite impulse response filter (IIR) will be designed to make the instrument response superposition in the time domain by a recursive operation. The basic theory will be the classical Z transformation long been used to treat digital signals which has (Robinson and Treitel, 1980).

First let us consider one of the filters in

equation (V-2-5):

$$F(s) = \frac{1}{s-d} .$$

The relationships among Fourier, Laplace, and Z transformations for this filter are easy to find,

$$e^{dt}U(t) \rightarrow \frac{1}{i\omega - d}$$

$$\rightarrow \frac{1}{s - d}$$

$$\rightarrow \frac{1}{1 - e^{d\Delta t}Z^{-1}},$$
(V-3-1)

where U(t) is the Heaviside step function. Because of the causality of the system, the poles in equation (V-3-1) should be located in the left half of the complex s-plane (Papoulis, 1962), or equivalently, inside the unit circle of the complex Z-plane. If we use the following definition for the Z-transform:

$$F(Z) = \sum_{n=0}^{\infty} f(n) Z^{-n}$$

the two-term filter will be stable for such a system, since if expanded in an infinite polynomial of Z^{-1} , the coefficients are all bounded. This stable property is necessary for any IIR to simulate the instrument response. Before we design such a filter, we must be certain to check that all of the poles of the Laplace transformation have a negative real part.

From equation (V-2-5), the instrument response can

be written as

$$F(s) = \sum_{i=1}^{n} \frac{C_i}{s - d_i} . (V-3-2)$$

The corresponding Z-transform will be

$$F(Z) = \sum_{i=1}^{n} \frac{C_i}{1 - e^{d_i \Delta t} Z^{-1}} . \tag{V-3-3}$$

This formula represents n number of two-term filters operating in parallel. We are not going to sum them up as a single filter (Seidl, 1980)

$$F(Z) = \frac{a_0 + a_1 Z^{-1} + \cdots + a_{n-1} Z^{-(n-1)}}{1 + b_1 Z^{-1} + \cdots + b_n Z^{-n}},$$

because the numerical error makes the determination of a's and b's unsatisfactory. Besides, this form is not easy to apply (Kulhanek, 1979).

It is known that the multiplication of Z^{-1} means a shift of the time sequence by one sampling period. Applying the filter (V-3-3) to the input data, x_k , we have the output y_k ;

$$y_{k} = F(Z) x_{k}$$

$$= \left(\sum_{i=1}^{n} \frac{C_{i}}{1 - e^{d_{i}\Delta t} Z^{-1}} \right) x_{k}$$

$$= \sum_{i=1}^{n} \left(\frac{C_{i}}{1 - e^{d_{i}\Delta t} Z^{-1}} x_{k} \right)$$

$$= \sum_{i=1}^{n} y_{k}^{(i)}$$

where

$$y_k^{(i)} = c_i x_k + e^{d_i \Delta t} y_{k-1}^{(i)}$$
.

k is the sampling point of the sequence, and i represents a different two-term recursive filter. This is the recursive formula describing the time domain operation.

Figure 38 gives the impulse response pulse and its spectrum for an LRSM 6284-13 seismograph. The impulse response waveform is obtained by passing an signal through a Z-transform IIR filter representing the instrument. Different amplitude response curves come from different time sampling intervals, which are 1.0, 0.5, and 0.0625 second for curves from top to bot-A theoretical response curve, which falls in the position of the curve with $\Delta t = 0.0625$, is also shown for Some restriction should be considered comparison. Fourier before using recursive filters. For the transformation, the sampling along the frequency axis, Δf , gives the periodicity in the time domain, causes the aliasing effect of the time series (Brigham, 1974). Equivalently, the time sampling Δt of the Ztransform causes the aliasing effect, but now in the frequency domain. Figure 38 shows the effect of aliasing at the ends of the frequency response curves for different sampling rates. Naturally smaller Δt values The reason for such an effect give better results. comes from $Z = e^{i\omega \Delta t}$ which is a harmonic function with period of 2π . To keep $\omega \Delta t = 2\pi$ smaller Δt gives higher Nyquist frequency, i.e., shifts the aliasing to

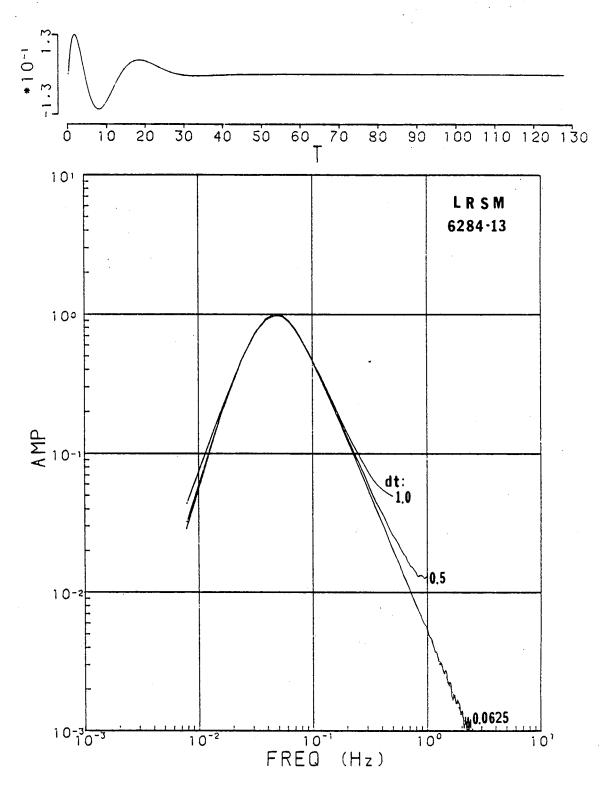


Figure 38. Simulation of an LRSM 6284-13 seismograph using the Z-transform method. Three response curves correspond to different sampling rates dt of 1.0, 0.5 and 0.0625 sec.

high frequencies.

Another way to alleviate the aliasing problem is to use the bilinear Z-transform. By substituting

$$s = \frac{2}{\Delta t} \frac{1 - Z^{-1}}{1 + Z^{-1}}$$

in equation (V-3-2), it is easy to find the recursive relation of a two-term filter for the bilinear Z-transform:

$$y_k^{(i)} = \frac{c_i}{2/\Delta t - d_i} (x_k + x_{k-1}) + \frac{2/\Delta t + d_i}{2/\Delta t - d_i} y_{k-1}^{(i)}$$

The bilinear Z-transform is a low frequency approximation (Oppenheim and Schafer, 1975, p.208). The distortion of the frequency axis at high frequencies offsets the true response. As shown in Figure 39, which describes the response of a short period WWSSN instrument, the bilinear Z-transform simulates the response well at low frequencies, but is poor for high frequencies. To obtain good results, a small Δt is again required. On the other hand, the filter response is found to be zero at the Nyquist frequency, which avoids any Gibb's phenomina in the time domain. A special frequency warping is used in practice to design the digital filter from its analytic form.

Using the time domain operation, we designed several IIR filters to model some currently used

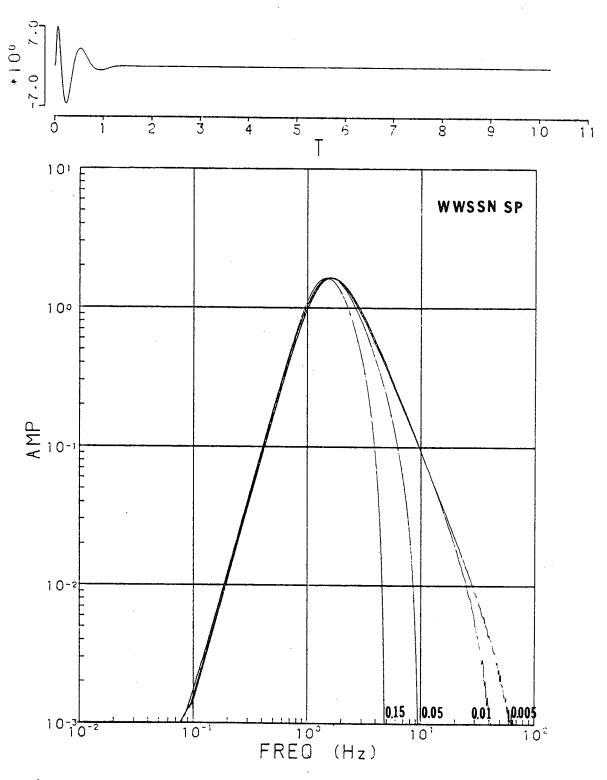
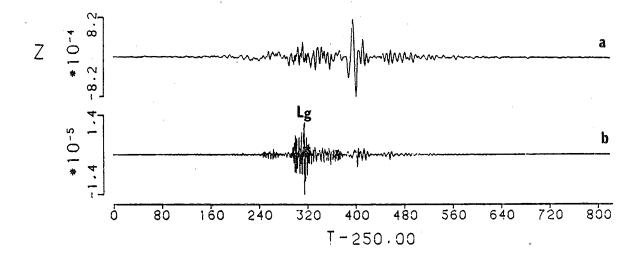


Figure 39. Simulation of a WWSSN SP instrument using the bilinear Z-transform method. Different response curves correspond to different sampling rates of 0.15, 0.05, 0.01 and 0.005 second.

instruments. Figure 40 illustrates one of the results. The seismograms plotted correspond to the time histories before and after the addition of a WWSSN SP instrument for two cases. The instrument response curve is the one given in Figure 39. It can be seen that the waveforms are totally altered. Low frequency signals are filtered out, and the remaining waveforms represent more likely the high frequency, higher mode Lg phases.



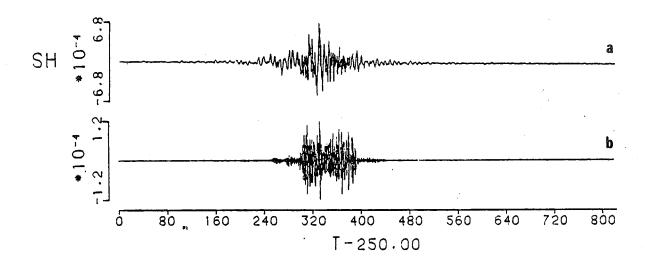


Figure 40. Seismograms showing the effect of the instrument. The two displays are for the vertical and tangential components, respectively. The CUS model and a distance of 2000 km are used. (a) is the ground displacement; and (b) is the seismogram after passing through a short-period instrument. The instrument response is that shown in Figure 39.

CHAPTER VI

COMPARISONS

To employ the synthetic seismogram method in actual applications, the validity of theory and numerical procedures should be checked. In this chapter, several comparisons of the wave integral method with totally different methods are presented. These examples illustrate the flexibility and reliability of the present method, and lend confidence to the new system.

Comparison to Generalized Ray Theory

First, the free surface displacements due to a double-couple source in a uniform half-space generated by the integral method are checked against the complete solutions from the Cagniard-de Hoop method (Johnson, 1974). Figures 41 and 42 show the computed radial components of ground velocity time histories at distances of 10, 25, 50, and 75 km for the vertical strike-slip and vertical dip-slip sources, respectively. Sections a in these figures come from a Cagniard-de Hoop ray summation and sections b from the integral method. The velocity structure has P velocity = 6.15 km/sec, S velocity = 3.55 km/sec, density = 2.8 gm/cm³. The depth of the point dislocation is 10 km. A seismic moment of

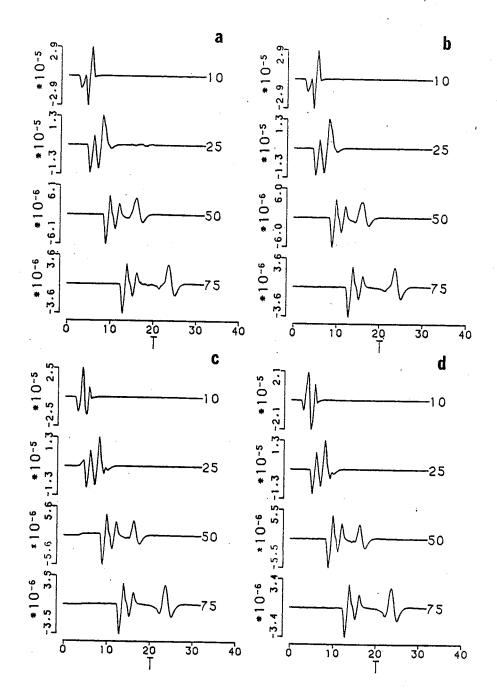


Figure 41. Comparison of Cagniard-de Hoop and wave theory solutions for a vertical strike-slip source at a depth of 10 km in a half-space with parameters of the first layer of SCM model. (a) Cagniard-de Hoop solution. (b) The complete wave theory solution. (c) The wave theory solution containing only the near-field and far-field P-SV terms. (d) The wave theory solution including only the far-field P-SV term.

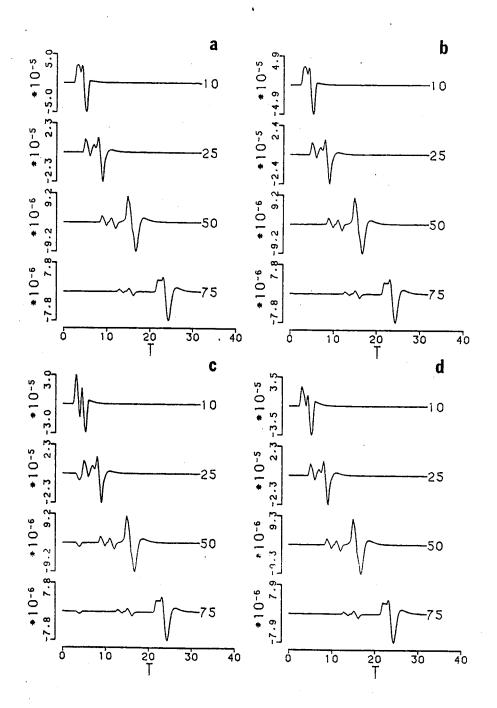


Figure 42. Results for the same model of Figure 41, but for the radial component due to a vertical dip-slip source.

1.0 E +20 dyne-cm and a parabolic source time function with τ = 0.5 sec are applied.

The agreement between the records of these methods is excellent. The comparisons are of nearperfect precision for both fault slip prescriptions. This convinces us that the method of wave integrals is highly reliable. Figures 41 and 42 also show effect of near-field terms (equation II-2-14) on the Sections c in these complete solutions. figures include the contributions of far-field and near-field P-SV terms (first two terms in equation II-2-14), where a noncausal, nonpropagating arrival exists as indicated by Herrmann (1978a), especially for the vertical dip-The records in sections d correspond to slip source. the solution using only the far-field term. From these figures, we can see that an apparent Rayleigh phase (fundamental mode only for half-space model) emerges from the P phase group as the distance increases. Since a clear distinction between the S arrival and surface wave phases is not possible, the separation between these two main phases on seismograms is not obvious.

Comparison to Finite-Element Method

The second example (Figure 43, 44) shows the accuracy of the present solution for a horizontally layered media. The compared solutions are obtained using the finite element method, which are adopted directly from

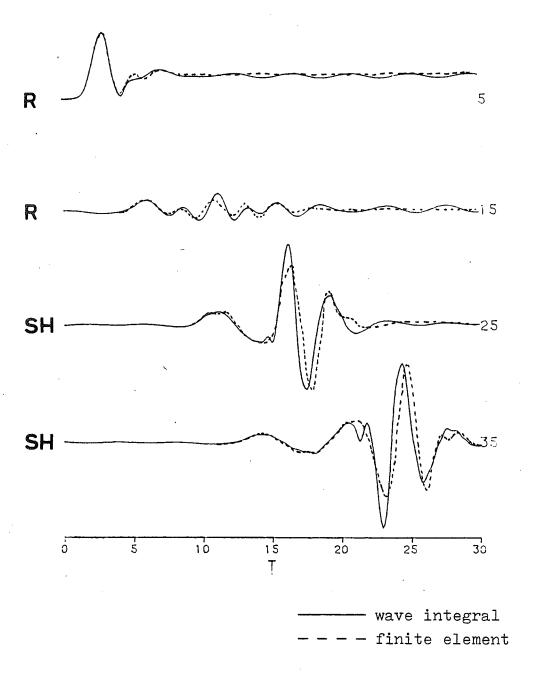


Figure 43. Comparison of wave integral solution with the finite element solution for the radial and tangential components due to a vertical strikeslip dislocation buried at a depth of 1 km in the two layers overlying half-space model of Table 3.

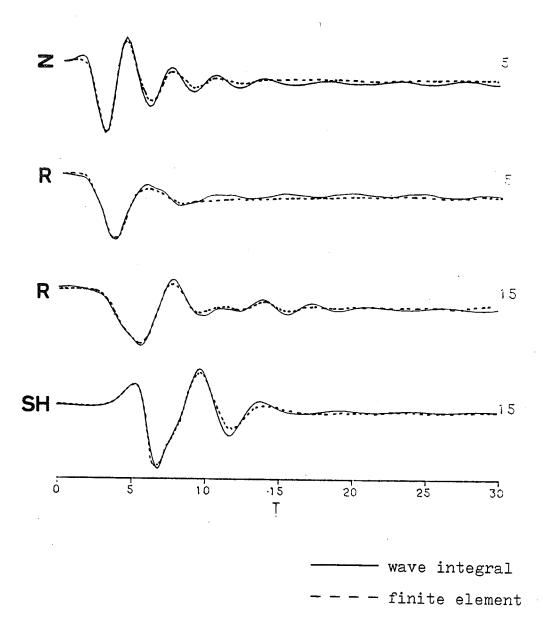


Figure 44. Results using the source and model of Figure 43, but for a source buried at the depth of 5 km.

Apsel (1979). The model consists of two layers overlying a half-space, as illustrated in Table 3. A vertical strike-slip dislocation source is considered and the source time function is a ramp of one second duration. The receivers with an azimuth of 22.5 degrees from the strike of the fault are located at the distances indicated at the ends of each traces. A low-pass filter with corner frequency Ø.5 Hz is applied to make the seismogram more visible. The results of the present method are computed up to 4 Hz before passing through the filter.

The comparsion results are shown in Figure 43 for a source at 1 km depth, and in Figure 44 for a source at 5 km. The components displayed are also indicated. The agreement in these figures is remarkable, especially in light of the vast differences between the two techniques. The slight deviation comes from the numerical errors of both methods and from slightly different low-pass filters used. However, the overall pattern still exhibits great consistency, especially for deeper sources, which is sufficient to confirm the success of the wave integral method.

Comparison to Modal Summation Method

In the previous examples we showed the comparisons with known complete solutions. The solutions include near-field as well as far-field terms, and the

TABLE 3

Two Layers Overlying Half-space Model

Thickness (km)	P vel (km/sec)	S vel (km/sec)	Density (g/cm³)
2	3.0	1.73	1.67
2	5.0	2.887	2.89
	6.0	3.46	3.46

integration is taken for different integral parts to create all kinds of signals. In the next example, several synthetic seismograms from Swanger and Boore (1978) and Heaton and Helmberger (1976,1977) are chosen to check against our pole contribution solutions, namely surface wave, as presented in chapter III. However, a solution from the locked mode approximation method is also included in an attempt to fit body wave signals.

Using the surface-wave modal superposition method Harkrider (1964,1970), Swanger and Boore (1978) simulated ground displacement records from earthquakes Imperial Valley of California. Since the sediments in this valley form a prominent wave guide, surface waves are thought to be dominant. Heaton and Helmberger (1976,1977) generated several synthetic seismograms using the generalized ray method for this area. They have used enough rays in their summation so that their synthetics can be considered to be a nearcomplete solution. The effort of these studies was directed to find a suitable source model by time domain waveform fitting. However, the synthetic waveforms created provide us with a check of our eigenfunction programs. Figures 45 and 46 are the results showing the comparison of these different methods.

Using the El Centro model listed in Table 4, we

TABLE 4
Earth Models

 			
Thickness (km)	P vel (km/sec)	S vel (km/sec)	Density (g/cm³)
	El Centro S	tructure	
2.9	_	1.50	1.5
_	_ '	3.30	2.5
	Imperial Valle	y Structure	
0.95	2.0	0.88	1.8
1.15	2.6	1.50	2.35
3.8	4.2	2.40	2.6
_	6.4	3.70	2.8

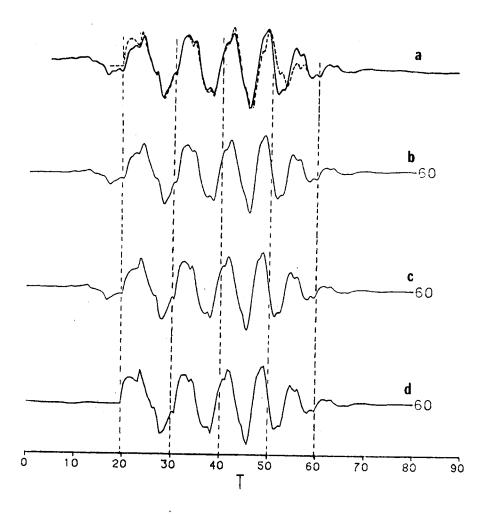


Figure 45. Comparison of Love wave synthetic ground displacements with Swanger and Boore's modal summation method (solid line in a) as well as Heaton and Helmberger's ray summation method (dashed line in a) of the 1968 Borrego Mountain earthquake. A vertical strike-slip source at 6 km depth and a symmetric triangular source time function of 1 second duration are used. The epicentral distance is 60 km and the azimuth is 8 degrees from the strike of the fault. model used is the El Centro structure listed in Table (b) is the result of eigenfunction programs but including only the first three modes which Swanger and Boore used. (c) is the result including all modes. (d) shows the synthetics from the locked mode approximation with a rigid layer at 200 km deep. Note that (d) successfully models the first arrival of ray theory solution.

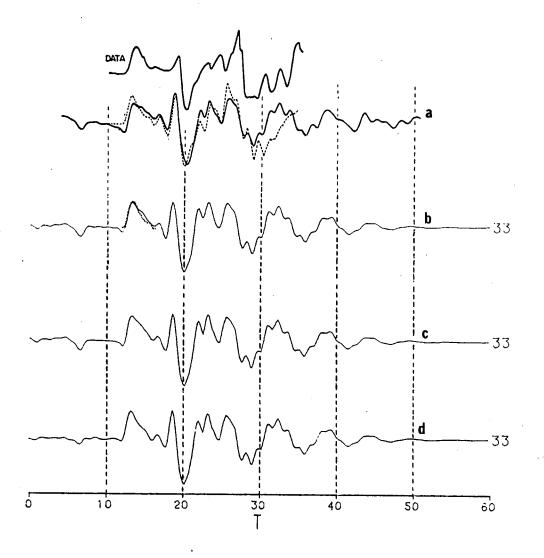


Figure 46. Same comparison as for Figure 45 but for the 1976 Brawley earthquake. The Imperial Valley structure (Table 4) of Heaton and Helmberger (1978) is used. The source is a vertical strike-slip point buried at 6.9 km, and the source time function is a 1.5 sec duration triangle. The top trace gives the real data (Heaton and Helmberger, 1978). Our solution (b), which is the summation of first five modes, shows a better fit of the first P arrival at 12 second than Swanger and Boore (1979). Again, the locked mode approximation (d) gives a good match to the ray theory solution in the front part of the record, but not in the rest.

generated surface wave seismograms after summing first three modes ((b) in Figure 45) or all the modes ((c) in Figure 45) by eigenfunction programs. The match of these seismograms to Swanger and Boore's result (solid trace in 45a) is obvious. The agreement superb and the phase coherence is nearly perfect. Furthermore, the seismogram from the locked approximation ((d) in Figure 45) fits the first P arrival of Heaton and Helmberger's solution (dashed in Figure 45a) quite well, although the later part of the trace seems not affected by the presence of a rigid cap layer.

Similar conclusions can be drawn for the comparison using the Imperial Valley model of Table 4. As shown in Figure 46, the result (b) of the present method seems to be a better fit comparison of the direct P arrival to ray theory solution than Swanger Boore obtained. The reason might be a longer time window we have used to avoid the time aliasing. locked mode approximation, again, shows a good match to the ray theory solution in the front part of the trace. These results serve to further validate the present method, despite the fact that various assumptions inherent in the other methods. In addition, the comparison of (b) and (c) in Figures 45 or 46 also reveals that the number of modes Swanger and Boore used to make up their seismograms was sufficient.

CHAPTER VII

SUMMARY AND CONCLUSIONS

A detailed and complete derivation of wave integral theory for studying wave propagation in plane multi-layered media has been presented. The three-dimensional wave propagation problem is formulated and solved in the wavenumber-frequency domain. The layer responses at any wavenumber and frequency are given in terms of Haskell's layer matrices and the corresponding compound forms. The formalism presented was shown to be stable and accurate for computations. A contour integration is taken over horizontal wavenumber so as to automatically include all kinds of waves. A fast Fourier transform then gives the final time history.

The technique is not new. However, after a close step-by-step development, several insights were revealed. The most important was the finding of symmetry properties of the layer matrix and its compound counterparts. These properties enable us to build up the relation of the present method with eigenfunction theory and most of other currently used wave integral methods.

Eigenfunction solutions as expressed in analytic forms have the advantage of computational accuracy and efficiency. The special cases of high frequencies and

complicated models can be handled without difficulty. New procedures were developed to find dispersion The eigenfunctions, energy integrals, and all other excited variables of surface waves could evaluated with sufficient precision. Furthermore, the theoretical development also provided a rigorous for classical eigenfunction theory. This in turn opened a new, and potentially powerful method. which can be used to extract more information about the wave propagation through the earth.

The contour integration over the real branch cut is a difficult part of simulating body wave-like signals. The variation of the integrand along the real-k axis reflects the influence of leaky modes. After an intensive study of this effect, we were able to design a numerical integration procedure for this integration path. We also found that body waves and surface waves cannot be separated from each other. They are interconnected.

To test the versatility of the eigenfunction program developed, the locked mode approximation was introduced to simulate the signals from the branch cut contribution. Except for the computational speed, the test was highly successful. This method is worthy of further investigation.

One significant value of the present study is its

ability to be extended to other methods. The reflection and reflectivity methods can be easily derived using the new system. As a result, the reflection and transmission properties of a layer boundary can be easily decomposed. With an arbitrary choice of reflection and/or transmission at interfaces, theoretical seismograms consisting of signals from a particular portion of the layered structure were obtained.

Source functions characterized by any first or second order point source were considered. A set of five fundamental forms was formed to represent all kinds of signals possibly excited in the earth. This expression, which isolates the direction-dependence in the final solution, is specially useful for focal mechanism studies.

A new way to find the instrument response, using least-squares inversion of calibration pulse waveforms, was developed. The method found the parameters which directly represent the instrument responses, even though the instrument-related constants are not given directly. These inversion procedures were performed in the time domain, and thus are easy to use. Finally, a digital z-transform technique, which was used to introduce the instrument effect in the time domain, was also discussed.

To verify the theory and numerical extension, a

set of three validation studies was provided. The results of comparison to other known complete solutions serve to lend confidence in our method.

Because the models treated are plane stratified media, the present method can describe the wave propagation in the earth up to an epicentral distance, say, of 3000 km. Beyond this distance, the earth flattening correction from Biswas and Knopoff (1970) or Chapman (1973) should be considered. Furthermore, the dispersion curves for the models with an apparent low velocity zone sometimes become extremely complicated. Special care is needed to handle these particular cases.

An overall view of this wave theory system shows its completeness and ease of adaptation. With the advent of future computers, the method will become more and more important. It is expected that the present development and associated computer programs will prove increasingly useful in various areas of theoretical seismology and earthquake engineering.

APPENDIX A

Layer Matrix

The elements of layer matrix α are

$$a_{11} = \gamma \cosh \nu_{\alpha} z - (\gamma - 1) \cosh \nu_{\beta} z$$

$$\alpha_{12} = k(\gamma - 1) \frac{\sinh \nu_{\alpha} z}{\nu_{\alpha}} - \frac{\gamma}{k} \nu_{\beta} \sinh \nu_{\beta} z$$

$$\alpha_{13} = \frac{k}{\rho} \cosh \nu_{\alpha} z - \frac{k}{\rho} \cosh \nu_{\beta} z$$

$$\alpha_{14} = -\frac{k^2}{\rho} \frac{\sinh \nu_{\alpha} z}{\nu_{\alpha}} + \frac{1}{\rho} \nu_{\beta} \sinh \nu_{\beta} z$$

$$a_{21} = -\frac{\gamma}{k} \nu_{\alpha} \sinh \nu_{\alpha} z + k (\gamma - 1) \frac{\sinh \nu_{\beta} z}{\nu_{\beta}}$$

$$\alpha_{22} = -(\gamma - 1) \cosh \nu_{\alpha} z + \gamma \cosh \nu_{\beta} z$$

$$\alpha_{23} = -\frac{\nu_{\alpha}}{\rho} \sinh \nu_{\alpha} z + \frac{k^2}{\rho} \frac{\sinh \nu_{\beta} z}{\nu_{\beta}}$$

$$a_{31} = -\rho \frac{\gamma(\gamma - 1)}{k} \cosh \nu_{\alpha} z + \rho \frac{\gamma(\gamma - 1)}{k} \cosh \nu_{\beta} z$$

$$\alpha_{32} = -\rho \ (\gamma - 1)^2 \ \frac{\sinh \nu_{\alpha} z}{\nu_{\alpha}} + \rho \ \frac{\gamma^2}{k^2} \ \nu_{\beta} \ \sinh \nu_{\beta} z$$

$$a_{41} = -\rho \frac{\gamma^2}{k^2} \nu_{\alpha} \sinh \nu_{\alpha} z + \rho (\gamma - 1)^2 \frac{\sinh \nu_{\beta} z}{\nu_{\beta}}$$

$$\alpha_{55} = \cosh \nu_{\beta} z$$
 $\alpha_{65} = \mu \nu_{\beta} \sinh \nu_{\beta} z$

$$\alpha_{56} = \frac{\sinh \nu_{\beta} z}{\mu \nu_{\beta}} .$$

The other elements are found by

$$\alpha_{ij} = \alpha_{5-j,5-i}$$

for $i, j = 1, \dots, 4$, and

$$a_{ij} = a_{11-j,11-i}$$

for i, j = 5, 6.

APPENDIX B

Compound Matrix

The compound matrix E^{-1} used in the main text is

$$E_{N}^{-1}|_{ij}^{12} = \frac{1}{4k \, \nu_{\alpha N} \, \nu_{\beta N}} \left[E^{-1}|_{12}^{12}, E^{-1}|_{13}^{12}, E^{-1}|_{14}^{12}, E^{-1}|_{23}^{12}, E^{-1}|_{24}^{12}, E^{-1}|_{34}^{12} \right]_{N}$$

where

$$E^{-1}|_{12}^{12} = \rho^2 \left[-\frac{\gamma^2}{k^2} \nu_{\alpha} \nu_{\beta} + (\gamma - 1)^2 \right]$$

$$E^{-1}|_{13}^{12} = -\rho \nu_{\alpha}$$

$$E^{-1}\big|_{14}^{12} = \rho \left[\frac{\gamma}{k} \nu_{\alpha} \nu_{\beta} - (\gamma - 1) k \right]$$

$$E^{-1}|_{23}^{12} = E^{-1}|_{14}^{12}$$

$$E^{-1}|_{24}^{12} = -\rho \nu_{\beta}$$

$$E^{-1}|_{34}^{12} = \nu_{\alpha}\nu_{\beta} - k^{2}$$
.

The compound matrix of layer matrix lpha is a 6x6 matrix whose components are

$$\alpha \mid_{12}^{12} = CPCQ + 1 - \alpha \mid_{14}^{14}$$

$$a \mid_{13}^{12} = (-CQX + k^2 CPY)/\rho$$

$$\alpha \mid_{14}^{12} = [(1 - CPCQ)(2\gamma - 1) + \frac{\gamma}{k^2}XZ + (\gamma - 1)k^2WY]k/\rho$$

$$a \mid_{23}^{12} = a \mid_{14}^{12}$$

$$a \mid_{24}^{12} = (k^2 CQW - CPZ)/\rho$$

$$\alpha \mid_{34}^{12} = [2(1 - CPCQ) k^2 + XZ + k^4 WY]/\rho^2$$

$$a \mid_{12}^{13} = \rho \left[-(\gamma - 1)^2 CQW + \frac{\gamma^2}{k^2} CPZ \right]$$

$$\alpha \mid_{13}^{13} = CPCQ$$

$$\alpha \mid_{14}^{13} = (\gamma - 1) k CQW - \frac{\gamma}{k} CPZ$$

$$a \mid_{23}^{13} = a \mid_{14}^{13}$$

$$\theta^{-\nu'}, \qquad \alpha \mid_{24}^{13} = WZ$$

$$a \mid_{12}^{14} = -\rho[(1 - CPCQ) \frac{\gamma}{k} (\gamma - 1)(2\gamma - 1) + \frac{\gamma^3}{k^3} XZ + (\gamma - 1)^3 k WY]$$

$$a \mid_{13}^{14} = k (\gamma - 1) CPY - \frac{\gamma}{k} CQX$$

$$a \mid_{14}^{14} = 1 + 2(1 - CPCQ) \gamma(\gamma - 1) + \frac{\gamma^2}{k^2} XZ + (\gamma - 1)^2 k^2 WY$$

$$a \mid_{23}^{14} = a \mid_{14}^{14} - 1$$

$$a \mid_{12}^{23} = a \mid_{12}^{14}$$

$$a \mid_{13}^{23} = a \mid_{13}^{14}$$

$$a \mid_{14}^{23} = a \mid_{23}^{14}$$

$$\alpha \mid_{12}^{24} = \rho \left[CQX \frac{\gamma^2}{k^2} - CPY (\gamma - 1)^2 \right]$$

$$\alpha \mid_{13}^{24} = XY$$

$$\alpha \mid_{12}^{34} = \rho^2 \left[2(1 - CPCQ) \frac{\gamma^2}{k^2} (\gamma - 1)^2 + \frac{\gamma^2}{k^4} XZ + (\gamma - 1)^4 WY \right].$$

The other components can be found by

$$a\mid_{kl}^{ij}=a\mid_{5-j,5-i}^{5-l,5-k}$$

where

$$\gamma = \frac{2k^2}{k_B^2}$$

 $CPCQ = \cosh \nu_{\alpha} z \cdot \cosh \nu_{\beta} z$

$$CPY = \cosh \nu_{\alpha} z \cdot Y$$

$$CPZ = \cosh \nu_{\alpha} z \cdot Z$$

$$CQW = \cosh \nu_{\beta} z \cdot W$$

$$CQX = \cosh \nu_{\beta} z \cdot X$$

$$XY = X \cdot Y$$

$$XZ = X \cdot Z$$

$$WY = W \cdot Y$$

$$WZ = W \cdot Z$$

with

$$W = \frac{\sinh \nu_{\alpha} z}{\nu_{\alpha}}$$

$$X = \nu_{\alpha} \sinh \nu_{\alpha} z$$

$$Y = \frac{\sinh \nu_{\beta} z}{\nu_{\beta}}$$

$$Z = \nu_{\beta} \sinh \nu_{\beta} z$$
.

APPENDIX C

Symmetry of Compound Matrix

Because of the particular symmetry existing in the layer matrix a , and those existing between a , E and their inverses respectively, we can reveal some interesting properties of a compound matrix obtained from these two kinds of matrices. In this appendix, these properties will be developed in somewhat rigorous mathematical terms for the compound matrices of order two which are used here. The equivalence of the third and fourth components of some of Haskell's compound matrices will consequently become apparent. the properties derived in this appendix were directly in the main text.

The special types of symmetry properties for the compound matrix can be summarized by two definitions:

(1) If two compound matrices A,B satisfy

$$A \mid_{kl}^{ij} = pB \mid_{ij}^{kl}$$
,

where p is a constant coefficient and ij \leftrightarrow ij:

then A and B are called 'skew-symmetric'.

(2) If two compound matrices A,B satisfy

$$A \mid_{kl}^{ij} = pB \mid_{5-j,5-i}^{5-l,5-k}$$

then A and B are called '5-complemental'.

When B=A and p=1, such symmetric properties are called 'self skew-symmetric' or 'self 5-complemental'. When B = A' and p= (-1)^{i+j+k+l} q, where q is a constant determined from A, such symmetric properties are called 'inverse skew-symmetric' or 'inverse 5-complemental'. 'Self skew-symmetry' is just a property indicating the symmetry of a compound matrix about its own skew diagonal axis if it is expanded in 6 by 6 form.

The compound matrix $a\mid_{kl}^{ij}$ defined in Appendix B is found to be self skew-symmetric as well as self 5-complemental. The 5-complemental property actually comes from the skew-symmetry of the corresponding simple matrix a_{ij} , i.e., $a_{ij}=a_{5-j,5-i}$. Now we can see that the third and fourth components of $a\mid_{kl}^{ij}$ must be equivalent in some manner. For example, the self skew-symmetry of $a\mid_{14}^{12}$ is $a\mid_{34}^{23}$ and its self 5-complement is $a\mid_{34}^{14}$. Hence

$$a \mid_{14}^{12} = a \mid_{34}^{23} = a \mid_{34}^{14} = a \mid_{23}^{12}$$
.

where the last equality comes from the self skew-symmetry of $\alpha|_{34}^{14}$. Besides this, the compound matrix α is also inverse skew-symmetric as well as

inverse 5-complemental with coefficient $p=(-1)^{i+j+k+l}$. Hence we have

$$a^{-1}|_{kl}^{ij} = (-1)^{i+j+k+l} a|_{ij}^{kl} = (-1)^{i+j+k+l} a|_{5-l,5-k}^{5-l,5-k} = (-1)^{i+j+k+l} a|_{kl}^{ij}.$$

The compound matrix α has extremely good symmetry properties.

Similarly, we find that matrix E defined in equation (II-1-13) is inverse skew-symmetric with p = (-1) $^{i+j+k+l}$ q and q = $4k^2\nu_{\alpha}\nu_{\beta}/\rho^2$. But it possesses neither the properties of self symmetry nor the property of inverse 5-complemental. With further investigation, compound matrix is found to be inverse 5this complemental only for some of its components, which are ij=12, 34 kl=13, 24 for $E^{-1}|_{kl}^{ij}$, and ij=13, 24 kl=12, 34 for $E \mid \mathcal{U}$ with p equal to $(-1)^{i+j+k+l}$ times q or 1/qrespectively. Other such properties occur for ij=14, 23 with p=-q for $E^{-1}|_{kl}^{ij}$ and kl=14, 23 with p=-1/q for $E\mid_{R}^{H}$. For example, $E^{-1}\mid_{R}^{H}$ and $E\mid_{R}^{H}$ are both inverse 5complemental. Since these two are also inverse skewsymmetric, their third and fourth components are equivalent, i.e., $E^{-1}|_{14}^{12}=E^{-1}|_{23}^{12}$ $E|_{34}^{14}=E|_{34}^{23}$.

With the properties of compound α , E and E^{-1} revealed, it is not difficult to find that Haskell's matrices R and X which are composed of E^{-1} and α 's, behave as E^{-1} , and Z which is composed of α 's, behaves as α . In summary,

- (1) R or χ are inverse skew-symmetric.
- (2) $R \mid_{kl}^{12} R \mid_{kl}^{14} R \mid_{kl}^{23} R \mid_{kl}^{34} R \mid_{l3}^{13} R \mid_{24}^{13}$ and $R^{-1} \mid_{kl}^{13} R^{-1} \mid_{kl}^{24} R^{-1} \mid_{l2}^{13} R^{-1} \mid_{l2}^{24} R^{-1} \mid_{l2$
- (3) z is self skew-symmetric and 5-complemental.
- (4) Z is inverse skew-symmetric and 5-complemental.

APPENDIX D

Perturbation of Surface Wave Energy Integrals

In this appendix, we will evaluate $\frac{\partial}{\partial k}R|_{12}^{12}$ by using the variational principles. First, the eigenfunctions are perturbed by varying the wavenumber k about its stationary value. In such a perturbation state, the stresses at the free surface are small but not zero, and still satisfy

$$K_N = RB_1$$

i.e.

$$\begin{bmatrix} 0 \\ 0 \\ A' \\ B' \end{bmatrix} = \begin{bmatrix} R_{11} & R_{12} & R_{13} & R_{14} \\ R_{21} & R_{22} & R_{23} & R_{24} \\ R_{31} & R_{32} & R_{33} & R_{34} \\ R_{41} & R_{42} & R_{43} & R_{44} \end{bmatrix} \begin{bmatrix} \varepsilon \\ 1 \\ T_{z_1} \\ T_{\tau_1} \end{bmatrix}$$

where the vectors K_N and B_i have been normalized by U_{z_i} . The first two rows give

$$T_{s_1} = -\left[\begin{array}{ccc} \varepsilon R \mid_{14}^{12} + R \mid_{24}^{12} \end{array} \right] / R \mid_{34}^{12}$$

$$T_{r_1} = \left[\begin{array}{ccc} \varepsilon R \mid_{13}^{12} + R \mid_{23}^{12} \end{array} \right] / R \mid_{34}^{12}$$
(D-1)

Note that ε is not the same ellipticity as defined in equation (III-1-4), since the system is no longer in an eigen state. If the system is really being characteristically excited, then $T_{z_1} = T_{r_1} = \emptyset$ and we have the same ellipticities as those in equation (III-1-4).

Return to the differential equation (II-1-11)

$$\frac{dT_z}{dz} = -\rho U_z + k T_r$$

$$\frac{dT_r}{dz} = \left(\rho - \xi \frac{k^2}{\omega^2}\right) U_r - k \sigma T_z$$

$$\frac{dU_z}{dz} = -k \sigma U_r + \frac{\omega^2}{\lambda} \sigma T_z$$

$$\frac{dU_r}{dz} = k U_z - \frac{\omega^2}{\mu} T_r ,$$
(D-2)

where

$$\xi = \frac{4\mu (\lambda + \mu)}{\lambda + 2\mu}$$
$$\sigma = \frac{\lambda}{\lambda + 2\mu} .$$

Multiply the first equation of (D-2) by U_z and integrate it with respect to z from \emptyset to infinity:

$$\begin{split} U_z \ T_z \,|_{0}^{\infty} &= \int\limits_{0}^{\infty} \left[\, -\rho U_z^{\,2} + k U_z \, T_r \, + \, T_z \, \frac{d}{dz} \, U_z \, \, \right] \, dz \\ \\ &= \int\limits_{0}^{\infty} \left[\, -\rho U_z^{\,2} + k U_z \, T_r \, - \, k \, \sigma U_r \, T_z \, + \, \frac{\omega^2}{\lambda} \sigma T_z^{\,2} \, \right] \, dz \quad . \end{split}$$

Replacing the integrals by the summation over layers represented by I_{ij} 's,

$$-U_{z_1}T_{z_1} = \sum_{n} -\rho_n(I_{22})_n + k(I_{24})_n - k\sigma_n(I_{13})_n + \omega^2 \frac{\sigma_n}{\lambda_n} (I_{33})_n, \qquad (D-3)$$

where because of the conservation of energy, the radiation condition requires that U_{z_1} and T_{z_1} vanish at great depth. Similarly, the second equation in (D-2) can be integrated to become

$$-U_{r_1}T_{r_1} = \sum_{n} \left(\rho_n - \xi_n \frac{k^2}{\omega^2} \right) (I_{11})_n - k \, \sigma_n (I_{13})_n + k \, (I_{24})_n - \frac{\omega^2}{\mu_n} (I_{44})_n . \tag{D-4}$$

Multiply equations (D-3) and (D-4) by ω^2 and mutually subtract,

$$\begin{split} &-\omega^2 \left(\ U_{r_1} T_{r_1} - U_{z_1} T_{z_1} \right) \\ &= \sum_n \left[\ \left(\rho_n - \xi_n \frac{k^2}{\omega^2} \right) (I_{11})_n + \rho_n (I_{22})_n - \omega^2 \frac{\sigma_n}{\lambda_n} \ (I_{33})_n - \frac{\omega^2}{\mu_n} (I_{44})_n \right] \, \omega^2 \, . \end{split}$$

In this equation, the terms on the right are nothing but the Lagrangian of Rayleigh waves as given by equation (III-1-13),

$$-\omega^2 \left(U_{r_1} T_{r_1} - U_{z_1} T_{z_1} \right) = L_R \equiv \omega^2 I_0 - k^2 I_1 - 2k I_2 - I_3 \quad . \tag{D-5}$$

Since all of the eigenfunctions are normalized (recall our definition of the energy integrals in equation III-1-10), $U_{z_1}=1$ and $U_{r_1}=\varepsilon$. The terms inside the parentheses on the left side of equation (D-5) can be further expanded after substitution from equation (D-1),

$$\begin{split} &U_{r_1}T_{r_1}-U_{z_1}T_{z_1}\\ &=\varepsilon T_{r_1}-T_{z_1}\\ &=\left[\varepsilon(\varepsilon R\mid_{13}^{12}+R\mid_{23}^{12})+\left(\varepsilon R\mid_{14}^{12}+R\mid_{24}^{12}\right)\right]/R\mid_{34}^{12}\\ &=\left[\varepsilon^2 R\mid_{13}^{12}R\mid_{13}^{12}+2\varepsilon R\mid_{14}^{12}R\mid_{13}^{12}+R\mid_{24}^{12}R\mid_{13}^{12}\right]/\left(R\mid_{34}^{12}R\mid_{13}^{12}\right), \end{split}$$

where we have used

$$R \mid_{14}^{12} = R \mid_{23}^{12}$$
.

Since

$$R \mid_{12}^{12} R \mid_{34}^{12} + R \mid_{14}^{12} R \mid_{23}^{12} = R \mid_{13}^{12} R \mid_{24}^{12} ,$$

equation (D-6) becomes

$$\varepsilon T_{r_1} - T_{z_1} = \frac{R \mid \frac{12}{12}}{R \mid \frac{12}{13}} + \frac{\left[\varepsilon R \mid \frac{12}{13} + R \mid \frac{12}{23}\right]^2}{R \mid \frac{12}{34} \mid R \mid \frac{12}{13}}.$$

Substitute into equation (D-5),

$$-\omega^{2}\left[\frac{R|_{12}^{12}}{R|_{13}^{12}} + \frac{\left[\varepsilon R|_{13}^{12} + R|_{23}^{12}\right]^{2}}{R|_{34}^{12}R|_{13}^{12}}\right] = \omega^{2}I_{0} - k^{2}I_{1} - 2kI_{2} - I_{3}.$$

Because the system slightly departs from the eigenstate, it follows that a perturbation in k results in

$$-\omega^2 \frac{\frac{\partial}{\partial k} R \mid_{12}^{12}}{R \mid_{13}^{12}} = -2kI_1 - 2I_2 ,$$

where we have set $R \mid_{12}^{12} \sim 0$ and consequently ε is very close to the value defined in equation (III-1-4), hence only the dominant term $\frac{\partial}{\partial k} R \mid_{12}^{12}$ is left. The group velocity as expressed in terms of integrals (Jeffreys, 1961) has the form

$$U = \frac{I_1 + I_2/k}{c I_0} .$$

The result, i.e., the amplitude factor in terms of energy integrals, is obtained

$$\frac{R \mid_{13}^{12}}{\frac{\partial}{\partial k} R \mid_{12}^{12}} \frac{k}{\omega^2} = \frac{1}{2(I_1 + I_2/k)} = \frac{1}{2c \ U \ I_0} . \tag{D-7}$$

For the Love waves, the derivation is similar. Start with the differential equation

$$\frac{d}{dz}T_{\vartheta} = (\mu k^2 - \rho \omega^2) U_{\vartheta} .$$

Multiplying both sides by $\boldsymbol{U}_{\boldsymbol{\vartheta}}$ and integrating, yields

$$-T_{\vartheta_1} = -\omega^2 I_0 + k^2 I_1 + I_2 . \tag{D-8}$$

Then from

$$\begin{split} \mathbf{K}_{\mathcal{N}} &= \mathbf{R} \; \mathbf{B}_1 \\ &= \mathbf{R} \; \left[\; \; \mathbf{1} \; , \; T_{\vartheta_1} \; \right]^T \; , \end{split}$$

the perturbed stress T_{ϑ_1} is

$$T_{v_1} = -R_{55} / R_{56}$$
.

Equation (D-8) becomes

$$\frac{R_{65}}{R_{66}} = -\omega^2 I_0 + k^2 I_1 + I_2 \ .$$

Again, the right side is the Love wave Lagrangian. After taking the derivative with respect to k, and putting $R_{66}\sim 0$ we have

$$\frac{k R_{56}}{\frac{\partial}{\partial k} R_{55}} = \frac{1}{2I_1} = \frac{1}{2c U I_0} , \qquad (D-9)$$

where the Love group velocity is just

$$U = \frac{I_1}{c I_0} \tag{D-10}$$

which is obtained by simply perturbing the Love wave Lagrangian.

APPENDIX E

Description of the Eigenfunction programs

To implement the eigenfunction theory as developed in chapter III, several FORTRAN computer programs have been set up using a DEC PDP 11/70 minicomputer in the department of Earth and Atmospheric Sciences, Saint Louis University. This appendix gives a brief description of these programs in order to facilitate their utilization and maintenance.

Because of the restriction of core size (about 60K bytes), we divided the computation into five separate programs which communicate via intermediate data files. The data files are stored in a high speed, large volume The designation skeleton is inherited Herrmann (1974, 1978b), although the theoretical and numerical techniques are modified. These programs suited, especially, to short periods (less than 1 second) and complicated structures. Figure 47 this sequence of programs and their input/output. Table 5 lists the information to be used when running these programs.

(1) surface81: This program searches for dispersion values over the range of phase velocity between $\emptyset.8\,\beta_{\rm min}$ and β_N for some fixed set of periods. It is suggested that these periods had better be arranged in

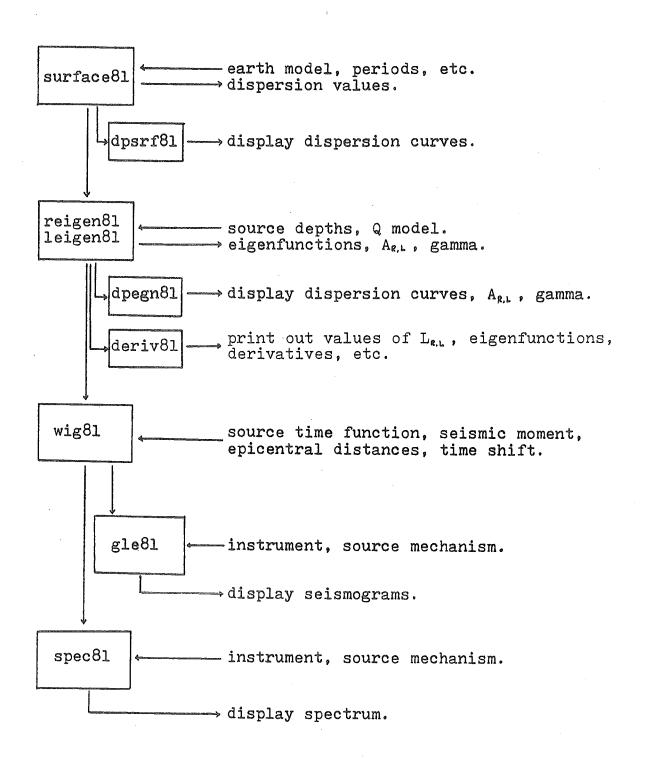


Figure 47. Flow chart of eigenfunction programs.

TABLE 5

Information for Synthesizing Seismograms

I.	Sampling Information:	
	l. sampling rate, dt (sec):	
	2. total sampling point:	
	3. frequency range (Hz):	
II.	Wave Type:	
	l. Rayleigh or Love?	······································
	2. component (Z,R,T,NS,EW)?	in de Arreston
	3. waveform (disp., vel., accel.)?	n's holden a
II	. Source Model:	
	1. source depth (Km):	
	2. seismic moment (dyne-cm):	
,	3. source mechanism (dip,slip,strike):	-15-25-0
	4. source time function:	
IV.	Earth model:	
	1. velocity model:	
	2. Q model:	
٧.	Receiver:	
	l. receiver position (r,Az):	
	2. original time shift (t0):	
	3. instrument (LP,SP,model):	

equally spaced frequency (inverse of period) so that a FFT can be used to transform spectral responses to time domain without using any sort of interpolation. The other input parameters are the structure which consists of the thickness of each layer, P-wave velocity, S-wave velocity and density. The number of layers and modes can be as large as 500.

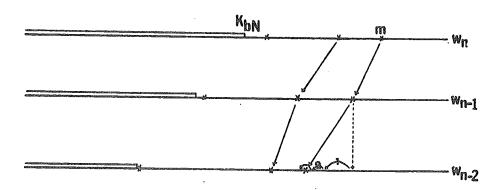
As pointed out in section 2.4, the search of poles easier to carry out in the wavenumber-frequency plane than in the phase velocity-period plane. The reason is that equation (II-4-6) can give reasonably estimated steps to progress for a particular mode durthe pole searching in the k-w plane. Figure 48 describes such a searching procedure. Note jump steps are frequency, wavenumber and mode dependence. In some cases, however, the dispersion values vary quite abruptly near the various P and S-wave layer velocities as illustrated in Figures 49 and 5Ø. figures display the dispersion curves of the CUS model at the high frequency range. The sharp bends and kinks result in very irregular pole spacing at a given frequency, which causes difficulty in pole searching. overcome this, two more constraints to the values provided by equation (II-4-6) are needed:

$$\Delta k_{\min} = (k_{\max} - k_{\min}) / (200 + 200 \cdot frequency)$$

 $\Delta k_{\text{max}} = 0.8 \cdot (k \mid_{\text{mode}=m} - k \mid_{\text{mode}=m-1}) \mid_{\text{at previous freq}}$

These serve as the lower and higher bounds for the

JUMPING METHOD



First jump:
$$\frac{\Delta}{\alpha}$$

Second jump:
$$k\frac{\Delta c}{c}$$

which is restricted by

$$\Delta k_{\min} < k \frac{\Delta c}{c} < \Delta k_{\max}$$

$$\Delta k_{\rm min} = (\ k_{\rm max} - k_{\rm min}\)/(\ 200 + 200 \cdot frequency\)$$

$$\Delta k_{\max} = 0.8 * (k \mid_{mode=m} - k \mid_{mode=m-1}) \mid_{at previous freq}$$

Figure 48. Jumping method for searching poles.

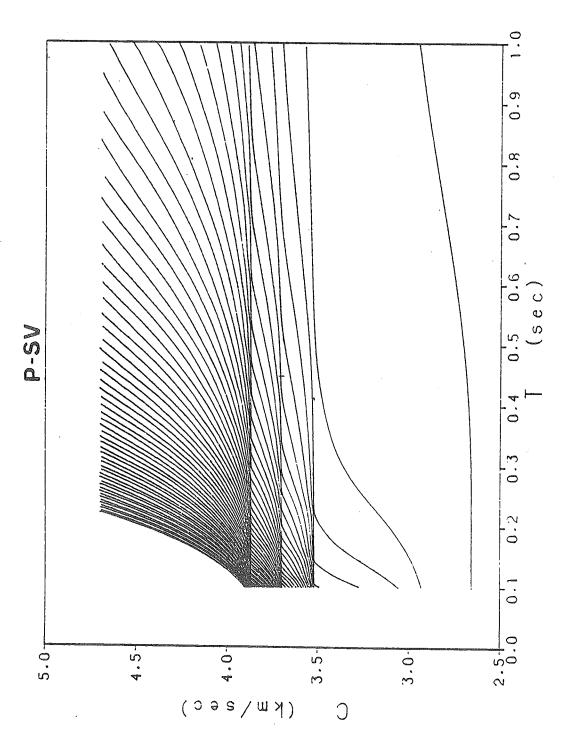


Figure 49. Phase velocity dispersion curves for the CUS model at short period range.

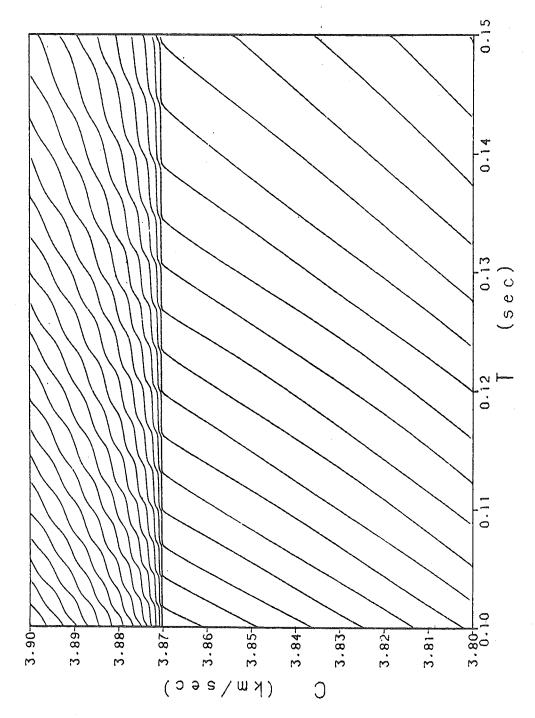


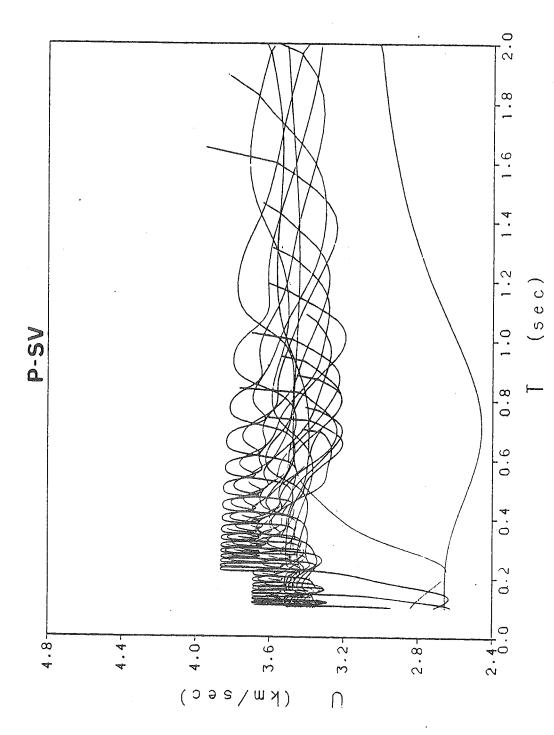
Figure 50. An expansion of Figure 49 for period between 0.10 and 0.15 second and phase velocity between 3.8 and 3.9 km/sec.

value of $\frac{k}{c}\Delta c$ from equation (II-4-6), i.e., if $\frac{k}{c}\Delta c$ is greater than Δk_{\max} or smaller than Δk_{\min} , Δk_{\max} or Δk_{\min} will be used to make several small jumps after the first jump $\frac{\Delta \omega}{c}$. This algorithm has been proved to be very efficient. For most models the pole is usually bracketed after not more than three or four searching jumps. The result can be displayed and checked by a plotting program 'dpsrf81'. All of the calculations in this program are performed in double precision.

(2) reigen81 (leigen81): This program computes the eigenfunctions and energy integrals which in turn give the group velocity, amplitude factor and attenuation coefficient. Reigen81 is for the P-SV case and leigen81 for SH. To save data file space, only the eigenfunction values at the source depth are stored. The source depth can be entered as several different values, and the output data for different source depths are stored in different output files. A Q model is input at The attenuation coefficients are determined using the perturbation theory of Anderson et al (1965). is not difficult to incorporate any sort of frequency dependent Q by slightly modifying the subroutine in this program (Mitchell, 1980, 1981). 'qamma' partial derivatives of phase velocity at different layer boundaries can also be stored for further study. To check the result, one can either print out the value the Lagrangian for each dispersion pair or use the

program 'dpegn81' to plot group velocity, amplitude factor, etc. For good results, the curves plotted should be smooth, such as those shown in Figure 51. All of the calculations in these programs are done in double precision.

- (3) wig81: This program computes the spectra stations at a given set of epicentral distances. A particular source time function needs be specified here. In order to be able to treat long duration time series, a separate program 'bigfft' which executes FFTinvolved. These two independent programs are connected by a 'system call'. Using this, the longest time series which can be created is 8192 points long, which is the largest dimension allowed in program 'bigfft'. In other programs only a vector with length of at most 1024 points is permitted. A linear interpolation applied to generate enough data for spectra before taking the inverse Fourier transform, if necessary. spectral data corresponding to different fundamental source types as listed in section 5.1 are generated and The contributions from different modes can be stored. separated as those used for higher mode study (Cheng and Mitchell, 1981).
- (4) gle81 (spec81): The program 'gle81' reads the spectra data created in the previous program and generates the seismograms for each component at each



Group velocity dispersion curves for the CUS model. Note that the low order higher modes have group velocities around 3.5 km/sec. Figure 51.

receiver location after calling 'bigfft' to take inverse FFT. The seismogram is then displayed by a plotter. The program 'spec81' just reads and plots the spectra data. The focal mechanisms, as well as the receiver azimuths, are entered to take into account the directivity dependence. Several instrument responses, long and short periods, are included to take into account the instrument effect.

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VITA AUCTORIS

Chien-Ying Wang was born on May 24, 1951 Tainan, Taiwan, Republic of China. He attended the National Central University, Chung-Li, Taiwan, from 1969 to 1973, and received his B.S. degree in Physics in 1973. After two years of military duty, he became a graduate student at the Oceanography Institute of National Taiwan University, Taiwan, and received an degree in Geophysics in 1977. During the years M.S. 1977 to 1981 he has been a graduate student and research assistant in the Department of Earth and Atmospheric Sciences of Saint Louis University, St. Louis, Missouri.

COMPUTER PROGRAMS

The computer programs listed here represent substantial improvements to those given by Herrmann (1978b). All programs are written in FORTRAN 77 and are run on a PDP 11/70 under the UNIX* operating system. This system makes use of the separate Instruction and Data space feature of the PDP 11/70 processor. Thus even though the PDP 11/70 is a 16 bit minicomputer, the text and data spaces can each be as large as 64 K bytes. However, because of the size limitations of even this machine, one program of Herrmann (1978b), wiggle, was split into two programs, wig81 and gle81. Double precision is used, which means that a floating point number uses 64 bits, or has about 16 significant figures.

The only non-standard usage is the

call system ('bigfft', kturn)

Under UNIX, this permits one program to initiate another program and waits until that program completes before proceeding. The program bigfft performs a Fast Fourier Transform of length 2n by performing a Fast Fourier Transform of length n using the constraint that the time series is purely real. The analog of "call system" may not exist on other computers, but then those computers may be more than 16 bit machines and hence the subterfuge used here may be removed by replacing the sequence of writing, reading and closing temporary files with array access and the "call system" by a

^{*}UNIX is a trademark of Bell Laboratories.

"call bigfft" where "bigfft" is now a subroutine rather than a separate program.

The plotting programs make use of subroutine calls to a CALCOMP* plotter. These are

open plotter CALL PLOTS (0, 0, LDEV) CALL PLOT (X, Y, ±IPEN) move plotter to X,Y with pen up (IPEN= ±3) or pen down (IPEN=±2) and reset origin at new position if IPEN=negative CALL FACTOR (FACT) multiply all movements by FACT use pen INP CALL NEWPEN (INP) CALL SYMBOL (X, Y, HEIGHT, ICHAR, ANGLE, + NCHAR) write CALL SYMBOL (X, Y, HEIGHT, INTEQ, ANGLE, -ICODE) text array plot symbol INTEQ with pen up/down, ICODE=-1/-1 during move CALL NUMBER (X, Y, HEIGHT, FLOATNUMB, ANGLE, ±NDEC) write number at X, Y CALL SCALE (ARRAY, AXLEN, NPTS, ±INC) define FIRSTV ARRAY (NPTS+1) DELTAV ARRAY (NPTS + INC + INC +1)CALL AXIS (X, Y, TITLE, ±NCHAR, AXLEN, ANGLE, FIRSTV, DELTAV) plot axis with label CALL LINE (XARRAY, YARRAY, NPTS, INC, LINTYP, INTEQ) draw line with/without symbol at each point

FIRSTV and DELTAV

array

must be defined for each

^{*}CALCOMP is a trademark.

Eigenfunction Programs

l.	surface81		•	•				•		٠	•		1
2.	dpsrf8l .		•	•							•		19
3.	reigen81.					•	•				•	•	24
4.	leigen81.				. •					•	•		43
5.	dpegn81 .	•	•		٠	•	•	•	•	•	•		52
6.	deriv81 .	•							•		•		57
7.	wig81	•					•	•	•		•	•	59
8.	gle81				•								72
9.	spec81				•								86
10.	bigfft												100

```
program surface81.f
C
C
      477 -i -I2 surfaceB1. f -o surfaceB1
r
C
      This program calculates the dispersion values for any
C
      layered model, any frequency, and any mode.
c
C
      This program will accept one liquid layer at the surface.
C
      In such case ellipticity of rayleigh wave is that at the
C
      top of solid array. Love wave communications ignore
C
      liquid layer.
C
C
      Program developed by Robert B Herrmann Saint Louis
C
      univ. Nov 1971, and revised by C. Y. Wang on Oct 1981.
C
C
      All processes are performed in the wavenumber-frequency
c
      domain.
C
      INPUT:
C
                       (eg. 2,1000)
      1. mmax, mode
c
         -mmax: >O number of layers including halfspace
                <0 end program</p>
c
                ≕O use previous model with new options
C
                    number of modes for which dispersion curves
         -mode:
C
                    are desired. (1000 is the maximum)
C
C
                                        40, 0, 6, 15, 3, 55, 2, 8
      2. d(i),a(i),b(i),rho(i)
                                  (ea.
C
                                             8. 09, 4. 67, 3. 3
                       i=1, mmax
C
         -d(i): layer thickness (km)
C
         -a(i): P wave velocity (km/sec)
C
         -b(i): S wave velocity (km/sec)
C
                                  (gm/cm3)
         -rho(i); density
C
C
      3 igphl, igphr, icut, idispl, idispr, ipunch
C
                                 1, 1, 1, 1, 1, 1, 3.0)
                           (eg.
C
                  =1 Love wave period equation plot
          -icohl
C
                  =O not plot
C
                  =1 Rayleigh wave period equation plot
          -iqphr
C
                  =0 not plot
C
                  =1 find the cutoff periods of Love and
          -icut
C
                     Rayleigh waves.
 C
                  =0 not
 C,
          -idispl =1 find the dispersion values and amplitude
 C
                     factor for Love wave
 C
                  =0 not
 C
          -idispr =1 find the dispersion values, amplitude
 C
                     factor and ellipticity for Rayleigh wave
 C
                  =O not
 C
          -ipunch =1 store and print the output of dispersion
 C
                     values.
 C
```

```
=O store the dispersion values only
C
                =-1 print the dispersion values only
C
C
      4. if igph1 or igphr =1
C
         kk, c1, c2, dc
C
         t(i) i=1, kk
                            (eq. 10,3,5,8,1,0,1
c
                                   1, 2, 3, 4, 5, 6, 7, 8, 9, 10
C
         -kk: number of periods along abscissa
C
         -c1: lower phase velocity limit
C
         -c2: upper phase velocity limit
C
         -dc: phase velocity increment
C
C
         -t(i): periods for which period equation is computted
C
C
      5. if icut = 1
C
C
         kmax, t1, dt, c1
                            (eg. 4, 18, 0, -0, 2, 4, 67)
         -kmax: number of cutoff periods to be found for cl
C
         -t1: initial period of search
C
               period increment (can be negative)
C
                phase velocity for which kmax cutoffs are desired
         --c1:
C
C
      6: if idispl or idispr =1
C
         kmax, t1, dt, c1
C
         if ti=0.0
C
         t(i) i=1, kmax (eg. 10, 0, 0, 0, 0, 3, 5, 0, 005
C
                  1, 1, 25, 1, 5, 1, 75, 2, 0, 2, 25, 2, 5, 2, 75, 3, 0, 3, 5, )
C
C
         -kmax: number of periods for which phase velocity
C
                 to be found
C
         -t1: =0.0 the array t(i) i=1, kmax are read in
C
              >0.0 initial period
C
                                 t(i)=t1+(i-1)*dt will
C
         -dt: period increment
             be generated
C
         -c1: initial phase velocity guess. better =0.8*b(1).
C
C
                (only when t1=0.0)
C
                kmax array of periods at which dispersion are
C
                calculated.
C
C
      7. if ipunch >=0
C
         name of Love wave output file (if idisp1 =1)
C
         name of Rayleigh wave output file (if idispr =1)
C
C
      8. stop the program (use -1.0)
         or go back to 1 for other models.
C
common d(500), a(500), b(500), rho(500), mmax, ipunch
      common t(4096),c(500),mode,llw,twopi
   10 format(1h , 5x, 4f10, 4)
   20 format(1h , 15x, 3f10, 4)
   30 format('\f')
   40 format(/, 1h , 19x, 'crustal model ', /1h )
```

```
50 format(1h ,7x,' thick
                                            s-vel
                                                      density(//)
                                  p-vel
      twopi=2. *3. 141592653
  100 write(6,*) 'enter mmax(layers), mode(modes)'
      read(5, *)mmax, mode
      mmax = number of layers to be read in, including
C
              the halfspace.
C
      mode = number of modes for which dispersion curves are
C
             desired can be as large as 1000.
C
      if(mmax) 400,300,200
  200 continue
      1 = mmax - 1
      d # thickness of layer in kilometers
C
C
      a = compressional wave velocity in km/sec
      b = transverse wave velocity in km/sec
C
      rho = density in gm/cc
C
      write(6,*) 'enter d.a.b.rho'
      do 220 i=1,1
      read(5, *) d(i), a(i), b(i), rho(i)
  220 continue
      read(5,*) a(mmax), b(mmax), rho(mmax)
      write(6,30)
      write(6,40)
      write(6,50)
      11w = 1
      if(b(1), le, 0, 0) llw=2
      do 240 i=1,1
  240 write(6,10) d(i),a(i),b(i),rho(i)
      write(6,20) a(mmax), b(mmax), rho(mmax)
  300 continue
      d(mmax) = 0.0
      write(6,*) 'igph1, igphr, icut, idisp1, idispr, ipunch'
             gt O calls love wave dispersion plot
C
              at O calls rayleigh wave dispersion plot
C
      iaphr
              gt O calls search for love and rayleigh higher
C
      icut
                   mode cutoff
C
      idispl qt O love wave dispersion curve
C
      idispr gt O rayl wave dispersion curve
C
C
      ipunch eq 1 store and print the dispersion values
              eq O store the dispersion values only
C
              eq -1 print the dispersion values only
C
      read(5,*) igph1, igphr, icut, idisp1, idispr, ipunch
    if(igphl.gt.O.or.igphr.gt.O) call gphdis(igphl.igphr)
      if(icut.gt.O) call cutoff
      if(idispl.qt.O, or.idispr.qt.O) call disper(idispl.idispr)
      go to 100
  400 continue
      stop
      end
C
c -
```

subroutine gphdis(igph1,igphr)
doubleprecision wvno.omega

```
common d(500), a(500), b(500), rho(500), mmax, ipunch
      common t(4096), c(500), mode, llw, twopi
      character*1 num(10), ks(60), kx(60)
      data num/(1/, /2/, /3/, /4/, /5/, /6/, /7/, /8/, /9/, / //
      This subroutine graphically displays the sign of the
C
      love or rayleigh wave period equation in the c-t plane.
C
      The dispersion curve is the line of zeroes.
C
      gphdis reads in up tp 59 different periods to form
C
      abscissa of plot.
C
      The ordinate varies from c1 to c2 in increments of dc.
C
      kk is the number of abscissa values
C
      c1 is less than c2
C
    · dc is positive
C
   10 format('\f')
   20 format(5x, 'period for abscissa of following graph'//)
   30 format(1x, 12(i3, f7, 2))
                                     function (//)
   40 format(19x, 'plot of
                             love
   50 format(19x, 'plot of rayleigh function'//)
   60 format(2x, f8, 4, 59(a1, a1))
      write(6,*) 'gphdis-kk,cmin,cmax,dc'
      read(5,*) kk,c1,c2,dc
      write(6,*) 'enter periods'
      read(5,*) (t(i), i=1, kk)
      if(kk.at.59) kk = 59
      if(dc. lt. 0. 0) dc = -dc
      if(c1.1t.c2) go to 100
      dum = c1
      c1 = c2
      c2 = dum
  100 continue
      write(6,10)
      write(6,20)
      write(6,30) (i, t(i), i=1, kk)
      do 700 ifunc = 1/2
      if(dc.eq.O.O) go to 700
      go to (200,300), ifunc
  200 if(igph1.le.0) go to 700
      write(6,10)
      write(6,40)
      go to 400
  300 if(igphr. le. 0) go to 700
      write(6,10)
      write(6,50)
  400 \text{ cc} = \text{c2}
  500 do 600 i=1,kk
       omega=dble(twopi)/dble(t(i))
      wvno=omega/dble(cc)
       del = dltar(wvno,omega,ifunc)
       1 = mmax
       110 = 1/10
       1 = 1 - 110*10
       ky(i)=num(1)
```

ks(i)='+'

```
if(del. lt. O. O) ks(i)='. '
      if(del. lt. 0. 0) kx(i)=''
  600 continue
      write(6,60) cc, (ks(1), kx(1), l=1, kk)
      cc = cc - dc
      if(cc.ge.c1) go to 500
  700 continue
      return
      end
C
C
      subroutine cutoff
      double precision wyno, omega, cc
      common d(500), a(500), b(500), rho(500), mmax, ipunch
      common t(4096), c(500), mode, llw, twopi
      kmax = number of cutoff periods to be found for phase
C,
C
             velocity c1.
C
      t1 = initial period in search
      dt = period increment negative if starting at high
C
C
           period and going toward shorter period
C
      c1 = phase velocity for which kmax cutoffs are being found
      this routine finds both love and rayleigh cutoffs
C
   10 format('\f'//)
  20 format(///16x, 'higher mode cutoff periods'//17x,
                 ',23x,' rayleigh '//2(11x,'period',10x,
           love
     2 'ph vel')///)
  30 format(1h ,7x,f10,4,6x,f10,4,7x,f10,4,6x,f10,4)
      write(6,*) 'cutoff-enter kmax,t1,dt,c1'
      read(5,*) kmax, t1, dt, c1
      cc =dble(c1)
      write(6,10)
      write(6,20)
      tt=t1
      do 700 ifunc=1,2
      do 600 j=1, kmax
      tcut=0.0
      t1=tt
      omega=dble(twopi)/dble(t1)
     wvno=omega/cc
     del1 = dltar(wvno.omega.ifunc)
 100 continue
     t2 = t1 + dt
     if(t2.1e.0.0) go to 700
     omaga=dble(twopi)/dble(t2)
     wyno≕omega/cc
     del2=dltar(wvno,omega,ifunc)
     if(sign(1.,del1).ne.sign(1.,del2)) go to 200
     t1 = t2
     del1 = del2
     go to 100
 200 if(abs (t1 + t2) - 0.0001) 500,500,300
 300 \ t3 = (t1 + t2) * 0.5
```

```
omega=dble(twopi)/dble(t3)
      wyno=omega/cc
      del3 = dltar(wyno,omega,ifunc)
      if(sign(1.,del1), ne. sign(1.,del3)) go to 400
      t1 = t3
      del1 = del3
      go to 200
 400 t2 = t3
      de12 = de13
      go to 200
  500 continue
      tcut=0.5*(t1+t2)
      if(t2.le.0.0) go to 700
      if(ifunc.eq.1) t(j)=tcut
      if(ifunc.eq.2) t(j+500)=tcut
  600 continue
  700 continue
      write(6,30) (t(j),c1,t(j+500),c1,j=1,kk)
      return
      end
C -
C
      subroutine disper(idispl/idispr)
      double precision omegaO, omega1
      double precision eroot(500), cphase(500)
      common d(500), a(500), b(500), rho(500), mmax, ipunch
      common t(4096), c(500), mode, llw, twopi
      common/stor/ eroot(500), cphase(500), nlost, index, nroot1
      character≉25 names
      the root determination section is one of interval halving
\mathbf{C}
      once a zero crossing has been found.
C
C
      the number of modes is allowed to be as large as 500,
C
      and the number of periods as 4096.
C
C
      t1 = initial starting period
C
      kmax = number of period s for which phase velocity is to
C
      be determined
C
      if t1 = 0 program reads in array of t(i) periods instead
c
      of computing them
      dt = period increment. next period t2 = t1 + dt
C
      c1 = initial phase velocity guess. make sure it is
C
      outside desired result
    5 format(a)
   10 format(1x, 'improper initial value no zero found ')
      write(6,*) 'enter kmax,t1,dt,c1 ; t1<≕O to read in array'
      read(5,*) kmax,t1,dt,cmin
      if(t1.qt.O.O) qo to 100
      write(6,*) 'disper-enter kmax periods'
      read(5, *) (t(j), j=1, kmax)
      go to 200
```

```
100 t(1) = t1
     do 150 i =2, kmax
 150 \ t(i) = t(i-1) + dt
 200 continue
     if(ipunch. lt. 0) go to 400
     if(idispl.le.O) go to 300
     write(6,*) 'enter name of Love wave output file: '
     read(5,5) names
     open(1, file=names, status='new', form='unformatted')
     rewind 1
 300 continue
     if(idispr. le. 0) go to 400
     write(6,*) 'enter name of Rayleigh wave output file:'
     read(5,5) names
     open(2, file=names, status='new', form='unformatted')
     rewind 2
 400 continue
      open(3, file='tmpsrf.d', status='scratch', form='unformatted')
      do 2000 ifunc=1,2
      ii = ifunc
      nlost = 0
      index = 0
      do 500 i=1,500
      c(i)=0.0
      eroot(i)=0.0
      cphase(i)=0.0
 500 continue
      rewind 3
      if(ifunc. eq. 1. and. idispl. le. 0) go to 2000
      if(ifunc. eq. 2. and. idispr. le. 0) go to 2000
      nroot1=1000
      do 1800 k = 1 / kmax
      t1 = t(k)
      omegal=dble(twopi)/dble(t1)
      if(k, gt. 1) omegaO=dble(twopi)/dble(t(k-1))
      index=index+1
CXXXXXXXX
      kmode=0
      call poles(ii,omegaQ,omega1,cmin)
      kmode=nroot1
CXXXXXXXX
      if(k. eq. 1. and. kmode. eq. 0) write(6, 10)
      if(k, eq. 1, and, kmode, eq. 0) go to 2000
      if(k.eq.1) lmode=kmode
      write(3) ifunc, kmode, t1
      do 1750 i=1, kmode
      cc=omega1/eroot(i)
      write(3) cc
 1750 continue
      i = -1
      tmp=0.0
      if(k, eq, kmax) write(3) i, i, tmp
 1800 continue
```

```
call output(ifunc, lmode)
     if(ipunch. lt. 0) go to 2000
    write(ifunc) mmax
     do 1810 i=1, mmax
    write(ifunc) d(i),a(i),b(i),rho(i)
1810 continue -
    write(ifunc) kmax
     rawind 3
1850 continue
     read(3) ifun, kmode, t1
     write(ifunc) ifun, kmode, tl
     if(ifun.le.Q) go to 1900
     do 1860 i=1, kmode
     read(3) cO
     write(ifunc) cO
1860 continue
     go to 1850
1900 continue
     close(ifunc)
2000 continue
     close(3, status='delete')
     return
     end
     subroutine output(ifunc, 1mode)
     common d(500), a(500), b(500), rho(500), mmax, ipunch
     common t(4096),c(500),mode,llw,twopi
                                    mode #', i3//21x,
  10 format(///15x, 'Love wave
                         phase vel')
          period
  20 format(1h , 12x, i3, 1x, f15, 10, 2x, f15, 10)
                                        mode #', i3//21x,
  30 format(///14x, 'Rayleigh wave
                           phase vel')
    1 ' period
  50 format('\f')
     write(6,50)
                        number of modes at the lowest period = ',
     write(6, 4) '
                  1mode
     nm≠O
     c1=b(mmax)
 100 continue
     nm=nm+1
     if(nm.gt.lmode) go to 600
     if(ipunch. ne. O. and. ifunc. eq. 1) write(6, 10) nm
     if(ipunch. ne. 0, and. ifunc. eq. 2) write(6, 30) hm
     rewind 3
     nt≖O
     do 200 i=1,500
     c(i)=0.0
 200 continue
     k k ≕O
 250 continue
     read(3) ifun, kmode, tO
```

```
if(ifun) 450,300,300
 300 continue
     nt=nt+1
      do 400 i=1,kmode
      read(3) cO
      if(i.ne.nm) go to 400
      c(nt)=c0
      kk=nt
 400 continue
      go to 250
  450 continue
      if(ipunch.eq.O) go to 100
      write(6,20) (i,t(i),c(i),i=1,kk)
      go to 100
  600 continue
      return
      end
C
C
      subroutine poles(ifunc.omegaO.omega.cmin)
      double precision eroot(500), cphase(500)
      double precision omegaO, omega, dk, domega, dkO, tadd
      double precision w/mn, w/mx, c1, c2, c3, c4, s, t, /phase, dc
      double precision wvm(500),c10
      common d(500), a(500), b(500), rho(500), mmax, ipunch
      common perd(4096),c(500),mode,llw,twopi
      common/stor/ eroot(500),cphase(500),nlost,index,nroot1
      this routine finds the roots of period equation using
c
      regular halving method to initialize the first two
C
      frequencies, then followed by jumping method.
C
C
      epi = 1.e-8
      freq = omega/twopi
      wvmx = omega/dble(cmin)
      wvmn = omega/dble(b(mmax))
      nmx=200 + (freq*200)
      dk = (w \vee m \times - w \vee m n) \wedge n m \times d k
      if(nlost, eq. 1001) go to 2000
      if(index. gt. 2) go to 3000
 2000 continue
      find the poles using the regular halving method
      nmx is chosen for a 40 km crustal model. for shallower
C
       thickness a proportionately smaller nex can be used
      search for roots of period equation
       if(index.gt.2) write(6,*) '-at period=', twopi/omega,
                       ' return halving method.'
       do 80 i=11w, mmax
      wvm(i)=omega/dble(b(i))
   80 continue
      nroot = C
       c2 = wvmx
       del2 = dltar(c2,omega,ifunc)
```

```
lyr=11w
      1=رر
      do 500 i=1, nmx
      JJ==JJ-1
      if(j), ne, O) go to 500
      c10 = w \vee m \times -i * dk
      if(i, eq. nmx) c10=w \vee mn+0. O1*dk
      JJ = I
      kk = 1
      if(c10.gt.wvm(lyr)) go to 90
      kk and jj represent a denser seraching when phase velocity
C
      = S wave velocity. Their values can be changed as kk=3*lyr
C
      JJ=8.
      kk = 10.0*(a(lyr+1)/a(lyr))+0.5*lyr
      lyr = lyr+1
      JJ = 5
   90 continue
      dkO = dk/float(kk)
      do 400 j=1, jj
      do 400 k=1, kk
      if(nroot, eq. mode) go to 510
      jk = kk*(j-1)*k
      c1 = c10+dk-dk0*float(jk)
      if(c1. lt. wvmn) go to 510
      del1 = dltar(c1, omega, ifunc)
      if(sign(1, 0, del1)*sign(1, 0, del2), qe, 0, 0) go to 350
      nroot = nroot + 1
      c4 = c2
      del4 = del2
      do 200 ii=1,15
      c3 = 0.5*(c1+c4)
      del3 = dltar(c3,omega,ifunc)
      if(sign(1.0,del1)*sign(1.0,del3).ge.0.0) go to 100
      de14 = de13
      c4 = c3
      go to 150
  100 \text{ del1} = \text{del3}
      c1 = c3
  150 continue
      if(abs(c4-c1), lt. epi) go to 250
  200 continue
  250 continue
      c3 = 0.5*(c1+c4)
      if(index.eq.1) cphase(nroot)=omega/c3
      if(nlost.eq.1) cphase(nroot)=omega/c3
      eroot(nroot) = c3
  350 c2 = c1
      del2 = deli
  400 continue
  500 continue
  510 continue
      nlost=nlost+1000
      go to 1250
```

```
3000 continue
      the jumping method.
c
      if this method fails, the control will return
C
      to the regular halving method.
C
      ****
C
      nroot = 0
      if(nroot1.eq.0) go to 1250
      dkO=0. 25*dk
      domega = omegaO-omega
      eroot(nroot1+1)=wvmn
      if(nroot1, eq. mode) eroot(nroot1+1)=eroot(nroot1)-5. *dk
      nlost
            == ()
      do 1200 i=1, nroot1
             = eroot(i)
      vphase = omegaO/s
             = -domega/vphase
      t
            = vphase-cphase(i)
      d c
      tadd = -dc*s/vphase
             = s-eroot(i+1)
      d c
      c1 = dabs(tadd)/tadd
      if(i.eq.1) go to 520
      if(dabs(tadd), lt. dkO) tadd=c1*dkO
       if(dabs(tadd).gt.dc) tadd=c1*O.5*dc
  520 continue
             ≕ 5+ti
       c2
       netrl = 5
       itrig = 0
       if(i,gt,1,and,c2,ge,aroot(i-1)) itrig=1
       if(itrig eq.O) go to 530
       if(dabs(tadd). lt.dkO) tadd=c1*dkO
       If roots are duplicated, a correction might be done here.
 C
       c2 = eropt(i-1)-0.1*abs(dc)
       netrl = 10
   530 continue
             = c2+tadd
       c 1
       ihalf = 20
       ncont = 0
   550 if(c2.le.wvmn) go to 1250
       del2 = dltar(c2,omega,ifunc)
       ntest = 0
   600 if(c1, le. wymn) go to 800
       del1 = dltar(c1, omega, ifunc)
       if(sign(1.0,del1)*sign(1.0,del2),le.0.0) go to 850
       ntest = ntest+1
       if(ntest.ge. nctrl) go to 650
       de12
              ≕ deli
               = c1
       c 2
       c 1
               = c1 + tadd
       go to 600
   650 ncont = ncont+1
        go to (700,720,750), ncont
   700 continue
       This is another kind of jumping procedure, which is
```

```
a remedy when the first jumping method fails.
C
      ihalf = 20
             == -dc/20.d+00
      tadd
             = 5+t
      02
             = 60
      nctrl
      if(i,eq.1) nctrl = 15
      if(itrig. eq. 1) c2=eroot(i-1)-0. 7*dk
      go to 740
  720 continue
      if(itrig.eq.1) go to 750
      if(i, eq. mode) go to 1200
      This is the third kind of jumping procedure.
             = 0.25*t
      tadd
             = s+tadd
      02
      if(i.eq.1) go to 730
      if(c2.1t.eroot(i-1)) go to 730
      tadd = -abs(s+t-eroot(i-1))/4.
      c2 = eroot(i-1)+tadd
  730 continue
      nctrl
            ≕ 3
      ihalf
              = 100
  740 c1
             = c2+tadd
      go to 550
  750 \text{ nlost} = \text{nlost+i}
      If all jumping methods fail, it goes back to the regular
C
      halving method.
       go to 2000
              = wymn+0.01*dk
  800 ci
       deli = dltar(c1, omega, ifunc)
       if(sign(1.0,del1)*sign(1.0,del2).le.0.0) go to 850
       go to 1250
  850 c4
              ≈ ¢2
       del4 = del2
       do 1000 ii=1, ihalf
             = 0.5*(c1+c4)
       del3 = dltar(c3, omega, ifunc)
       if(sign(1.0,del1)*sign(1.0,del3) .ge. 0.0) go to 900
             = del3
       del4
       c 4
              ≖ ୯3
       go to 950
  900 del1
              = del3
              ≖ c3
       c1
  950 continue
       if(abs(c4-c1). lt. epi) go to 1050-
  1000 continue
  1050 continue
       c3 = 0.5*(c1+c4)
       nroot = nroot+1
       eroot(nroot) = c3
       cphase(nroot) = vphase
  1200 continue
  1250 continue
       if(nroot.gt.nroot1) nroot≕nroot1
```

```
nroot1=nroot
      return
      ខ្មាជ
C
C
C
      function dltar(wyno, omega, kk)
      double precision wyno, omega
      common d(500), a(500), b(500), rho(500), mmax, ipunch
      common t(4096), c(500), mode, 11w, twopi
C
      go to (10,20), kk
      love wave period equation
C
   10 dltar = dltar1(wvno.omega)
      return
      rayleigh wave period equation
C
   20 dltar = dltar4(wyno,omega)
      return
       end
C
C
C
       function dltar1(wvno.omega)
       double precision xnor, ynor, wyno, omega
       double precision wyno2, xkb, rb, e1, e2, xmu, q, rho1, beta1
       double precision sinq, cosq, y, z, exqp, exqm, e10, e20
       haskell-thompson love wave formulation from halfspace
C
c
       common d(500), a(500), b(500), rho(500), mmax, ipunch
       common d1(4096), d2(500), mode, 11w, twopi
       wyno2=wyno*wyno
       beta1=dble(b(mmax))
       rhoi=dble(rho(mmax))
       xkb=omega/beta1
       rb = dsqrt(dabs(w∨no2-xkb*xkb))
       e1=rho1*rb
       e2=1. d+00/(beta1*beta1)
       mmm1 = mmax - 1
       do 600 m=mmm1, 11w, -1
       beta1=dble(b(m))
       rhoi=dble(rho(m))
       xmu=rhoi*betai*beta1
       xkb=omega/betal
       rb = dsqrt(dabs(wvno2-xkb*xkb))
       q = dble(d(m))*rb
       if(wyno-xkb) 100,200,300
   100 \sin q = d \sin(q)
       y = sinq/rb
       z = -rb*sinq
       cosq = dcos(q)
       qo to 500
   200 cosq=1.0d+00
        u=dble(d(m))
```

```
z=0.0d+00
      go to 500
  300 continue
      if(q.gt.40.) go to 400
      exqp=dexp(q)
      exqm=1./exqp
      sinq = (exqp-exqm)/2.
      y = sinq/rb
      z = sinq*rb
      cosq = (exqp + exqm)/2.
      go to 500
  400 continue
      y = 0.5d + 00/rb
      z = 0.5d+00*rb
      cosq = 0.5d+00
  500 continue
      e10=e1*cosq+e2*xmu*z
      e20=e1*y/xmu+e2*cosq
      xnor=dabs(e10)
      ynor=dabs(e20)
      if(ynor.gt.xnor) xnor=ynor
      if(xnor.1t.1.d-30) xnor=1.Od+00
      e1=e10/xnor
      e2=e20/xnor
  600 continue
      dltar1=e1
      return
      end
C
      function dltar4(wvno.omga)
      double precision e(5), ee(5), ca(5,5)
      double precision wyno, omga, omega, wyno2, rho1
      double precision xka,xkb,ra,rb,t,gam,gammk,gamm1
      double precision exa, p, q, dpth, w, wO, cosp, cr
      double precision a0, cpcq, cpy, cpz, cqw, cqx, xy, xz, wy, wz
      common d(500), a(500), b(500), rho(500), mmax, ipunch
      common per(4096), c(500), mode, llw, twopi
      common/ovrflw/ a0, cpcq, cpy, cpz, cqw, cqx, xy, xz, wy, wz
      omega≕omga
      if(omega. lt. 1. Od-5) omega=1. Od-5
      w∨no2=w∨no*w∨no
      xka=omega/dble(a(mmax))
      xkb=omega/dble(b(mmax))
      ra≖dsqrt(dabs(w∨no2-xka*xka))
      rb=dsqrt(dabs(w∨no2-xkb*xkb))
      t = dble(b(mmax))/omega
      gammk = 2. *t*t
      gam = gammk*wvno2
      gammi = gam - 1.
      rho1=dble(rho(mmax))
      e(1)=rho1*rho1*(gamm1*gamm1-gam*gammk*ra*rb)
```

```
e(2)=-rho1*ra
      e(3)=rhoi*(gamm1-gammk*ra*rb)
      e(4)=rho1*rb
      e(5)=wvno2-ra*rb
      matrix multiplication from bottom layer upward
C
      mmm1 = mmax-1
      do 500 \text{ m} = \text{mmm1}, 11\text{w}, -1
      xka = omega/dble(a(m))
      xkb = omega/dble(b(m))
      t = dble(b(m))/omega
      gammk = 2.*t*t
      qam = gammk*w∨no2
      ra = dsqrt(dabs(wyno2-xka*xka))
      rb = dsqrt(dabs(wvno2-xkb*xkb))
      dpth=dble(d(m))
      rho1=dble(rho(m))
      p=ra*dpth
      g=rb*dpth
      beta=b(m)
      call var(p,q,ra,rb,wvno,xka,xkb,dpth,w,cosp,exa,beta)
      call dnka(ca, wvno2, gam, gammk, rhoi)
      do 200 i=1,5
      cr=0. Qd+QQ
       do 100 j=1,5
      cr=cr+e(j)*ca(j,i)
  100 continue
       ee(i)=cr
  200 continue
       call normc(ee.exa)
       do 300 i = 1.5
       e(i)=ee(i)
  300 continue
  500 continue
       \omega\Omega = 0.0
       if(11w.eq.1) go to 600
       xka = omega/dble(a(1))
       ra = dsqrt(dabs(wvno2-xka*xka))
      dpth=dble(d(1))
       rhoi=dble(rho(1))
       p ≈ ra*dpth
       beta = b(1)
       call var(p, q, ra, rb, wvno, xka, xkb, dpth, w, cosp, exa, beta)
       wO≕-rho1*w/cosp
  600 continue
       dltar4=e(1)+w0*e(2)
       return
       end
C
C
C
       subroutine var(p, q, ra, rb, wyno, xka, xkb, dpth, w, cosp, exa, beta)
       double precision p.q.ra.rb.wvno.xka.xkb.dpth
```

double precision w.x.y.z.cosp.cosq.sinp.sinq

```
double precision exa, expp, expm, exqp, exqm
    double precision a0, cpcq, cpy, cpz, cqw, cqx, xy, xz, wy, wz
    common/ovrflw/
                      aO, cpcq, cpy, cpz, cqw, cqx, xy, xz, wy, wz
    exa=0.0d+00
    a0≈1.0d+00
    if(wvno-xka) 100,200,300
100 sinp=dsin(p)
    w=sinp/ra
    x=-ra*sinp
    cosp=dcos(p)
    go to 500
200 cosp=1.0d+00
    w=ooth
    x=0. Od+00
    go to 500
300 if(p. gt. 40.0) go to 400
    expp=dexp(p)
    expm=1./expp
    sinp=(expp-expm)*O.5d+OO
    cosp=(expp+expm)*O.5d+OO
    w=sinp/ra
    x=ra*sinp
    go to 500
400 exa=p
    x=ra*0.5d+00
    w=0.5d+00/ra
    cosp=0.5d+00
    a0=0.0d+00
    if(exa. lt. 70.0) a0=1. /dexp(exa).
500 continue
    if(beta, lt. 1, e-5) return
    if(wvno-xkb) 600,700,800
600 sinq=dsin(q)
    z=-rb*sing
    y=sing/rb
    cosq=dcos(q)
    go to 1000
700 cosq=1.0d+00
    y=dpth
    z=0.0d+00
    go to 1000
800 if(q.gt.50.0) go to 900
    exqp=dexp(q)
    exqm=1./exqp
    sinq=(exqp-exqm)*0.5d+00
    cosq=(exqp+exqm)*O.5d+OO
    y=sinq/rb
    z=rb*sing
    go to 1000
900 y=0.5d+00/rb
    z=0.5d+00*rb
    cosq=0.5d+00
    p+exe=exe
```

```
a0=0.0d+00
      if(exa.lt.70.0) a0=1./dexp(exa)
1000 continue
      cpcq=cosp*cosq
      cpy=cosp*y
      cpz=cosp*z
      U¥X≔UX
      x = x + z
      աս≕ա×ս
      WIZ₩W¥I
      cqw≕cosq≯w
      cqx=cosq*x
      return
      end
C
C
c
      subroutine normc(ee,ex)
      This routine is an important step to control over- or
C
      underflow.
C
      The Haskell or Dunkin vectors are normalized before
C
      the layer matrix stacking.
c
      Note that some precision will be lost during normalization.
C
C
      double precision ee(5), ex, t1, t2
      ex = 0.0d+00
      t1 = 0.0d+00
      do 10 i = 1.5
      if(dabs(ee(i)), gt, t1) t1 = dabs(ee(i))
   10 continue
      if(t1, 1t, 1, d-30) t1=1, d+00
      do 20 i = 1.5
      t2=ee(i)
      t2=t2/t1
      ee(i)=t2
   20 continue
      ex=dlog(t1)
      return
      end
C
C
C
      subroutine dnka(ca, wvno2, gam, gammk, rho)
       double precision ca(5,5), wyno2, gam, gammk, rho
       double precision gamm1, twgm1, gmgmk, gmgm1, gm1sq, rho2, t
       double precision a0, cpcq, cpy, cpz, cqw, cqx, xy, xz, wy, wz, a0pq
       common/ ovrflw / aO, cpcq, cpy, cpz, cqw, cqx, xy, xz, wy, wz
       gamm1 = gam-1.
       twom1=qam+qamm1
       qmqmk=qam*qammk
       gmgm1=gam*gamm1
       gmisq=gammi*gammi
       rho2=rho*rho
```

```
aOpq=aO-cpcq
     ca(1,1)=cpcq-2.*gmgm1*aOpq-gmgmk*xx-wvno2*gm1sq*wy
     ca(1,2)=(wvno2*cpy-cqx)/rho
     ca(1,3)=-(twgm1*aOpq+gammk*xz+wvno2*gamm1*wy)/rho
     ca(1,4)=(cpz-wvno2*cqw)/rho
      ca(1,5)=-(2.*wvno2*aOpq+xz+wvno2*wvno2*wy)/rho2
      ca(2,1)=(qmqmk*cpz-qm1sq*cqw)*rho
      ca(2,2)=cpcq
      ca(2,3)=qammk*cpz-gamm1*cqw
      ca(2,4)=-wz
      ca(2,5) = ca(1,4)
      ca(4,1)=(gm1sq*cpy-gmgmk*cqx)*rho
      ca(4,2)=-xu
      ca(4,3)=qamm1*cpy-gammk*cqx
      ca(4,4)=ca(2,2)
      ca(4,5) = ca(1,2)
      ca(5,1)=-(2,*gmgmk*gm1sq*aOpq+gmgmk*gmgmk*xz+
                gmisq*gmisq*wy)*rho2
      ca(5,2)=ca(4,1)
      ca(5,3)=-(ganmk*gamm1*twgm1*aOpq+gam*gammk*gammk*xz+
                qamm1*qm1sq*wy)*rho
      ca(5,4)=ca(2,1)
      ca(5,5) = ca(1,1)
      t=-2, *wvno2
      ca(3,1)=t*ca(5,3)
      ca(3,2)=t*ca(4,3)
      ca(3,3) = a0+2. *(cpcq-ca(1,1))
      ca(3,4)=t*ca(2,3)
      ca(3,5)=t*ca(1,3)
      return
      end
C ******
```

```
dpsrf81.f
C
C
      f77 -i -12 dpsrf.f -o dpsrf -lcalcomp12
C
C
      plot dispersion curve
C
      get data from surface81
C
r
      dimension t(2048), v(2048), mode(500)
      dimension d(500),a(500),b(500),rho(500)
      character*20 names
      character*4 yorn
      format(a)
  5
      format(4f12.5)
  15
      pi=2. *3. 141592653
      write(6,*) ' '
      write(6,*) 'Rayleigh(1) or Love(2): '
      read(5, *) 11rr
      write(6,*) ' '
      write(6,*) 'enter the input file name (from surface81):'
      read(5,5) names
      open(1, file=names, status='old', form='unformatted')
      write(6, *)
          'how many modes and which modes wanted? (3, 1, 3, 5)'
      write(6,*) '(if all modes wanted, answer 1,0 )'
       read(5,*) nmode,(mode(i),i=1,nmode)
       if(mode(1), ne. 0) go to 40
       nmode=500
       do 30 i=1,500
      mode(i)=i
  30
      continue
      continue
  40
                                     C--freq(2) K-C(3)
                                                            C-K(4)'
                          C-perd(1)
       write(6,*) 'plot
                                     freq-K(6): '
       write(6,*) '
                          K-freq(5)
       read(5, *) ipg
       call plots(0,0,7)
       write(6,*) 'enter ipen: '
       read(5,*) ipen
       call newpen(ipen)
       yorn='y'
       call plot(2, 5, 2, 1, -3)
       write(6,*) ' '
       write(6,*) 'model:'
       do 600 kk=1, nmode
       rewind 1
       read(1) nmax
       do 50 ii=1, nmax
       read(1) d(ii), a(ii), b(ii), rho(ii)
       continue
       if(kk, eq. 1) write(6, 15) (d(i), a(i), b(i), rho(i), i=1, nmax)
       read(1) nper
   100 continue
       ident=mode(kk)
       nt=0
```

```
itrig=0
 200 continue
     read(1) ifunc, kmode, t1
      if(ifunc. 1t. 0) go to 400
      if(kmode. le. 0) go to 200
      nt=nt+1
      do 300 i=1, kmode
      read(1) cO
      wyne0=2, *3, 141592653/(c0*t1)
      if(i.ne.ident) go to 300
      itrig=1
      nn=nt
      t(nn)=t1
      v(nn)=c0
 300 continue
      go to 200
 400 continue
      if(itrig. eq. O. and. nmode. eq. 500) go to 600
      do 500 i=1, nn
      go to (500, 482, 487, 488, 485, 486), ipg
 482 t(i)=1./t(i)
      go to 500
 485 t(i)=1./t(i)
      \forall(i)=pi*t(i)/\forall(i)
      go to 500
 486 tmp=1./t(i)
      t(i) #pi*tmp/v(i)
      ∨(i)=tmp
      go to 500
 487 tmp=v(i)
      \forall(i)=pi/(t(i)*\forall(i))
      t(i) = tmp
      go to 500
  488 t(i)=pi/(t(i)*v(i))
  500 continue
      call disp(yorn, itrig, kk, ipg, ident, nn, t, v)
      call plot(0.0,8.0,999)
      close(1)
      close(2)
      write(6,*) 'job finished'
      stop
      end
C
      subroutine disp(yorn, itrig, kk, ipg, ident, nn, x, y)
      dimension x(1), y(1)
      character*4 yorn, nory
      character*11 alpha(2,6),alp,bet
      data alpha/
                                      ', 'C
     讲
         ľC
            (km/sec)', 'T
                             (sec)
                                             (km/sec)','f
                                                            (Hz)
                             (km/sec)','C
                                             (km/sec)','K
                                                            (rad/km)',
            (rad/km)', 'C
         7 K
                                      ', 'f
                                             (Hz)
                                                      ', 'K (rad/km)' /
       'K (rad/km)', 'f
                             (Hz)
      if(itrig.eq.O)
```

```
write(6,*) 'mode ', ident, ' is not generated. '
   if(itrig.eq.O) return
   alp=alpha(1, ipg)
   bet=alpha(2, ipg)
   if(yorn.eq. 'n') go to 100
   uorn='n'
   xlen=6.0
   ulen=4.5
    xmin=1.e+37
    x = -1. e + 37
   umin=1. e+37
    ymax=-1.e+37
    do 80 i=1, nn
    if(x(i), ge, xmax) xmax=x(i)
    if(x(i), le, xmin) xmin=x(i)
    if(y(i).ge.ymax) ymax=y(i)
    if(y(i), le.ymin) ymin=y(i)
80
   continue
    write(6,*) 'ymin=',ymin,' ymax=',ymax
    if(ipg.ge.9) write(6.*) ' -y value in log sacle'
    write(6,*) 'xmin=',xmin,' xmax=',xmax
    write(6,*) '
    write(6,*) 'enter ymin,ymax,yinc,xmin,xmax,xinc'
                                    -y should be integer. '
    if(ipg.ge.9) write(6,*) '
    read(5,*) yO,y1,yinc,x0,x1,xinc
    xinut=(x1-x0)/x1en
    qinut=(g1-g0)/glen
100 continue
    if(kk.ne.1) go to 500
    write(6,*) 'plot axis? (y/n)'
    read(5,5) nory
5
    format(a)
    write(6,*) ' '
    write(6,*) 'wait. It is processing.'
    if(nory.ne.'y') go to 500
    call plot(xlen, 0, 0, 2)
    call plot(xlen,ylen,2)
    call plot(0.0,ylen,2)
    call plot(0.0,0.0,2)
    call plot(0.0,-0.05,2)
    if(x0.1t.0.0) xshif=-0.25
     if(x0, ge, 0, 0) xshif=-0.13
     if(x0, ge, 10, 0) xshif=-0.23
    if(xinc.ge. 0.1) call number(xshif, -0.17, 0.1, x0, 0.0, 1)
     if(xinc. lt. O. 1)
   *call number(xshif-0.04,-0.16,0.09,x0,0.0,2)
     orid=xinc/xinut
     xi=1.
200 xx = xi*grid
     if(abs(xx).gt.xlen+0.02) go to 250
     call plot(xx, 0.0,3)
     call plot(xx,-0.05,2)
     xsum=xO+xinc*xi
```

```
if(xsym. 1t. 0. 0) xshif=-0.25
    if(xsym.ge.O.O) xshif=-0.13
    if(xsym.ge.10.0) xshif=-0.23
    if(xsym.ge. 100.0) xshif=-0.33
    if(xinc. qe. 0.1)
   *call number(xx+xshif, -0.17, 0.1, xsym, 0.0, 1)
    if(xinc. lt. 0. 1)
   *call number(xx+xshif-0.04,-0.16,0.09,xsym,0.0,2)
    xi = xi+1.
    go to 200
250 continue
    call symbol(2, 5, -0, 4, 0, 16, bet, 0, 0, 11)
    call plot(0,0,0,0,3)
    call plot(-0.05,0.0,2)
    if(ipg.ge.9) go to 260
    if(y0.1t.0.0) xshif=-0.45
    if(y0, ge, 0, 0) xshif=-0.35
    if(y0.ge.10.0) xshif=-0.45
    if(qinc.ge. 0.1)
   *call number(xshif, -0.05, 0.1, y0, 0.0, 1)
    if(qinc. 1t. 0. 1)
   *call number(xshif-0.03,-0.05,0.09,y0,0.0,2)
    go to 270
260 continue
    call number(-0.34,-0.06,0.1,10.0,0.0,-1)
    call number (-0. 16, 0. 02, 0. 06, y0, 0. 0, -1)
270 continue
    arid=yinc/yinut
     xi=1.
300 xx=xi*grid
     if(abs(xx), gt, ylen+0, 02) go to 400
    call plot(O.O,xx,3)
     call plot(-0.05, xx, 2)
     call plot(O.O,xx,3)
     xsym=yO+yinc*xi
     if(ipg.ge.9) go to 310
     if(xsym. lt. 0. 0) xshif=-0. 45
     if(xsym.ge, 0.0) xshif=-0.35
     if(xsym.ge. 10.0) xshif=-0.45
     if(xsym.ge. 100. 0) xshif=-0.55
     if(yinc, ge. 0.1)
    *call number(xshif,xx-0.05,0.1,xsym,0.0,1)
     if(uinc. 1t. 0. 1)
    *call number(xshif-0.04, xx-0.05, 0.09, xsym, 0.0, 2)
     go to 320
310 continue
     call number (-0.34, xx-0.06, 0.1, 10.0, 0.0, -1)
     call number (-0.16, xx+0.02, 0.06, xsym, 0.0, -1)
320 continue
     xi = xi+1.
     go to 300
400 continue
     call symbol(-0.55,1.5,0.16,alp,90.0,11)
```

```
reigen81.f
      f77 -i -I2 reigen81.f -o reigen81
C
      This program calculates the eigenfunctions of
C
      Rayleigh wave for any plane multi-layered model.
C
c
      Body wave Q model can be included.
C
C.
      The propagator-matrix, instead of numerical-integration
C
      method is used, in which the Haskell rather than
C
C
      Harkrider formalisms are concered.
C
      Such a revision was developed to cover a large range of
C
      frequencies, say 200 Hz, and to improve the calculation
C
      efficiency. The layer thickness is not limitted.
C
C
      For the sake of space saving, only the values of
C
       eigenfunctions at the source depths are stored.
C
       However, several files with different source depths
C,
C
       can be set up.
C
 C
       -Oct 10, 1981
 C
       double precision sumiO, sumi1, sumi2, sumi3
       dimension nos(100), dphs(100), dphq(100), qa(100), qb(100)
       dimension depth(100)
                       d(100), a(100), b(100), rho(100), qa1(100),
       common/model/
                         qb1(100), xmu(100), xlam(100), mmax, ll
       common/eigfun/ ur(100), uz(100), tz(100), tr(100), uu0(4),
                          dcda(100), dcdb(100), dcdr(100)
                        sumiO, sumi1, sumi2, sumi3, flagr, are, ugr
      *
        common/sumi/
        character*1 dd
        character*50 names
        format(a)
        format(/2x,'M',4x,'DPTH',2x,' D',3x,' A',3x,' B',
  5
                 3x, ' RHO', 3x, ' QA ', 4x, ' QB ', 4x, ' MU ', 3x, 'LMDA')
  10
        format(i3,1x,5f7,2,2f8,2,2f7,2)
        format(3x, '-source is on the top of this layer. ',
  20
   30
                  / source depth=',6.2)
        format(i3, 1x, f7, 2, 7x, 3f7, 2, 2f8, 2, 2f7, 2)
   40
        format(3x, '-source is inside the half space.',
   50
                  / source depth=',f6.2)
        form at('at the source depth = ', f6.2, '
   60
        write(6, *)
            'enter the input file name: (from surface81)'
        read(5,5) names
        open(1, file=names, status='old', form='unformatted')
        open(3, file='tmp. 1', status='scratch', form='unformatted')
        rewind 3
  C.
        enter source depths and Q-model:
```

C C

C

```
C
      write(6,*) 'enter the source depths: '
      write(6, *)
         '(no. of source depths, sredph(1), sredph(2), ...)'
      write(6,*)
         '*** if no of source depths is negative, no output'
      write(6,*) / files will be generated. ***
      This can be used as dividing layers.
C
      read(5, \Rightarrow) kks, (dphs(i), i=1, iabs(kks))
      ks=abs(kks)
      write(6,*) 'enter Q-model here(1), from a file(2), '
      write(6,*) 'or NOT considered(3): '
      read(5, *) kk
      kg=1
      go to (100,110,170), kk
      continue
 100
      write(6,*)
          'enter d(i), Qa(i), Qb(i) use d(i)=0.0 for halfspace'
      icode=5
      go to 120
      continue
 110
      write(6,*) 'enter the name of the file storing Q-model:'
      read(5,5) names
      open(2, file=names, status='old', form='formatted')
       rewind 2
       icode=2
 120 continue
       i = 1
       base=0.0
 140
       continue
       read(icode, *) dq(), qa(i), qb(i)
       if(dq0. le. 0. 0) go to 150
       base=base+dqO
       dphq(i)=base
       i = i + 1
       go to 140
  150
       kq≖i
       do 160 i=1,100
       qa1(i)=qa(kq)
       qb1(i)=qb(kq)
       continue
  160
  170
       continue
       write(∆,*) 'Store the derivatives? (y/n):'
       read(5,5) dd
       if(dd.eq. 'n') go to 175
       write(6,*) 'enter the output file name for derivatives: '
       read(5,5) names
       open(9, file=names, status='new', form='unformatted')
       rewind 9
       continue
  175
       do 180 i=1,100
       depth(i)=10000000.0
  180 continue
```

```
C
      obtain the earth model:
C
Ċ.
      read(1) mmax
      write(6,10)
      base≕O.O
      depth(1)=0.0
      do 190 i=1, mmax
      read(1) d(i),a(i),b(i),rho(i)
      base=base+d(i)
       depth(i+1)=base
       xmu(i)=rho(i)*b(i)*b(i)
       xlam(i)=rho(i)*(a(i)*a(i)-2.*b(i)*b(i))
       continue
 190
C
       insert the G-model into the velocity model.
C
       insert the source depth as an interface of layers.
C
 C
       kas=ka+ks-1
       do 300 k0=1, kqs
       dphO=dphq(kO)
       is=kO-kq+1
       if(kO.ge.kq) dphO=dphs(is)
       do 200 i=1, mmax
       k ≋ i
       if(dphO.eq.depth(i)) go to 250
       if(dphO.gt.depth(i).and.dphO.lt.depth(i+1)) go to 210
       continue
  500
       k1 = k+1
  210
        dphk1=depth(k1)
        do 220 i=mmax, k,-1
        i1=i+1
        d(i1)=d(i)
        a(i1)=a(i)
        b(i1)=b(i)
        rho(i1)=rho(i)
        xmu(i1)=xmu(i)
        xlam(i1)=xlam(i)
        depth(ii)=depth(i)
        if(kO, ge, kq) qal(i1)=qal(i)
        if(kO, ge, kq) qb1(i1)=qb1(i)
        continue
   220
        d(k)=dphO-depth(k)
        d(k1)=dphk1-dphO
        depth(k1)=depth(k)+d(k)
        mmax≕mmax+1
        if(kO.ge.kq) go to 240
        if(k0, eq. 1) ns=1
        do 230 j=ns/k
        qa1(j)=qa(k0)
        qb1(j)=qb(kO)
        continue
   230
        ns=k1
   240
```

```
go to 280
250
                continue
                if(kO.ge.kq) go to 270
                 if(kO.eq.1) ns=1
                 do 260 i=ns/k-1
                 qa1(i)=qa(kO)
                 qb1(i)=qb(kO)
                 continue
260
                 ns≕k
270
                 nos(is)=ns
280
 300
                 continue
                  1=1
                 do 310 i=1, mmax-1
                 write(6,20) i,depth(i),d(i),a(i),b(i),rho(i),qa1(i),
                                                                qb1(i), xmu(i), xlam(i)
                  if(i.ne.nos(j)) go to 310
                 write(6,30) dphs(j)
                  J=1+1
 310
                  continue
                  i = mmax
                  write(6,40)
                         i, depth(i), a(i), b(i), rho(i), qa1(i), qb1(i), xmu(i), xlam(i)
                  if(nos(j), eq. mmax) write(6,50) dphs(j)
                   if(dd.eq.'y')
                            write(9) mmax, (d(i), a(i), b(i), rho(i), i=1, mmax)
                  write(6,*) ' '
                   write(6,*) 'wait.'
                   if(b(1), 1e, 0, 0) 11=2
                   read(1) nper
   400
                   continue
C
                   read in the dispersion values.
C
C
                    read(1) ifunc, mode, t
                    write(3) ifunc, mode, t
                    if(dd.eq.'y') write(9) ifunc, mode, t
                    if(ifunc. 1t. 0) go to 700
                    if (mode, le, O) go to 400
                    do 600 k=1, mode
                    read(1) c
 C
                    main part.
 C
 C
                     omega=6.2831853/t
                     wyno=omega/c
                     call syfunc(omega, wyno)
                     call energy(omega, wyno)
                     omega2=omega*omega
                     machine series and machine and
                     do 450 i=11, mmax
                     ur(i)=ur(i)*w∨no
                      tz(i)=tz(i)*omega2
```

```
fr(i)=tr(i)*wvomg2
450
      continue
      if(kk.ne.3) call gammap(omega.wvno.gamma)
      if(dd.eq.'n') go to 510
      output the derivatives.
C
      xiO=sumiO
      xil=sumi1
      xi2=sumi2
      xi3=sumi3
      write(9) 000(1),000(3),c,ugr,xi0,xi1,xi2,xi3,are,flagr
      do 500 i=1, mmax
      write(9) depth(i), ur(i), uz(i), tz(i), tr(i),
                  dcda(i), dcdb(i), dcdr(i)
 500
      continue
 510
      continue
      do 560 i=1,ks
      j=nos(i)
      urs=ur(j)
      uzs=uz(j)
      durs=-wvho*uzs+tr(j)/xmu(j)
      durs=(wvno*xlam(j)*urs+tr(j))/(xlam(j)+2.*xmu(j))
      ur0=ur(11)
C
      output
C
C
      write(3) wyno,urO,are,ugr,gamma
      write(3) urs, durs, uzs, duzs
 560
       continue
       continue
 600
       go to 400
 700
      continue
       output the eigenfunction files for different source depths.
C
C
       if(kks.le.0) go to 950
       do 900 i=1,ks
       rewind 3
       write(6, *)
         'enter the name of output file storing the eigenfunctions'
       write(6,60) dphs(i)
       read(5,5) names
       open(4, file=names, status='new', form='unformatted')
       rewind 4
       write(4)
          mmax, (d(j), a(j), b(j), rho(j), qa1(j), qb1(j), j=1, mmax)
       write(4) nper,dphs(i)
  800
       continue
       read(3) ifunc, mode, t
       write(4) ifunc, mode, t
       if(ifunc. 1t. 0) go to 870
       if (mode, le. 0) go to 800
       do 860 k=1, mode
       do 850 j=1, ks
```

```
read(3) qa(j),qb(j),dphq(j),depth(j),dcda(j)
      read(3) ur(j), tr(j), uz(j), tz(j)
850
      continue
      write(4) qa(i), qb(i), dphq(i), depth(i), dcda(i)
      urite(4) ur(i), tr(i), uz(i), tz(i)
860
      continue
      go to 800
 870
      continue
      close(4)
 900
      continue
 950
      continue
      close(1)
      close(2)
      close(3, status≕'delete')
      write(6, *) ' '
      write(6,*) 'reigen81 finished'
      write(6,*) ' '
      stop
      end
C
C
C
      subroutine gammap(omega,wyno,gamma)
      This routine finds the attenuation gamma value of
C
C
      surface wave.
C
                       d(100), a(100), b(100), rho(100), qa1(100),
      common/model/
                         qb1(100), xmu(100), xlam(100), mmax, ll
      common/eigfun/ ur(100), uz(100), tz(100), tr(100), uuO(4),
                         dcda(100), dcdb(100), dcdr(100)
      x=0.0
      do 100 i=11, mmax
      x=x+dcda(i)*a(i)/qa1(i)+dcdb(i)*b(i)/qb1(i)
 100
      continue
      c≕omega/w∨no
      gamma=0.5*w∨no*x/c
      return
       end
C
C
C,
       subroutine syfunc (omega, wyno)
      This routine combines the Haskell vector from sub down and
C
C
      Dunkin vector from sub up to form the eigenfunctions.
C
       double precision exe(100), exa(100), ext, fact
                       d(100), a(100), b(100), rho(100), qa1(100),
       common/madel/
                         gb1(100), xmu(100), xlam(100), mmax, ll
       common/eigfun/ ur(100), uz(100), tz(100), tr(100), uuO(4),
                         dcda(100), dcdb(100), dcdr(100)
                       υυ(100,5), exe, exa
       common/dunk/
       common/hask/
                       \vee \vee (100, 4)
       common/water/ waterO,ktrig
```

```
ktrig=1
      call up (omega, wyno, fr)
      ktrig=0
      call down(omega, wyno)
      のヘン・ログドラインのネートファ
      omega2≕omega*omega
      f5=uu(11,4)
      uuO(1)=wvno*uu(11,3)/f5
      000(2)=1.0
C
      υυΟ(3)=tz is actually the period equation.
      uuO(3)≕fr
C
      uuO(4)=tr should be zero.
      UUQ(4)=fr
      ur(11)=uu(11,3)/f5
      uz(11)=1.0
      tr(11)=-waterO
      tr(11)=0.0
      do 200 i=11+1/mmax
      i1 = i - 1
      սս1 ≕
     * \v(i1,2)*\uu(i,1)+\v(i1,3)*\uu(i,2)
                                                   +vv(i1,4)*uu(i,3)
     * --vv(i1,1)*uu(i,1)-vv(i1,3)*uu(i,3)*wvno2+vv(i1,4)*uu(i,4)
      ∪∪3 ==
     * -vv(i1,1)*uu(i,2)+vv(i1,2)*uu(i,3)*wvno2+vv(i1,4)*uu(i,5)
      UU4 =
     * -vv(i1,1)*vu(i,3)--vv(i1,2)*vu(i,4)
                                                  - \lor \lor (i1,3) * \upsilon \upsilon (i,5)
      ext=0.0
      do 100 k=11, i1
      ext=ext+exa(k)-exe(k)
 100
     continue
      fact=0.0
      if(ext.gt.-80.0) fact=dexp(ext)
      ur(i)=uu1*fact/f5
      U1(i)=UU2*fact/f5
      tr(i)=uu3*fact/f5
      tr(i)=004*fact/f5
 200
      continue
      return
      end
C
C
C
      subroutine up(omega, wvno, fr)
      This routine finds the values of the Dunkin vectors at
C
C
        each layer boundaries from bottom layer upward.
C
      dimension eeO(5)
      double precision exe(100), exa(100), ex1, ex2, p, q, rab, dept
      double precision wd.wra.wcosd2.wsin2d
      common/model/ d(100),a(100),b(100),rho(100),qa1(100),
                         qb1(100), xmu(100), xlam(100), mmax, 11
      common/dunk/
                      uu(100,5), exe, exa
```

```
common/save/
                      dd(5, 5), aa(4, 4), exi, ex2
      common/aamatx/ ww(100), xx(100), yy(100), zz(100),
                       cospp(100); cosqq(100)
      common/water/ waterO,ktrig
      common/engerw/ wd, wra, wcosd2, wsin2d
      onvw*onvw=Sonvw
      xka=omega/a(mmax)
      xkb=omega/b(mmax)
      ramsgrt(abs(wyno2-xka*xka))
      rb=sqrt(abs(wvno2-xkb*xkb))
      t == b(mmax)/omega
      gammk = 2. *t*t
      gam = gammk*w∨no2
      gamm1 = gam - 1.
      if(wvno.lt.xka) write(6,*) ' imaginary nua'
      if(wyno.lt.xkb) write(6,*) ' imaginary nub'
      uu(mmax,1)=wvno2-ra*rb
      uu(mmax,2)=-rho(mmax)*rb
      uu(mmax,3)=rho(mmax)*(gamm1-gammk*ra*rb)
      uu(mmax, 4)=rho(mmax)*ra
      υυ(mmax,5)=rho(mmax)*rho(mmax)*(gamm1*gamm1-gam*gammk*ra*rb)
      matrix multiplication from bottom layer upward
C
      mm \times 1 = mma \times -1
      do 400 k=mmx1,11,-1
      k1 = k + 1
      dpth=d(k)
      xka = omega/a(k)
      xkb = omega/b(k)
      t ≕ b(k)/omega
      gammk = 2. *t*t
      gam = qammk*w∨no2
      ramabs(wyno2-xka*xka)
      rab=dble(ra)
      rab=dsqrt(rab)
      dept=dble(dpth)
      p≕rab*dept
      ra=sngl(rab)
      rb=abs(wvno2-xkb*xkb)
      rab=dble(rb)
      rab=dsqrt(rab)
      q≕rab*dept
     ·rb≕snql(rab)
      call var(k,p,q,ra,rb,wvno,xka,xkb,dpth)
      call dnka(wvno2; gam, gammk, rho(k))
      exa(k)=ex2
      do 200 i=1,5
      cc=0.0
      do 100 j=1,5
      cc=cc+dd(i,j)*uu(k1,j)
  100 continue
      eeO(i)=cc
  200 continue
      call normc(eeO, rab)
```

```
exe(k)=ex1+rab
      do 300 i=1,5
      uu(k,i) = eeO(i)
  300 continue
  400 continue
      water0≕0.0
      if(11, eq. 1) go to 500
      xka=omega/a(1)
      ra=abs(w∨no2-xka*xka)
      rab=dble(ra)
      rab=dsqrt(rab)
      wra≕rab
      dept=dble(d(1))
      p=rab*dept
      ra=sngl(rab)
      call var(100, p. O. O. ra. O. O. wvno, xka, O. O. d(1))
      waterO describes the surface water layer effect.
C
      water0=rho(1)*ww(100)/cospp(100)
      if(ktrig.eq.O) go to 500
C
      prepare for subroutine wenerg which takes the energy
C
      integral over water layer.
      q=2, *ex2
      rab=0.0
      if (q, gt, -80, 0, and, q, lt, 80, 0) rab=1. /dexp(q)
      wd=rab*dept
      wcosd2=2. *cospp(100)*cospp(100)
      p = 2. *p
      ra=2. *wra
      call var(100, p. O. O, ra. O. O, wvno, xka, O. O, d(1))
      g≕ex2-g
      rab=0.0
      if(q, qt, -80.) rab=dexp(q)
      wsin2d=rab*ww(100)
      if(wvno.lt,xka) wra=-wra
  500 continue
      fr=uu(11,5)+waterO*uu(11,4)
      return
      end
C
C
C
      subroutine down(omega, wyno)
      This routine finds the values of the Haskell vectors at
C
C
      each layer boundaries from top layer downward.
C
      dimension aaO(5)
      double precision exe(100), exa(100), ex1, ex2
      common/model/ d(100), a(100), b(100), rho(100), ga1(100),
                         qb1(100), xmu(100), xlam(100), mmax, ll
      common/dunk/
                      uu(100,5), exe, exa
                      vv(100,4)
      common/hask/
      common/save/
                      dd(5,5), aa(4,4), exi, ex2
      common/aamatx/ ww(100), xx(100), yy(100), 2z(100),
```

```
cospp(100), cosqq(100)
      wyno2=wyno*wyno
      do 100 j=1,4
      \vee \vee (11, 1) = 0.0
100
      continue
      \vee\vee(11,4)=1.0
      aa0(5)=0.0
      mm \times 1 = mma \times -1
      do 500 k=11, mmx1
      k1 = k-1
      if(k, eq. 11) k1=11
      t=b(k)/omega
      qammk=2. *t*t
      gam=gammk*wvno2
      ա≕աա(k)
      x=xx(k)
      ų≕yų(k)
      2== 2 2 ( k )
      cosp=cospp(k)
      cosa=cosaa(k)
      call hska(w.x.y.z.cosp.cosq.wvno2.gam.gammk.rho(k))
      do 300 j=1,4
      ¢ c = 0. 0
      do 200 i=1.4
      cc=cc+vv(k1,i)*aa(i,j)
200
      continue
      aaQ(j)=cc
300
      continue
      call normc(aa0,ex2)
      exa(k)=exa(k)+ex2
      do 400 i=1.4
      \vee\vee(k,i)=aaO(i)
 400
      continue
 500
      continue
      return
      end
С
C
      subroutine var(m, pp, qq, ra, rb, wvno, xka, xkb, dpth)
      This routine calculates the values of exp(p), exp(q)...
C
      Since exp(88) \cong 10.**38, the exponential power terms p,q
C
      must be controlled to prevent possible overflow.
C
c
      double precision exa, ex, pp, qq, qpp, qmp
     , double precision expp, expm, exqp, exqm, sinp, cosp, sinq, cosq
                        dd(5,5),aa(4,4),exa,ex
      common/save/
       common/aamatx/ w0(100),x0(100),y0(100),z0(100),
                      cospO(100), cosqO(100)
       common/ovrflw/ a0, cpcq, cpy, cpz, cqw, cqx, xy, xz, wy, wz
       e x ≈ 0. 0
       a0=1.0
       if(wyno-xka) 100,200,300
```

```
100 sinp=dsin(pp)
    w=sinp/dble(ra)
    x=-dble(ra)*sinp
    cosp=dcos(pp)
    go to 500
200 \text{ cosp} = 1.0d + 0
    w≕dpth
    x = 0, 0
    go to 500
300 if(pp.gt.40.0) go to 400
    expp=dexp(pp)
    expm=1./expp
    sinp=(expp-expm)*0.5
    cosp=(expp+expm)*0.5
    w=sinp/dble(ra)
    x=dble(ra)*sinp
    go to 500
400 ex=pp
    x=ra#0.5
    w=0.5/ra
    cosp≈0.5
    a0=0.0
    if(ex. 1t. 75. 0) a0=1. /dexp(ex)
500 continue
    if(m. eq. 100) go to 1100
    if(wyno-xkb) 600,700,800
600 sinq=dsin(qq)
    z=-dble(rb)*sinq
    y=sing/dble(rb)
    cosq=dcos(qq)
    go to 900
700 cosq=1.0d+0
    y=dpth
    z =0. O
    go to 900
800 if(qq.gt.40.0) go to 1000
    exqp=dexp(qq)
    exqm=1./exqp
    sing=(exap-exam)*0.5
    cosq=(exqp+exqm)*0.5
    y≕sinq/dble(rb)
    z=dble(rb)*sing
900 cpcq=cosp*cosq
    cpy=cosp*y
    cpz=cosp*z
    cqw=cosq*w
    cqx=cosq*x
    Xy≠x∻y
    X 2 == X == Z
    աս≕ա∗ս
    wz≕w*z
    cosq≕aO∻cosq
    z = a O * z
```

```
u≕aO⊁u
      e x a = e x
      go to 1100
1000 qpp=qq+pp
      qmp=qq-pp
      exqp=0.0
      if(abs(qmp), lt, 60.) exqp=dexp(qmp)
      exqm=0.0
      if(qpp. lt. 60.) exqm=1./dexp(qpp)
      sinq=(exqp-exqm)*0.5
      cosq=(exqp+exqm)*O.5
      y=sinq/dble(rb)
      z=dble(rb)*sing
      22=rb*0.5
      yy=0.5/rb
      ccosq#0.5
      еха≕ех⊹а्а
      a0=0.0
      if(exa. 1t. 75. 0) a0=1. /dexp(exa)
      cpcq=cosp*ccosq
      cpy=cosp*yy
      cpz=cosp*zz
      xu=x*uu
      XZ=X*ZZ
      ហប្≕ឃ¥ប្ប
      wz=w*zz
      cqw=ccosq*w
      cqy=ccosq*x
 1100 continue
      ωO(m)≔w
      x = (m) \circlearrowleft x
      gO(m)=g
      zO(m)=z
      cospO(m)=cosp
      cosqO(m)=cosq
      return
      end
C
C
C
      subroutine normc(ee,ex)
      This routine is an important step to control over- or
C
C
      The Haskell or Dunkin vectors are normalized before
C
C
      the layer matrix stacking.
C,
      Note that some precision will be lost during normalization.
C
      dimension ee(5)
      double precision ex, t1, t2
      t0 = 0.0
      do 10 i = 1.5
      if(abs(ee(i)), gt, tO) tO = abs(ee(i))
   10 continue
```

```
ti=dble(tO)
       if(t1, 1t, 1, d-30) t1=1, d+00
       do 20 i = 1.5
       t2=dble(ee(i))
       t2=t2/t1
       ee(i) = sngl(t2)
    20 continue
       ex=dlog(t1)
       return
       end
C
C
C
       subroutine hska(w,x,y,z,cosp,cosq,wvno2,gam,gammk,rho)
       double precision ex1, ex2
       common/save/ dd(5,5),aa(4,4),ex1,ex2
       qamm1 = qam-1.
       temp = x-w vno2*y
       aa(2/3) = temp/rho
       temp = temp*gammk
       aa(4,3) = temp+y
       aa(4,1) = -(gamm1*y+gam*aa(4,3))*rho
       aa(2,1) = -wvno2*aa(4,3)
       temp = cosq-cosp
       aa(1,3) = temp/rho
       aa(2,4) = -wvno2*aa(1,3)
       aa(4,2) = rho*qammk*qamm1*temp
       aa(3,1) = -wvno2*aa(4,2)
       temp = temp*gam
       aa(1,1) = cosq-temp
       aa(2,2) = cosp+temp
       aa(3,3) = aa(2,2)
       aa(4,4) = aa(1,1)
       temp = z-wyno2*w
       aa(1,4) = temp/rho
       aa(1/2) = -w-gammk*temp
       aa(3,4) = -wvno2*aa(1,2)
       aa(3,2) = rho*(gam*aa(1,2)-gamm1*w)
       return
       end
C
C
       subroutine dnka(wvno2, gam, gammk, rho)
       double precision ex1, ex2
       common/save/ dd(5,5),aa(4,4),ex1,ex2
       common/ovrflw/ a0, cpcq, cpy, cpz, cqw, cqx, xy, xz, wy, wz
       gamm1 = gam-1.
       twgm1=gam+gamm1
       gmgmk=gam*gammk
       gmgm1=qam*qamm1
       gm1sq=gamm1*gamm1
       rho2=rho*rho
```

```
aOpq=aO-cpcq
      dd(1,1)=cpcq-2.*gmgm1*aOpq-gmgmk*xz-w∨no2*gm1sq*wy
      dd(2,1)=(gmisq*cqw-gmgmk*cpz)*rho
      dd(3,1)=-(gammk*gamm1*twgm1*aOpq+gam*gammk*gammk*xz+
                gammi*gmisq*wy)*rho
      dd(4,1)=(gmgmk*cqx-gm1sq*cpy)*rho
      dd(5,1)=-(2.*gmgmk*gm1sq*aOpq+gmgmk*gmgmk*xx*
                gmisq*gmisq*wy)*rho2
      dd(1,2)=(cqx-wvno2*cpy)/rho
      dd(2,2)=cpcq
      dd(3,2)=qammk*cqx-gamm1*cpy
      dd(4,2)=-xy
      dd(5,2)=dd(4,1)
      dd(1,4)=(wvno2*cqw-cpz)/rho
      dd(2, 4) = -wz
      dd(3,4)=qamm1*cqw-gammk*cpz
      dd(4,4) = dd(2,2)
      dd(5,4)=dd(2,1)
      dd(1,5)=-(2.*wvno2*aOpq+xz+wvno2*wvno2*wy)/rho2
      dd(2,5)=dd(1,4)
      dd(3,5)=-(twgm1*aOpq+gammk*xz+wvno2*gamm1*wy)/rho
      dd(4,5)=dd(1,2)
      dd(5,5) = dd(1,1)
      t=-2, *wvno2
      dd(1,3) = t*dd(3,5)
      dd(2,3)=t*dd(3,4)
      dd(3,3)=a0+2.*(cpcq-dd(1,1))
      dd(4,3)=t*dd(3,2)
      dd(5,3) = t * dd(3,1)
      return
      end
C
C
C
      subroutine energy(omega, wyno)
      This routine takes the energy integral by an analytic
C
      way using the eigenfuntions found above.
C
C
      double precision wynoO, omegaO
      double precision wyno2, omega2, wyomg2, sum, xka, xkb, ra, rb
      double precision dpth.daa.dbb.drho.dlam.dlamu.dmu
      double precision urur, urtz, uzuz, uztr, trtr, urduz, uzdur
      double precision durdur, duzduz, dldl, dldm, dldk, dldr
      double precision sumiO, sumi1, sumi2, sumi3
      double complex t(4,4), tt(4,4), ff(6), pp(4)
      double complex nua, nub, c1, c2
      common/coef/
                      t, tt, ff, pp
                      d(100), a(100), b(100), rho(100), qa1(100),
      common/model/
                         qb1(100), xmu(100), xlam(100), mmax, ll
      common/eigfun/ uu(100,4),uu0(4),dcda(100),dcdb(100),
                         dcdr(100)
      common/sumi/ - sumiQ, sumil, sumi2, sumi3, flagr, are, ugr
      wynoO=dble(wyno)
```

```
omegaO=dble(omega)
Omega2=omega0*omega0
wvomq2=wvnoO*omega2
w>no2=w>no0*w>no0
sumi0=0.0d+00
sumi1=0.0d+00
sumi2=0.0d+00
sumi3=0.0d+00
do 300 k=11, mmax
k1 = k+1
kk = k
if(k,eq.mmax) kk=101
daa=dble(a(k))
dbb=dble(b(k))
drho=dble(rho(k))
dlam=dble(xlam(k))
dmu≕dble(xmu(k))
dlamu=dlam+2. *dmu
dpth=dble(d(k))
xka=cmegaO/daa
xkb=omegaO/dbb
qammk=dbb/omegaO
gammk=2. *gammk*gammk
qam=gammk*wvno2
ra=dsgrt(dabs(wvno2-xka*xka))
rb=dsqrt(dabs(wvno2-xkb*xkb))
if(ra.1t.1,d-6) ra=1.d-6
if(rb. lt. 1. d-6) rb=1. d-6
nua=dcmplx(ra, 0.0)
nub=demplx(rb, 0, 0)
if(wynoO.lt.xka) nua=dcmplx(O.O.ra)
if(wvnoO, 1t, xkb) nub≈dcmp1x(O, O, rb)
call tminus(nua, nub, wvno2, gam, gammk, drho)
call tplus(nua, nub, wvno2, gam, gammk, drho)
call intval(kk, nua, nub, dpth)
do 200 i=1,2
i2=i+2
c 1 = O. O
c2=0.0
do 100 \text{ j} = 1.4
c1=c1+tt(i, j)*uu(k1, j)
c2=c2+tt(i2,j)*uu(k,j)
continue
pp(i)=c1
pp(i2)=c2
continue
urur=sum(kk,1,1)*w∨no2
urtz=sum(kk, 1, 3)*wvomg2
uzuz=sum(kk,2,2)
uztr=sum(kk,2,4)*wvomg2
trtz=sum(kk,3,3)*omega2*omega2
trtr=sum(kk,4,4)*wvomg2*wvomg2
urduz=(wvnoO*dlam*urur+urtz)/dlamu
```

100

200

```
uzdur=-wynoO*uzuz+uztr/dmu
     durdur=w√no2*uzuz-2. *w√no0*uztr/dmu+trtr/(dmu*dmu)
     durduz=(wyno2*dlam*dlam*urur+2. *wyno0*dlam*urtz+
                tztz}/(dlamu*dlamu)
    ij.
     sumiO=sumiO+drho*(uzuz+urur)
     sumil=sumil+dlamu*urur+dmu*uzuz
     sumi2=sumi2+dmu*uzdur-dlam*urduz
     sumi3=sumi3+dlamu*duzduz+dmu*durdur
      dldl=-wvno2*urur+2. *wvno0*urduz-duzduz
     dldm=-wvno2*(2, *urur+uzuz)-2, *wvno0*uzdur-(2, *duzduz+durdur)
      dldr=omega2*(urur+uzuz)
      dcda(k)=2.*drho*daa*omegaO*dldl/w∨no2
      dcdb(k)=2.*drho*dbb*omegaO*(d1dm-2.*d1d1)/wvno2
      dcdr(k)=dldr+dlam*dldl/drho+dmu*dldm/drho
      dcdr(k) #dcdr(k) *omegaO/wvno2
300
      continue
      if(b(1), le. O. O) call wenerg(wvnoO)
      dldk=-2.*(wvnoO*sumi1+sumi2)
      do 400 k=11, mmax
      dcda(k)=dcda(k)/dldk
      dcdb(k)=dcdb(k)/dldk
      dedr(k)=dedr(k)/dldk
400
      continue
      flagr=omega2*sumi0-wvno2*sumi1-2.*wvno*sumi2-sumi3
      ugr=(wvnoO*sumil+sumi2)/(omegaO*sumiO)
      are=wvnoO/(2.*omegaO*ugr*sumiO)
      return
      end
C
C
      function sum(kk,i,j)
C
      The analytic forms of the solution of integral:
r
      Integral U*U dz = T-matrix * eigenfaction * integral-coefs
C
C
      double precision sum
      double complex t(4,4), tt(4,4), ff(6), pp(4)
      double complex sumO, sum1, sum2, sum3, sum4, sum5, sum6
      common/coef/ t, tt, ff, pp
      if(kk.eq.101) go to 100
      sum1 = t(i, 1) * t(j, 1) * pp(1) * pp(1) + t(i, 3) * t(j, 3) * pp(3) * pp(3)
      sum1=sum1*ff(1)
      sum2=t(i,2)*t(j,2)*pp(2)*pp(2)+t(i,4)*t(j,4)*pp(4)*pp(4)
      sum2=sum2*ff(2)
      sum3=(t(i,1)*t(j,2)+t(i,2)*t(j,1))*pp(1)*pp(2)
             +(t(i,3)*t(j,4)+t(i,4)*t(j,3))*pp(3)*pp(4)
      (E) 44*Emue=Emue
      sum4=(t(i, 1)*t(j, 4)+t(i, 4)*t(j, 1))*pp(1)*pp(4)
             +(t(i,3)*t(j,2)+t(i,2)*t(j,3))*pp(2)*pp(3)
      sum4=sum4*ff(4)
      sum5=(t(i,1)*t(j,3)+t(i,3)*t(j,1))*pp(1)*pp(3)
       sum5=sum5*ff(5)
```

```
sum6=(t(i,2)*t(j,4)+t(i,4)*t(j,2))*pp(2)*pp(4)
      sumb=sumb*ff(6)
      sumQ=sum1+sum2+sum3+sum4+sum5+sum6
      sum=real(sumO)
      return
 100 continue
      sum1=t(i,3)*t(j,3)*pp(3)*pp(3)*ff(1)
      sum2=t(i,4)*t(j,4)*pp(4)*pp(4)*ff(2)
      sum3=(t(i,3)*t(j,4)+t(i,4)*t(j,3))*pp(3)*pp(4)
      sum3=sum3*ff(3)
      SumO=sum1+sum2+sum3
      sum=real(sumO)
      return
      end
C
C
C
      subroutine intval(k, nua, nub, dpth)
C
      This routine finds the coeficients needed for integrals.
C
C
      double precision dpth
      double complex t(4,4), tt(4,4), ff(6), pp(4)
      double complex nua, nub, p, q, pq, expp, exqq
      common/coef/ t, tt, ff, pp
      if(k.eq. 101) go to 100
      p≖nua≯dpth
      q=nub*dpth
      pq=(nua+nub)*dpth
      call ifpu(p,p+p,expp)
      ff(1)=(1, O-expp)/(2, *nua)
      call ifpq(q,q+q,exqq)
      ff(2)=(1, 0-exqq)/(2, *nub)
      call ifpq(pq/2 ,pq,expp)
      ff(3)=(1, O-expp)/(nua+nub)
      call ifpq(p/2, p, expp)
      call ifpq(q/2.,q,exqq)
      ff(4)=(exqq-expp)/(nua-nub)
      ff(5)=dpth*expp
      ff(6)=dpth*exgq
      return
 100
      continue
      ff(1)=0.5/nua
      ff(2)=0.5/nub
      ff(3)=1./(nua+nub)
      return
      end
C
C
C
      subroutine ifpq(p,pq,expq)
      double complex p.pq.expq
      if(real(p), lt. 40.0) go to 100
```

```
expq=0.0d+00
      go to 200
      continue
100
      expq=zexp(-pq)
      continue
500
      return
      end
C
C
C
      subroutine tplus(nua, nub, wvno2, gam, gammk, rho)
C
      T matrix.
C
C
      double precision gam, gammk, rho, wyno2
      double complex t(4,4), tt(4,4), ff(6), pp(4)
      double complex nuarnub
      common/coef/ t, tt, ff, pp
      t(1,1)=-1./rho
      t(1,2)=nub/rho
       t(1,3)=t(1,1)
       t(1,4)=-t(1,2)
       t(2,1)=-nua/rho
       t(2,2)=wvno2/rho
       t(2,3)=-t(2,1)
       t(2,4)=t(2,2)
       t(3,1)=1.-gam
       t(3,2)=gam*nub
       t(3,3)=t(3,1)
       t(3,4)=-t(3,2)
       t(4,1)=-gammk*nua
       t(4,2)=-t(3,1)
       t(4,3) = -t(4,1)
       t(4,4)=t(4,2)
       do 100 i=1,4
       do 100 j=1,4
       t(i,j)=0.5*t(i,j)
  100
       continue
       return
       end
 C
 C
       subroutine tminus(nua, nub, wvno2, gam, gammk, rho)
 C
       T-inverse matrix
 C
 C
       double precision gam, gamm1, gammk, rho, wvno2
       double complex t(4,4), tt(4,4), ff(6), pp(4)
        double complex nuarnub
        common/coef/ t, tt, ff, pp
        gamm1=gam-1.0
        tt(1,1)=-rho*gam
```

```
tt(1,2)=rho*gamm1/nua
      tt(1,3)=1.0
      tt(1,4)=-wvno2/nua
      tt(2,1) = -rho*qamm1/nub
      tt(2,2)=rho*gammk
      tt(2,3)=1. /nub
      tt(2,4)=-1.0
      tt(3,1)=tt(1,1)
      tt(3,2) = -tt(1,2)
      tt(3,3)=1.0
      tt(3,4)=-tt(1,4)
      tt(4,1) = -tt(2,1)
      tt(4,2)=tt(2,2)
      tt(4,3) = -tt(2,3)
      tt(4,4) = -1.0
      return
      end
C
c
C
      subroutine wenerq(wvno)
      calculate energy trapped in the top water layer.
C
      double precision wyno, wd, wra, wra2, wcosd2, wsin2d
      double precision urur, uzuz, urduz, duzduz, wvno2, dr, dlam
      double precision sumiO, sumi1, sumi2, sumi3
      common/model/ d(100), a(100), b(100), rho(100), qa1(100),
                        qb1(100), xmu(100), xlam(100), mmax, ll
      common/sumi/ _ sumiO, sumi1, sumi2, sumi3, flagr, are, ugr
      common/engerw/ wd.wra.wcosd2.wsin2d
      wyno2=wyno*wyno
      wra2=wra*dabs(wra)
           =(wsin2d-wd)/(wra2*wcosd2)
      UTUT
             =urur*wvno2
      UTUT
             =(wsin2d+wd)/wcosd2
      UZUZ
      urduz =(wsin2d-wd)/wcosd2
      urduz =urduz*w∨no
      duzduz=(wsin2d-wd)*wra2/wcosd2
      dr=dble(rho(1))
      dlam=dble(xlam(1))
      sumiQ=sumiO+dr*(urur+uzuz)
      sumil=sumil+dlam*urur
      sumi2=sumi2-dlam*urduz
      sumi3=sumi3+dlam*duzduz
      return
      end
****
```

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```
leigen81.f
C
C
      f77 -i -12 leigen81.f -o leigen81
C
C
      This is a new version of program leigen which calculates
C
      the eigenfunctions of Love wave for any plane layered
C
      model.
C
C
      Some analytic ofms for energy integrals are used instead
C
      of taking numerical integration directly.
C
C.
      The I/O are mostly the same as those of program leigen,
c
      except that the eigenfunctions calculated are located at
C
      the top of each layer, not in the middle.
r
r
      The Q-model is taken into account.
C
      for the sake of space saving, only the values of
С
      eigenfunction at the source depth are stored.
c
      However, several files with different source depths
c
      can be set up.
C
c
      --Oct 10, 1981
      dimension nos(100), dphs(100), dphq(100), qa(100), qb(100)
      dimension depth(100)
                      d(100), a(100), b(100), rho(100), qa1(100),
      common/model/
                         qb1(100), xmu(100), xlam(100), mmax, ll
      common/eigfun/ ut(100), tt(100), dcdb(100), dcdr(100), uu0(4)
                      sumiO, sumi1, sumi2, flagr, ale, ugr
      common/sumi/
       character*1 dd
       character*50 names
 5
       format(a)
       format(/2x, 'M', 3x, 'DPTH', 3x, ' D ', 3x, ' A ', 3x, ' B ',
 10
                3x, ' RHO', 3x, ' QA ', 4x, ' QB ', 4x, ' MU ', 3x, 'LMDA')
       format(i3,1x,5f7.2,2f8.2,2f7.2)
 20
       format(3x, '-source is at the top of this layer. ',
 30
                ' source depth = ', f6.2)
       format(i3,1x,f7,2,7x,3f7,2,2f8,2,2f7,2)
 40
       format(3x, '-source is inside the half space.',
 50
                 'source depth = ', f6.2)
       format(' at the source depth = ',f6.2,' km')
 60
       write(6,*) 'enter the input file name: (from surface81)'
       read(5,5) names
       open(1, file=names, status='old', form='unformatted')
       open(3, file='tmp. 1', status='scratch', form='unformatted')
       rewind 3
       enter the source depths and Q model.
 C
       write(6,*) 'enter the source depths: '
       write(6,*)
           '(no. of source depths, srcdph(1), srcdph(2),...)'
```

```
write(6, *)
         '*** if no. of source depths is negative, no output'
     write(6,*) '
                     files will be generated. ***'
     read(5,8) kks, (dphs(i), i=1, iabs(kks))
     ks#iabs(kks)
     write(6,*) 'enter G-model here(1), or from file(2) or'
     write(6,*) 'NOT consider Q in this program(3): '
     read(5,*) kk
     kq=1
     go to (100, 110, 170), kk
100
     continue
     write(6, *)
         'enter d(i), Qa(i), Qb(i) use d(i)=0 for halfspace'
      icode=5
     go to 120
110 continue
     write(6,*) 'enter the name of the file storing Q-model: '
     read(5/5) names
     open(2, file=names, status='old', form='formatted')
     rewind 2
      icode=2
     continue
120
      i = 1
      base=0.0
140
     continue
      read(icode,*) dqO,qa(i),qb(i)
      if(dq0, le, 0, 0) go to 150
      base=base+dqO
      dphq(i)=base
      i = i + 1
      go to 140
150
      kq=i
      do 160 i=1,100
      qa1(i)=qa(kq)
      qb1(i)=qb(kq)
160
      continue
170
      continue
      write(6,*) 'Store the derivatives? (y/n)'
      read(5,5) dd
      if(dd.eg. 'n') go to 175
      write(6,*) 'enter the output file name for derivatives: '
      read(5,5) names
      open(9, file=names, status='new', form='unformatted')
      rewind 9
 175
      continue
      do 180 i=1,100
      depth(i)=1000000.0
 180
      continue
C
      enter the earth model.
C
      read(1) mmax
      write(6,10)
```

```
base=0.0
      depth(1)=0.0
      do 190 i=1, mmax
      read(1) d(i), a(i), b(i), rho(i)
      base=base+d(i)
      depth(i+1)=base
      xmu(i) = rho(i) *b(i) *b(i)
      xlam(i)=rho(i)*(a(i)*a(i)-2.*b(i)*b(i))
 190
      continue
C
      insert the G-model into the velocity model.
C
      insert the source depth at the boundary of a layer.
С
C
      kqs=kq+ks-1
      do 300 k0=1, kqs
      dphO=dphq(kO)
      is=kO-kq+1
      if(kO, ge, kg) dphO=dphs(is)
      do 200 i=1, mmax
      if(dphO.eq.depth(i)) go to 250
      if(dphO, gt, depth(i), and, dphO, lt, depth(i+1)) go to 210
 200
      continue
      k1=k+1
 210
      dohk1=depth(k1)
      do 220 i=mmax, k, -1
      i1 = i+1
      d(i1)=d(i)
      a(i1)=a(i)
      b(i1)=b(i)
      rho(i1)=rho(i)
       xmu(i1) = xmu(i)
       xlam(i1)=xlam(i)
       depth(i1)=depth(i)
       if(kO, ge, kq) = qal(il) = qal(i)
       if(kO, ge, kq) qbi(i1)=qbi(i)
       continue
 220
       d(k) = dphO - depth(k)
       d(k1) = dphk1 - dph0
       depth(k1)=depth(k)+d(k)
       mmax=mmax+1
       if(kO.ge.kq) go to 240
       if(k0, eq. 1) ns=1
       do 230 j=ns, k
       qa1(j)=qa(kO)
       qb1(j)=qb(kO)
 230
       continue
 240
       ns=k1
       go to 280
 250
       continue
       if(kO, ge, kq) go to 270
       if(k0, eq. 1) ns=1
       do 240 i=ns,k-1
```

```
qa1(i)=qa(k0)
      qb1(i)=qb(kQ)
 240
      continue
 270
      ns=k
 280
      nos(is)=ns
 300
      continue
      1=1
      do 310 i=1, mmax-1
      write(6,20) i,depth(i),d(i),a(i),b(i),rho(i),qa1(i),qb1(i),
                      xmu(i), xlam(i)
      if(i.ne.nos(j)) go to 310
      write(6,30) dphs(j) -
      J=1+1
 310
      continue
      i=mmax
      write(6,40) i,depth(i),
            a(i), b(i), rho(i), qa1(i), qb1(i), xmu(i), xlam(i)
      if(nos(j), eq. mmax) write(6,50) dphs(j)
      if(dd.eq.'y')
         write(9) mmax, (d(i), a(i), b(i), rho(i), i=1, mmax)
      write(6,*) ' '
      write(6,*) 'wait.'
      11=1
      if(b(1), le. 0, 0) 11=2
      read(1) nper
 400
      continue
C
C
      read in the dispersion values.
C
      read(1) ifunc, mode, t
      write(3) ifunc, mode, t
      if(dd.eq.'y') write(9) ifunc, mode, t
      if(ifunc. 1t. 0) go to 700
      if(mode.le.O) go to 400
      do 600 k=1, mode
      read(1) c
C
      main part.
C
c
      omega=6, 2831853/t
      wvno=omega/c
      call shfunc(omega, wvno)
      call energy(omega/wyno)
      if(kk.ne.3) call gammaq(omega,wvno,gamma)
      if(dd.eq.'n') go to 510
C
      output the derivatives.
      write(9) dum, uuO(2), c, ugr, sumiO, sumi1, sumi2, dum, ale, flagr
      do 500 i=1, mmax
      write(9) depth(i), ut(i), tt(i), dum, dum, dum, dcdb(i), dcdr(i)
 500
      continue
 510
      continue
      do 560 i=1,ks
      j≕nos(i)
```

```
uts=ut(j)
      duts=tt(j)/xmu(j)
      write(3) wyno, ale, ugr, gamma, uts, duts
 560
      continue
      continue
 600
      ap to 400
      continue
 700
C
      output the data files with different source depths.
C
C
      if(kks.le.0) go to 950
      do 900 i=1,ks
      rewind 3
      write(6, *)
     * 'enter the name of output file storing the eigenfunctions'
      write(6,60) dphs(i)
      read(5,5) names
      open(4, file=names, status='new', form='unformatted')
      rewind 4
      write(4)
          mmax, (d(j), a(j), b(j), rho(j), qa1(j), qb1(j), j=1, mmax)
      write(4) nper, dphs(i)
      continue
 800
      read(3) ifunc, mode, t
      write(4) ifunc, mode, t
      if(ifunc. lt. O) go to 870
      if(mode. Le. O) go to 800
      do 860 k=1, mode
      do 850 j=1,ks
      read(3) qa(j),qb(j),dphq(j),ut(j),tt(j),dcdb(j)
 850
      continue
      write(4) qa(i), dummy, qb(i), dphq(i), ut(i)
      write(4) tt(i), dcdb(i), dummy, dummy
 860
      continue
      go to 800
 870
      continue
      close(4)
      continue
 900
      continue
 950
      close(1)
      close(2)
      close(3, status='delete')
      write(6,*) ' '
      write(6,*) 'leigen81 finished'
      write(6,*) ' '
       stop
       end
C
C
C
       subroutine gammaq(omega,wvno,gamma)
       This routine finds the attenuation gamma value.
C
```

```
common/model/
                       d(100), a(100), b(100), rho(100), qa1(100),
                         qb1(100), xmu(100), xlam(100), mmax, ll
      common/eigfun/ ut(100), tt(100), dcdb(100), dcdr(100), uu0(4)
      x # O. Ø
      do 100 i=11, mmax
      x = x + dcdb(i) + b(i) / qb1(i)
 100
      continue
      c=omega/wyno
      qamma=0.5*w∨no*x/c
      return
      end
C
С
C
      subroutine shfunc(omega, wyno)
C
      This routine evaluates the eigenfunctions by calling sub
C
C
      double precision exl(100), ext, fact
      common/model/ d(100), a(100), b(100), rho(100), qa1(100),
                         qb1(100), xmu(100), xlam(100), mmax, 11
      common/eigfun/ uu(100,2),dcdb(100),dcdr(100),uuQ(4)
      common/save/
                       ex1
      call up(omega, wvno, fl)
      000(1)=1.0
      oud(2)=stressO is actually the value of period equation.
C
C
      UUC(3) is used to print out the period evation value before
      the root is refined.
~
      UUO(2)=f1
      uu0(3)=0.0
      uuO(4)=0.0
      ext=0.0
      do 100 k=11+1, mmax
      ext=ext+ex1(k-1)
      fact=0.0
      if(ext. lt. 85.0) fact=1, /dexp(ext)
      υυ(k,1)=υυ(k,1)*fact/υυ(ll,1)
      uu(k,2)=uu(k,2)*fact/uu(11,1)
 100
      continue
      00(11,1)=1.0
      00(11,2)=0.0
      return
      end
C
C
C
      subroutine up(omega, wvno, fl)
C
      This routine calculates the elements of Haskell matrix,
C
      and finds the eigenfunctions by analytic solution.
C
      double precision ex1(100), qq, rr, ss, exqm, exqp, sinq, cosq
      common/model/ d(100), a(100), b(100), rho(100), qa1(100),
```

```
qb1(100), xmu(100), xlam(100), mmax, ll
     common/eigfun/ uu(100,2),dcdb(100),dcdr(100),uu0(4)
     common/save/
                     exl
     wyno2=wyno*wyno
     xkb=omega/b(mmax)
     rb=sqrt(abs(wvno2-xkb*xkb))
     if(wyno.lt.xkb) write(6,*) ' imaginary nub'
     uu(mmax, 1)=1.0
     uu(mmax,2)=-xmu(mmax)*rb
     mm \times 1 = mma \times -1
     do 500 k=mmx1,11,-1
     k1 = k+1
     doth=d(k)
     xkb=omega/b(k)
     rb=abs(wvno2-xkb*xkb)
     rr#dble(rb)
     rr=dsart(rr)
     ss=dble(dpth)
     qq=rr#ss
     if(wyno-xkb) 100,200,300
100
     sinq=dsin(qq)
     cosq=dcos(qq)
     y#sinq/rr
     z=-rr*sinq
     qq=0. O
        to 400
     90
500
     qq=0.0
     cosq=1.0d+0
     y≕dpth
     z=0.0
     go to 400
300
     if (qq. g(, 40, 0) go to 350
     exqp=1.
     exqm=1./dexp(qq+qq)
     sinq=(exqp-exqm)*0.5
     cosq=(exqp+exqm)*0.5
     y=sinq/rr
     z=rr*sinq
     go to 400
350
     continue
     u=0.5/rr
     z=0.5*rr
     cosq=0.5
400
     continue
     ampO=cosq*uu(k1,1)-y*uu(k1,2)/xmu(k)
     strO=cosq*uu(k1,2)-z*xmu(k)*uu(k1,1)
     rr=abs(ampQ)
     ss=abs(strO)
     if(ss.gt.rr) rr#ss
     if(rr. 1t. 1. d-30) rr=1. d+00
     ex1(k)=dlog(rr)+qq
     uu(k,1)=ampO/rr
     uu(k,2)=str0/rr
```

```
continue
 500
      £1=00(11,2)
      return
      end
C
C
C
      subroutine energy(omega, wvno)
      This routine calculates the values of integrals IO, I1,
C
      and I2 using analytic solutions. It is found
c
      that such a formulation is more efficient and practical.
C
C
      double precision wynoO, omegaO, c, sumiO, sumi1, sumi2
      double precision xkb, rb, dbb, drho, dpth, dmu, wvno2, omega2
      double precision upup, dupdup, dcb, dcr
      double complex nub, xnub, exqq, top, bot, f1, f2, f3
                      d(100), a(100), b(100), rho(100), qa1(100),
      common/mode1/
                         qb1(100), xmu(100), xlam(100), mmax, ll
      common/eigfun/ uu(100,2),dcdb(100),dcdr(100),uu0(4)
                      xiO, xi1, xi2, flagr, ale, ugr
      common/sumi/
      wynoO=dble(wyno)
      omegaO=dble(omega)
      c=omega0/w∨no0
      omega2=omega0*omega0
      wvno2=wvno0*wvno0
      sumiO=0.0d+00
      sumi1=0.0d+00
      sumi2=0.0d+00
      do 300 k=11, mmax
      k1 = k+1
       dbb≈dble(b(k))
       drho=dble(rho(k))
       dpth=dble(d(k))
       dmu=dble(xmu(k))
       xkb=omegaO/dbb
       rb=dsgrt(dabs(wvno2-xkb*xkb))
       if(k, eq. mmax) go to 100
       nub=dcmplx(rb)O.O)
       if(wynoO.lt.xkb) nub=dcmplx(O.O,rb)
       dun*umb≖duux
       top=uu(k,1)-uu(k,2)/xnub
       bot=uu(k1,1)+uu(k1,2)/xnub
       f3=nub*dpth
       exqq=0.0
       if(real(f3), lt. 40) exqq=zexp(-2, %f3)
       f1=(1, -exqq)/(2, *nub)
       exqq≖0. O
       if(real(f3), lt.80) exqq=zexp(-f3)
       f2=dpth*exqq
       f1=0. 25*f1*(top*top+bot*bot)
       f2=0.5 #f2*top*bot
       53+1+52
       upup=real(f3)
```

```
43=xnub*xnub*(f1-f2)
     dupdup=real(#3)/(dmu*dmu)
     go to 200
100 continue
     upup ≈0.5/rb*uu(mmax,1)*uu(mmax,1)
     dupdup=0.5%rb*uu(mmax,1)*uu(mmax,1)
     continue
200
     sumiO=sumiO+drho*upup
     sumi1=sumi1+dmu*upup
     sumi2=sumi2+dmu*dupdup
     dcr=-0.5%c*c*c*upup
     dcb=0.5*c*(upup+dupdup/wvno2)
     dcdb(k)=2. *drho*dbb*dcb
     dcdr(k)=dcr+dbb*dbb*dcb
300
     continue
     do 400 k=11, mmax
      dcdb(k)=dcdb(k)/sumi1
     dcdr(k)=dcdr(k)/sumil
     continue
400
     flagr=omega2*sumi0-wvno2*sumi1-sumi2
     ugr=sumi1/(c*sumi0)
      ale=0.5/sumil
      xiO=sumiO
      xil=sumil
      xi2=sumi2
      return
      end
***
```

```
dpegn81.f
C
c
      f77 -i -12 dpegn81 f -o dpegn81 -lcalcomp12
C
C
      plot dispersion curve
c
      get data from reigen81 or leigen81
C
C
      dimension t(1000), v(1000), mode(500)
      dimension d(500), a(500), b(500), rho(500)
      character*20 names
      character*4 yorn, xlog, ylog
      common/ ctrl / yorn, xlog, ylog
      format(a)
      format(4f12.5)
  15
      format(5x, i5, 3x, e10, 4, 3x, e10, 4)
  25
      pi=2. *3. 141592653
      write(6,*) ' '
      write(6,*) 'Rayleigh(1) or Love(2)?'
      read(5,*) llrr
      write(6,*) ' '
      if(11rr. eq. 1) write(6,*)
          'enter input file name: (from reigen81)'
      if(11rr. eq. 2) write(6,*)
          'enter input file name: (from leigen81)'
      read(5,5) names
      open(1, file=names, status='old', form='unformatted')
      write(6,*)
          'how many modes and which modes wanted? (e.g. 3,1,3,5)'
      write(6,*) '(if all modes wanted, answer 1,0 )'
      read(5,*) nmode,(mode(i),i=1,nmode)
      if(mode(1), ne. 0) go to 40
      nmode=500
       do 30 i=1,500
      mode(i)=i
  30
      continue
  40
       continue
                                                           C-frea(4)'
                                    U-freq(2)
                                                C-perd(3)
       write(6,*) 'plot U-perd(1)
                                                          - C-K(B)'
                         K-freq(5)
                                    freq-K(6)
                                                K-C(7)
       write(6, *)
                         Ar-perd(9)
                                            Ar-freq(10)'
       write(6,%)
                                            gamma-freq(12)'
                         gamma-perd(11)
       write(6, *)
                         Ar*atten-perd(13) Ar*atten-freq(14): '
       write(6, *)
       read(5,*) ipg
       write(6,*) 'plot X axis in log scale? (y/n): '
       read(5,5) xlog
       write(6,*) 'plot Y axis in log scale? (y/n): '
       read(5,5) ylog
       call plots(0,0,7)
       write(6,*) 'enter ipen: '
       read(5,*) ipen
       call newpen(ipen)
       gorn='g'
       call plot(2.5,2.3,-3)
       write(6,*) ' '
```

```
write(6,≯) 'model:'
    do 600 kk=1, nmode
    rewind 1
    read(1) nmax, (d(i), a(i), b(i), rho(i), ga1, qb1, i=1, nmax)
    if(kk, eq. 1) write(6, 15) (d(i), a(i), h(i), rho(i), i=1, nmax)
    read(1) nper, dphs
    kkmode=0
100 continue
    ident=mode(kk)
    nt=0
    itrig=0
200 continue
    read(1) ifunc, kmode, t1
    if(ifunc. 1t. 0) go to 400
    if(kmode, le. 0) go to 200
    nt=nt+1
    do 300 i=1, kmode
    read(1) wyno0, ur0, are0, u0, gama0
    cOmpi/(t1*w∨noO)
    read(1) urO, durO, uzO, duzO
    if(i.ne.ident) go to 300
    itrig=1
    nn=nt
    t(nn)=t1
    \vee (nn) = c0
    if(ipg, le, 2) \lor (nn) = 00
    if(ipg. eq. 9. or. ipg. eq. 10) ∨(nn)=areO/sqrt(w∨noO)
    if(ipq.eq. 11. or. ipg.eq. 12) v(nn)=gamaO
    if(ipg. eq. 13. or. ipg. eq. 14)
        \forall(nn)=areO/(sqrt(\omega\noO)*exp(gamaO*1000.O))
300 continue
    go to 200
400 continue
    if(itrig.eq.O. and. nmode.eq. 500) go to 600
    if(t(nn), gt. 5000.0) nn=nn-1
    do 550 i=1, nn
    go to (500,482,500,482,485,486,487,488,500,482,500,482,
               500,482), ipq
482 t(i)=1./t(i)
    go to 500
485 \ t(i)=1./t(i)
    \vee(i)=pi*t(i)/\vee(i)
    go to 500
486 \text{ tmp} = 1./t(i)
    t(i)=pi*tmp/v(i)
    \forall(i)=tmp
    go to 500
487 tmp=v(i)
    \vee(i)=pi/(t(i)*\vee(i))
    t(i)=tmp
    go to 500
488 t(i)=pi/(t(i)*v(i))
500 continue
```

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```
i\hat{\tau}(x\log, eq, 'y') = t(i) = log 10(t(i))
      if(ylog eq. 'y') \lor (i)=logiO(\lor(i))
 550 continue
      call disp(itrig, kk, ipg, ident, nn, t, v)
 600 continue
      call plot(0.0,8.0,999)
      close(1)
      close(2)
      write(6,*) ' '
      write(6,*) 'Job finished'
      write(6,*) ' '
      stop
      end
C
      subroutine disp(itrig, kk, ipg, ident, nn, x, y)
      dimension x(1), y(1)
      character*4 yorn, nory, xlog, ylog
      character*11 alpha(2,14), alp, bet
      common/ ctrl / yorn, xlog, ylog
      data alpha/
                                                          (Hz)
                                           (km/sec)','f
                                     1, 10
             (km/sec)','T
                            (sec)
     *
         40
                                           (km/sec)','f
                                                          (Hz)
                                     1,10
           (km/sec)','T
                            (sec)
                                                          (rad/km)'
                                                    17 1K
                                     1,14
                                           (Hz)
         'K (rad/km)', 'f
                            (Hz)
                                                          (rad/km) '.
                            (km/sec)','C
                                           (km/sec)', 'K
            (rad/km)/,/C
     ×
                                    ', 'Amp Factor ', 'f
                                                          (Hz)
        'Amp Factor ', 'T
                            (sec)
                                                    1,14
                                                          (Hz)
                                           Gamma
                            (sec)
                      ', 'T
     솭
             Gamma
                                                                   11
                                     ', ' Ar*atten ', 'f
                                                          (Hz)
                      1, 17
                            (sec)
          Ar*atten
      if(itrig eq.O) write(6,*) 'mode ',ident,' is not generated.'
      if(itrig.eq.O) return
      alp=alpha(1, ipg)
      bet=alpha(2, ipg)
       if(yorn, eq. 'n') go to 100
      uorn='n'
       x1en=6.0
       glen=4.5
       xmin=1. e+37
       x = -1.e + 37
       umin=1.e+37
       ymax = -1.e + 37
       do 80 i=1, nn
       if(x(i).ge.xmax) xmax=x(i)
       if(x(i), le. xmin) xmin=x(i)
       if(q(i).ge.ymax) ymax=y(i)
       if(u(i), le. ymin) ymin=y(i)
  80
       continue
       write(6,*) 'ymin=',ymin,' ymax=',ymax
       if(ylog.eq.'y') write(6.*) ' -y value in log sacle'
       write(6,*) 'xmin=',xmin,' xmax=',xmax
       if(xlog.eq.'y') write(6,*) ' -x value in log sacle'
       write(6,*)
       write(6,*) 'enter ymin,ymax,yinc,xmin,xmax,xinc'
       if(ylog.eq.'y') write(6,*) ' -x should be integer.'
       if(ylog.eq.'y') write(6,*) '
                                            -y should be integer. '
```

```
read(5,*) y0,y1,yinc,x0,x1,xinc
    xinut=(xi-xO)/xlen
    qinut=(g1-g0)/glen
100 continue
    if(kk.ne.1) go to 500
    write(6,*) 'plot axis? (y/n)'
    read(5,5) nory
    format(a)
5
    write(6, *) ' '
    write(6,*) 'wait. It is processing.'
    if(nory.ne.'y') go to 500
    call plot(xlen, 0, 0, 2)
    call plot(xlen,ylen,2)
    call plot(0.0,ylen,2)
    call plot(0.0,0.0,2)
    call plot(0.0,-0.05,2)
    if(xlog.eq.'y') go to 110
    if(x0, 1t, 0, 0) xshif=-0.25
    if(x0, ge, 0, 0) xshif=-0, 13
    if(x0, ge. 10, 0) xshif=-0.23
    call number(xshif, -0.17, 0.1, x0, 0.0, 1)
    go to 120
110 call number (-0, 12, -0, 18, 0, 1, 10, 0, 0, 0, -1)
    call number(999.,-0.1,0.06,x0,0.0,-1)
120 continue
    grid=xinc/xinut
    x i = 1.
200 xx=xi*grid
     if(abs(xx).gt.xlen+0.02) go to 250
     call plot(xx,0.0,3)
     call plot(xx,-0.05,2)
     xsym=xO+xinc*xi
     if(xlog.eq.'y') go to 210
     if(xsym. 1t. 0. 0) xshif=-0. 25
     if(xsym.ge, O. O) xshif=-0.13
     if(xsym.ge. 10.0) xshif=-0.23
     if(xsym.ge. 100.0) xshif=-0.33
     call number(xx+xshif, +0, 17, 0, 1, xsym, 0, 0, 1)
     go to 220
210 continue
     call number(xx-0, 12, -0, 18, 0, 1, 10, 0, 0, 0, -1)
     call number (999., -0.1, 0.06, xsym, 0.0, -1)
220 continue
     xi = xi+1.
     go to 200
       continue
250
     call symbol(2, 5, -0, 5, 0, 16, bet, 0, 0, 11)
     call plot(0.0,0.0,3)
     call plot(-0.05,0.0,2)
     if(ylog.eq.'y') go to 260
     if(y0, 1t. 0. 0) xshif=-0. 45
     if(y0.ge.0.0) xshif=-0.35
     if(y0, ge, 10, 0) \times shif=-0.45
```

```
call number(xshif, -0.05, 0.1, y0, 0.0, 1)
     go to 270
 260 continue
     call number (-0.34, -0.06, 0.1, 10.0, 0.0, -1)
     call number (-0, 16, 0, 02, 0, 06, y0, 0, 0, -1)
 270 continue
     orid=uinc/yinut
     x i = 1.
 300 xx=xi*grid
      if(abs(xx), gt. ylen+0.02) go to 400
      call plot(O.O,xx,3)
      call plot(-0.05, xx, 2)
      xsum=qO+qinc*xi
      if(ylog.eq.'y') go to 310
      if(xsym. 1t. 0. 0) xshif=-0. 45
      if(xsym.ge. 0. 0) xshif=-0.35
      if(xsym.ge, 10.0) xshif=-0.45
      if(xsym. ge. 100. 0) xshif=-0.55
      call number(xshif, xx-0, 05, 0, 1, xsym, 0, 0, 1)
      go to 320
 310 continue
      call number(-0.34, xx-0.06, 0.1, 10.0, 0.0, -1)
      call number(-0.16, xx+0.02, 0.06, xsym, 0.0, -1)
 320 continue
      xi = xi+1.
      go to 300
 400 continue
      call symbol(-0.55, 1.5, 0.16, alp, 90.0, 11)
 500 continue
      do 550 i=1, nn
      if(u(i), qt, y1) y(i)=y1
      if(x(i), gt, x1) x(i)=x1
      if(y(i), lt, y0) y(i)=y0
      if(x(i), lt, xO) x(i)=xO
  550 continue
      x(nn+1)=x0
      x(nn+2)=xinut
      y(nn+1)=40
      y(nn+2)=yinut
      call line(x, y, nn, 1, 0, 0)
      return
      end
***
```

```
deriv81.f
C
r
      f77 -i -I2 deriv81.f -o deriv81
C
      This program reads the eigenfunctions and derivatives
C
      at different depths from an output file of r(1)eigen81.
C
r
      common/model/ d(500), a(500), b(500), rho(500)
      common/eigfun/ depth(500), ur(500), uz(500), tz(500), tr(500),
                        dcda(500), dcdb(500), dcdr(500)
      character*50 names
      format(a)
  5
  10 format(///22x,'M',3x,' D ',3x,' A ',3x,' B ',5x,'RHO')
  20
      format(20x, i3, 1x, 4f7, 2)
      format(20x, i3, 8x, 3f7, 2)
  30
  40
      format(////)
      format(////30x/'T(SEC) = ', f7.3/)
  45
      format(10x)' = ",e11.4,' C(KM/S) = ',f7.4,
                   U(energy) = ', f7.4
      format(7x, 'C(KM/S) = ', f7.4, ' U(energy) = ', f7.4,
  51
                  ALE = (, e11.4)
      format(10x, ' IO = ', e11. 4, ' II = ', e11. 4, '
  55
               e11.4)
      56
               e11.4, ' L = ', e11.4)
      format(10x, ' I3 = ',e11.4, ' ARE = ',e11.4, ' L = ',
  60
               e11.4)
      format(//
                  M', 5x, 'DEPTH', 5x, 'UR', 10x, 'UZ', 10x, 'TZ', 10x,
                'TR', 9x, 'DCDA', 8x, 'DCDB', 8x, 'DCDR'/)
      format(/'
                  M', 8x, 'DEPTH', 6x, 'DISP', 8x, 'STRESS', 8x,
  71
                'DC/DB', 8x, 'DC/DR'/)
     format(1x, i3, f10, 2, 7(1x, e11, 4))
  75
      format(1x, i3, f13, 2, 4(2x, e11, 4))
  76
      format(/9x, 'STRESSO = ', e11.4)
  80
      write(6,*) 'Rayleigh(1) or Love(2): '
      read(5,*) llrr
      write(6.*) 'enter the input file name: (from r<1>eigen81)'
      read(5,5) names
      open(1, file=names, status='old', form='unformatted')
      read(1) mmax, (d(i), a(i), b(i), rho(i), i=1, mmax)
      write(6,*) ' '
      write(6,10)
      do 100 i=1, mmax-1
      write(6,20) i,d(i),a(i),b(i),rho(i)
  100 continue
      i =mmax
      write(6,30) i,a(i),b(i),rho(i)
  200 continue
      read(1) ifun, mode, per
      if(ifun. 1t. 0) go to 500
      if(mode.le.O) go to 200
      write(6,40)
```

```
omega=6.2831853/per
     do 400 k=1/mode
      read(1) e, ee, c, ugr, xiO, xi1, xi2, xi3, are, flagr
      wyno=omega/c
      write(6,45) per
      go to (210,220), llrr
 210 continue
      write(6,50) e.c.ugr.
      write(6,55) xiQ,xi1,xi2
      write(6,60) xi3, are, flagr
      go to 250
 220 continue
      write(6,51) c.ugr.are
      write(6,56) xiO, xi1, xi2, flagr
 250 continue
      if(11rr.eq. 1) write(6,70)
      if(11rr.eq. 2) write(6,71)
      do 300 i=1, mmax
      read(1) depth(i), ur(i), uz(i), tz(i), tr(i),
                 dcda(i), dcdb(i), dcdr(i)
      if(11rr.eq.1)
     * write(6,75) i,depth(i),ur(i),uz(i),tz(i),tr(i),
                     dcda(i), dcdb(i), dcdr(i)
      if(11rr.eq.2)
     * write(6,76) i,depth(i),ur(i),uz(i),dcdb(i),dcdr(i)
 300 continue
      write(6,80) ee.
  400 continue
      go to 200
  500 continue
      stop
      end
****
```

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```
wig81.f
c
C
      f77 -i -12 wig81 f -o wig81
C
C
      This program ganerates ten basic types of synthetic
C
      spectrum after combining the eigenfunctions from
c
      reigen81 and/or eigen81 with the source spectrum.
C
C
      The output can be a seismogram which will input to
C
      the program gle81 or a spectrum which will input to
C
      the program spec81 for plotting.
C
C
      The dimension needed has been reduced to the minimum by
C
      using the system call.
C
      The program 'bigfft' to do fast Fourier transform should
C
      exist in the present working directory.
C
C
      The maximum point for time histories is 8192.
c
      The maximum mode number at any period is 400.
c
C
      dimension rr(20), tshif(20), np(20)
      common/sectim/ sec(500)
      common/ctrl/ ms, sr, si, xmom, rO, tO, npO, kkr, kkl, kkf, ino
                     df, bmax, mode, peri, per3(200), xx(20), yy(20)
      common/resp/
      character*50 names
  5
      format(a)
      open(7, file='t1.d', status='scratch', form='unformatted')
      open(8, file='t2.d', status='scratch', form='unformatted')
      write(6, *)
     * 'BE SURE bigfft exist in the present directory.'
      write(6,*) ' '
      write(6, 4)
          'enter the name of input file: (from reigen81)'
       write(6,*) '(if not use, answer none )'
       read(5,5) names
       kkr=O
       if(names, eq. 'none') go to 100
       open(1, file=names, status='old', form='unformatted')
   100 write(6,*)
          'enter the name of input file: (from leigen81)'
       write(6,*) '(if not use, answer none )'
       kk1=0
       read(5,5) names
       if(names, eq. 'none') go to 120
       open(2, file=names, status='old', form='unformatted')
   120 continue
       write(6,*)
          'seismograms(1), spectrum(2), or both(3) wanted?'
     . read(5,*) ityp
       if(itup.eq.2) go to 140
       write(6, *)
```

```
* 'enter the name of output file for seismogram:'
   read(5,5) names
   open(3, file=names, status='new', form='unformatted')
   rewind 3
140 if(itup.eq.1) go to 160
   write(6,*)
       'enter the name of output file for spectrum: '
   read(5,5) names
    open(4, file=names, status='new', form='unformatted')
    rewind 4
   write(6,*) 'what kind of spectrum wanted?'
   write(6, *)
       'all modes(1), fund-high modes separated(2): '
    read(5,*) kkf
160 continue
    write(6,*) 'enter source seismic moment(in 1.e+20): '
    read(5,*) xmom
    write(6,*) 'enter source type (step:1 bell:2 readin:3):'
    read(5,*) ms
    write(6,*) 'enter dt: '
    read(5,*) dt
    if(ms. eq. 2. or. ms. eq. 3) call source(ms. dt)
170 continue
    write(6,*) 'enter station locations- r, tshift, npt: '
    write(6, *)
       '(use r, tshift, -npt if no interpolation wanted.'
    write(6,*) ' use -1,0,0 to stop this sequence. )'
    i ≕ 1
200 continue
    read(5,*) rr(i), tshif(i), np(i)
    if(rr(i), le. 0. 0) go to 260
    if(np(i).1t.0) go to 250
    npt=np(i)
    np t0=1
230 continue
    npt0=2*npt0
    if(npt.eq.nptO) go to 250
    npt1=2*npt0
    if(npt. lt. npt1, and, npt. gt. npt0) go to 240
    go to 230
240 continue
    np(i)=nptO
    write(6,*)
   * 'npt is not in power of 2. adjust to ', np(i)
250 continue
    i = i + 1
    go to 200
260 nsta=i-1
    if(ityp, eq. 1, or, ityp, eq. 3) write(3) dt, nsta, kkr, kkl
    if(ityp.eq.2.or.ityp.eq.3) write(4) dt.nsta.kkr.kkl.kkf
    kO≕O
    zero=0.0
    open(9, file='bigfft.d', status='new', form='unformatted')
```

```
do 400 k1=1, nsta
       k0=k0+1
       ro=rr(k1)
       tOmtshif(k1)
      np0=np(k1)
       in0=1
       if(npO, le. 0) in0=2
       npO=abs(npO)
       df=1./(npO*dt)
       if(ms. eq. 1) go to 350
       generate the source spectrum.
C
       if(k1, ne. 1, and, npO, eq. np(k1-1)) go to 350
       isign=-2
       rewind 9
       write(9) npO,dt,df,isign ✓
       do 300 i=1, np0, 2
       i1 = i + 1
       if(i, le, 500) write(9) src(i), src(i1)
       if(i.gt.500) write(9) zero, zero
  300 continue
       close(9)
       call system('bigfft', kretn)
       open(9, file='bigfft.d', status='old', form='unformatted')
   350 continue
       itp=ityp+2
       if(itup.eq.3) itp=3
       ino=inO
       if(itp.eq.3) ino=1
Ċ
       the main call.
C
       call main(kO, k1, itp)
       if(itup.eq.3) call main(kO,k1,4)
   400 continue
       close(1)
       close(2)
       if(itup.eq. 1. or. itup.eq. 3) close(3)
       if(ityp. eq. 2. or. ityp. eq. 3) close(4)
       close(7, status='delete')
       close(8, status='delete')
       write(6,*) ' '
       write(6,*) 'wig81 finished'
       write(6,*) ' '
       stop
       end
 C
 C
 C
       subroutine main(kO,k1,itp)
       dimension d(80), a(80), b(80), rho(80), qa1(80), qb1(80)
       common/ctrl/ ms, sr, si, xmom, rO, tO, npO, kkr, kkl, kkf, ino
       common/resp/ df,bmax,mode,per1,per3(200),xx(20),yy(20)
       format(/2x, 'M', 5x, 'D', 7x, 'A', 7x, 'B', 6x, 'RHQ', 5x, 'QA',
   10
                  6x, 'QB')
```

```
20
    format(i3,6f8,2)
30
   format('source depth =', f7.2/
              'no. of periods generated in surface81 =', i5)
    rewind 7
    rewind 8
    k0=k0+100
    kk=kkr+kk1
    do 200 k2=1, kk
    icode=k2
    if(kk.eq.1.and.kkl.eq.1) icode=2
    rewind icode
    if(k0*k2.eq.101) write(6,10)
    read(icode)
           mmax, (d(i), a(i), b(i), rho(i), gal(i), gbl(i), i=1, mmax)
    bmax=b(mmax)
    do 100 i=1, mmax
100 if(k0*k2 eq. 101)
           write(6,20) i,d(i),a(i),b(i),rho(i),ga1(i),gb1(i)
    read(icode) nper, dphsrc
    if(kO*k2, eq. 101) write(6,30) dphsrc, nper
    if(k0*k2, eq. 101, or. k0*k2, eq. 201) write(itp) dphsrc, nper ___
    if(k2.eq.1) write(itp) r0,t0,np0
    write(6,*) ' '
    if(itp. eq. 3. and. icode. eq. 1)
       write(6,*) (
                     RAYLEIGH seismogram for station # ', k1,
                   ' is under calculation.'
    if(itp. eq. 3. and. icode. eq. 2)
       write(6,*) /
                       LOVE seismogram for station # ', k1,
                   ' is under calculation.'
   ≯
    if (itp. eq. 4. and. icode. eq. 1)
                     RAYLEIGH spectrum for station # ', k1,
       write(6,*) '
                   ' is under calculation.'
    if(itp.eq.4. and. icode.eq.2)
       write(6,*) ' LOVE spectrum for station # ', k1,
                   ' is under calculation.'
    the main routine.
    if(ino.eq.1) call inter(itp.icode)
    if(ino.eq.2) call noint(itp,icode)
200 continue
    rewind 7
    rewind B
    go to (300,500), itp-2
300 continue
    if(kkr.eq.1) read(7) per,(xx(i), i=1,16)
    if(kkl.eq.1) read(8) per,(xx(i),i=17,20)
    write(3) (xx(i), i=1,20)
    if(per. le. O. O) go to 600
    go to 300
500 continue
    if(kkr.eq.1) read(7) per,(xx(i),i=1,16)
    if(kkl, eq. 1) read(B) per,(xx(i), i=17,20)
```

C

C

```
write(4) per,(xx(i), i=1,20)
      if(kkf.eq. 1. and. per. lt. O. O) return
      if(kkf.eq. 1) go to 500
      if(kkr.eq.1) read(7) per,(xx(i), i=1,16)
      if(kkl.eq.1) read(8) per.(xx(i),i=17,20)
      write(4) per,(xx(i), i=1,20)
      if(per. le. 0. 0) go to 600
      go to 500
  600 continue
      return
      end
C
C
C
      subroutine inter(itp,icode)
      This routine sets the end values for interpolation.
C
      common/dimen/ ur(200,2), dur(200,2), ur(200,2), dur(200,2),
                  wyno(200,2), ur0(200,2), are(200,2), gamma(200,2)
                     ms, sr, si, xmom, rO, tO, npO, kkr, kkl, kkf, ino
      common/ctrl/
                     df, bmax, modeO, per1, per3(200),
      common/resp/
                       xr(16), xl(4), yr(16), yl(4)
      if(ms.eq.1) go to 90
      rewind 9
      read(무) npx,dtx,dfx,isign
  90
      continue
      np2=np0/2+1 .
      icode1=icode+6
      read(icode) ifunc.mode2.per2
      if(ifunc.lt.O) write(6,*) 'no data in the eigen file.'
      do 200 j=1, mode2
      read(icode) wyno(j,2),urO(j,2),are(j,2),ugr,gamma(j,2)
      read(icode) ur(j,2), dur(j,2), uz(j,2), duz(j,2)
  200 continue
      itrig=0
      -do 800 i=np2,1,-1
      do 300 j=1,16
      xr(j)=0.0
      yr(j)=0.0
  300 continue
      do 310 j=1,4
      x1(j)=0.0
      u1(j)=0.0
  310 continue
      per=-1.0
      if(i.eq.1) go to 750
      if(ms.ne.1) read(9) sr.si
      per=1. /((i-1)*df)____
      if(i.eq.2) go to 320
      dper=1. /((i-2)*df)-per
  320 continue
      perO=per-O.005*dper
      if(i, eq. np2) per0=per+0, 005*dper
  330 if(per0.1t.per2) go to 400
```

```
if(itrig.eq.2) go to 750
     per1=per2
     mode1=mode2
     read(icode, end=750) ifunc, mode2, per2
     itriq=1
      if(ifunc.gt.O) go to 340
     itrig=2
     qo to 330
 340 continue
     do 350 j=1, mode1
     wvno(j,1) = wvno(j,2)
     urO(j,1) = urO(j,2)
      are(j,1) = are(j,2)
      gamma(j, 1) = gamma(j, 2)
      ur(j, 1)
                = ur(j,2)
                = dur(1,2)
      dur (j. 1)
      uz (j. 1)
               = uz(j,2)
                = duz(j,2)
      duz (j. 1)
 350 continue
      do 360 j=1, mode2
      Sreq=(1)=per2
      read(icode) wyno(j,2),urO(j,2),are(j,2),ugr,gamma(j,2)
      read(icode) ur(j,2),dur(j,2),uz(j,2),duz(j,2)
 360 continue
      go to 330
  400 if(itrig.eq.0) go to 750
      mod=mode1-mode2
      modeO≕mode2
 700 continue
      kk = itp-2
      if(itp. eq. 4. and. kkf. eq. 1) kk=4
      go to (710,720), icode
  710 call excitr(per/kk)
      if(itp.eq.4.and.kkf.eq.2) call excitr(per,3)
      go to 750
  720 call excitl(per/kk)
      if(itp.eq.4. and. kkf.eq.2) call excit1(per,3)
  750 continue
      go to (760,770), icode
  760 write(7) per/xr
      if(itp.eq.4. and. kkf.eq.2) write(7) per/yr
      go to 800
  770 write(8) per,xl
      if(itp.eq.4.and.kkf.eq.2) write(8) per/yl
_800 continue
      return
      end
C
C
C
      subroutine noint(itp,icode)
      common/dimen/ ur(200,2), dur(200,2), uz(200,2), duz(200,2),
                  wynn (200, 2), uro (200, 2), are (200, 2), gamma (200, 2)
```

```
ms, sr, si, xmom, rO, tO, npO, kkr, kkl, kkf, ino
    common/ctr1/
                   df, bmax, mode, per1, per3(200),
    common/resp/
                     xr(16), xl(4), yr(16), yl(4)
    icode1=icode+6
300 continue
    do 310 j=1,16
    xr(j)=0.0
    ur(j)=0. Q
310 continue
    do 320 j=1,4
    x1(j)=0.0
    q1(j)=0.0
320 continue
    read(icode, end=900) ifunc, mode, per
    if(ifunc.le.O) per=-1.O
    if(ifunc. le. 0) go to 750
    abomit=1, OCE ob
    read(icode) wyno(j,1),urO(j,1),are(j,1),ugr,gamma(j,1)
    read(icode) ur(j,1),dur(j,1),uz(j,1),duz(j,1)
    wyno(j,2)=wyno(j,1)
    ur0(j,2)=ur0(j,1)
    are(j, 2) =are(j, 1)
    gamma(j, 2) = gamma(j, 1)
    ur(j, 2)=ur(j, 1)
    dur(1,2)=dur(1,1)
    uz(j,2)=uz(j,1)
    duz(j, 2) = duz(j, 1)
330 continue
    if(ms. eq. 1) go to 650
    no2=no0/2+1
    per1=1./((np2-1)*df)
    if(per.1t.per1) go to 600
    rewind 9
    read(9) npy, dtx, dfx, isign
    read(9) sr1, si1
    j=np2-2
400 continue
    if(j.eq.0) go to 600
    per2=1. /(j*df)
    read(9) sr2, si2
    if(per.gt.per2.and.per.le.per1) go to 500
    per1=per2
    sr2=sr1
    si2#si1
     j=j-1
    go to 400
500 sr=sr1+(per-per1)/(per2-per1)*(sr2-sr1)
    si=si1+(per-peri)/(per2-peri)*(si2-si1)
    go to 650
600 write(6,*) 'The period range of source pulse spectrum'
    write(6,*) 'does not cover the period ≈', per
    write(6,*) 'Now r=',rO,' npt=',npO,' dt=',dtx
    write(6,*) 'and the period range is ',1./((np2-1)*df),
```

```
1-1, 1. /df
     write(6,*) 'TRY DIFFERENT npt, dt. '
     stop
 650 continue
      kk=itp-2
      if(itp. eq. 4. and. kkf. eq. 1) kk=4
      go to (710,720), icode
 710 call excitr(per/kk)
      if(itp.eq.4.and.kkf.eq.2) call excitr(per.3)
      go to 750
 720 call excitl(per/kk)
      if(itp.eq.4.and.kkf.eq.2) call excitl(per.3)
 750 continue
      go to (760,770), icode
 760 write(7) per/xr
      if(itp.eq. 4. and. kkf.eq. 2) write(7) per,yr
      go to 800
 770 write(8) per,xl
      if(itp. eq. 4. and. kkf. eq. 2) write(8) per/yl
 800 continue
      go to 300
  900 continue
      return
      end
C
C
C
      subroutine excitr(per,ident)
      This routine generates the Z and R components of
C
      seismogram for
C
       1: 45-deg dip-slip source 2: strike-slip source
c
                                     4: explosion source.
       3: dip-slip source
C
c
      The number 200 in dimension declaration limits the
C
        number of mode at a particular frequency not over 200.
C
C
      dimension dk(4), dkk(4), vz(2,4), vr(2,4)
      common/dimen/ ur1(200,2), dur1(200,2), ur1(200,2),
                        duz1(200,2), wyno1(200,2), ur01(200,2),
      4
                        are1(200,2), gamma(200,2)
                      ms, srO, siO, xmom, rx, tx, npx, kkr, kkl, kkf, ino
       common/ctrl/
                      df, bmax, mode, per1, per3(200), xz(2, 4),
       common/resp/
                        xr(2, 4), xt(2, 2), yz(2, 4), yr(2, 4), yt(2, 2)
       do 90 i=1/2
       do 90 j=1,4
       \forall z(i,j)=0.0
       \forall r(i,j)=0.0
  90
      continue
       X T = X X T
       spectra normalized to a distance of 1000 km.
C
       if(ident.ge.2) rxx=1000.0
       J1=1
       if(ident, eq. 3) j1=2
```

```
J2≕mode
if(ident.eq.2) j2=1
if(j1.gt.j2) return
do 200 j≐j1,j2
rat=0.0
if(ino.eq.1) rat=(per-per1)/(per3(j)-per1)
wyno=wyno1(j,1) + (wyno1(j,2)-wyno1(j,1))*rat
atn=gamma(j,1)*rx
fact1=0.0
if(atn. lt. 80. 0) fact1=1. /exp(atn)
atn=gamma(j,2)*rx
fact2=0.0
omega=6.2831853/per
vi == arei(j, i)/sqrt(wvnoi(j, i)*rxx)
v2 = are1(j,2)/sqrt(wvno1(j,2)*rxx)
w1 = duz1(j, 1)+0.5*wvno1(j, 1)*ur1(j, 1)
\omega 1 = \omega 1 * v 1 * fact 1
u1 = u1*urO1(j,1)
w2 = duz1(j,2)+0.5*wvno1(j,2)*ur1(j,2)
w2 = w2*v2*fact2
u2 = u2*ur01(j,2)
dk(1) = w1 + (w2 - w1) * rat
dkk(1) = u1 + (u2 - u1) * rat
w1 = w \lor no1(j, 1) * ur1(j, 1)
w1 = w1*v1*fact1
u1 = ui*urO1(j,i)
w2 = wvno1(j,2)*vr1(j,2)
mS = mS*AS*tactS
u2 = u2*urO1(j,2)
dk(2) = \omega_1 + (\omega_2 - \omega_1) * rat
dkk(2) = u1 + (u2 - u1) * rat
w1 = w \vee noi(j, 1) * uzi(j, i) + duri(j, 1)
w1 = w1*v1*fact1
u1 = wi*urOi(j,i)
w2 = wvno1(j, 2)*v21(j, 2)*dur1(j, 2)
w2 = w2*v2*fact2
u2= w2*ur01(J,2)
dk(3) = \omega 1 + (\omega 2 - \omega 1) * rat
dkk(3) = u1 + (u2 - u1) * rat
\omega 1 = \operatorname{dur1}(j, 1) - \operatorname{wvno1}(j, 1) * \operatorname{ur1}(j, 1)
w1 = w1*v1*fact1
ui = ui*urOi(j,1)
\omega 2 = dur1(j, 2) - \omega v no1(j, 2) * ur1(j, 2)
w2 = w2*v2*fact2
u2 = w2*ur01(j,2)
 dk(4) = \omega 1 + (\omega 2 - \omega 1) * rat
 dkk(4)= u1+(u2-u1)*rat
 t0=xmom/2.5066283
 ti=omega*tx-wyno*rx-0.7853981632
 t2=omega*tx-wvno*rx-2.356194489
 ct1=cos(t1)
 st1=sin(t1)
```

```
ct2=cos(t2)
      st2=sin(t2)
      do 100 k=1,4
      ct=cti
      if(k,eq.3) ct=-st1
      st=st1
      if(k, eq. ③) st=ct1
      V_2(1, k) = V_2(1, k) - dk(k) *tO*ct
      v_{z}(2,k)=v_{z}(2,k) - dk(k)*t0*st
      ct=ct2
      if(k, eq. 3) ct=-st2
      st=st2
      if(k,eq.3) st=ct2
      \forall r(1,k) = \forall r(1,k) + dkk(k) * tO * ct
      \forall r(2,k) = \forall r(2,k) + dkk(k) * tO * st
 100 continue
 200 continue
      if(ms.eq.1) go to 220
      sr=srO
      si=si0
      do to 240
 220 sr=0.0
      si = -per/6.2831853
 240 continue
      do 300 k=1,4
      tQ=vz(1,k)
      ∨2(1, k)=sr*∨2(1, k)-si*∨2(2, k)
      \forall z (2, k) = sr*\forall z (2, k) + si*t0
      t0=vr(1, k)
      \forall r(1,k) = sr \neq \forall r(1,k) - si \neq \forall r(2,k)
      \forall r(2,k) = sr*\forall r(2,k) + si*t0
  300 continue
      do 400 i=1.2
      do 400 j=1,4
      if(ident.ne,3) xz(i,j)=vz(i,j)
      if(ident.ne.3) xr(i,j)=vr(i,j)
      if(ident.eq.3) yz(i,j)=vz(i,j)
      if(ident.eq.3) yr(i,j)=vr(i,j)
  400 continue
      return
       end
C
C,
C
       subroutine excitl(per,ident)
       This routine generates the T component of seismograms
C
                                      2: strike-slip source.
       for 1: dip-slip source
C
C
                                                 j≖1,2
                                                          SOUTCE
       vt(i,j) i=1,2 real & imag part
C.
       dimension dk(2), vt(2,2)
       common/diman/ ut1(200,2), dut1(200,2), dr1(200,2),
                          duz1(200,2), wyno1(200,2), ur01(200,2),
```

```
ale1(200,2),gamma(200,2)
                     ms, srO, siO, xmom, rx, tx, npx, kkr, kkl, kkf, ino
    common/ctr1/
                     df, bmax, mode, per1, per3(200), xx(16),
    common/resp/
                        xt(2,2), qq(16), qt(2,2)
    do 90 i=1,2
    do 90 j=1,2
    vt(i, j) = 0.0
90
    continue
    צידדצ
    if(ident.ge.2) rxx=1000.0
     j1 = 1
    if(ident.eq.3) J1=2
     J2=mode
    if(ident.eq.2) j2=1
     if(j1.gt.j2) return
     do 100 j=j1, j2
    rat=0.0
     if(ino, eq. 1) rat=(per-per1)/(per3(j)-per1)
    wyno=wynoi(j,1) + (wynoi(j,2)-wynoi(j,1))*rat
     atn=gamma(j,1)*rx
     fact1=0.0
     if(atn. 1t. 80. 0) fact1=1. /exp(atn)
     atn=gamma(j,2)*rx
     fact2=0.0
     if(atn. lt. 80.0) fact2=1. /exp(atn)
     omega=6.2831853/per
     v1 = als1(j,1)/sqrt(wvno1(j,1)*rxx)
     \sqrt{2} = alei(j,2)/sqrt(wvnoi(j,2)*rxx)
     wi = dut1(j, 1)*vi*fact1
     w2 = dut1(j, 2)*v2*fact2
     dk(1) = w1 + (w2 - w1) * rat
     w1 = w \vee no1(j, 1) * vt1(j, 1)
     w1 = w1*v1*fact1
     w2 = wvno1(j, 2)*ut1(j, 2)
     w2 = w2*v2*fact2
     dk(2) = \omega 1 + (\omega 2 - \omega 1) * rat
     t0=xmom/2,5066283
     ti=omega*tx-wyno*rx+0.7853981632
     cti=cos(ti)
     st1=sin(t1)
     \forall t(1,1) = \forall t(1,1) + dk(1) * tO * st1
     \forall t(2,1) = \forall t(2,1) - dk(1) * t0 * ct1
     \forall t(1,2) = \forall t(1,2) + dk(2) * t0 * ct1
     \forall t(2,2) = \forall t(2,2) + dk(2)*t0*st1
100 continue
     if(ms. eq. 1) go to 120
     sr=sr0
     si=si0
     go to 140
120 sr=0.0
     si = -per/6.2831853
 140 continue
     do 200 k=1,2
```

```
t0=∨t(1, k)
      \forall t(1,k) = sr * \forall t(1,k) - si * \forall t(2,k)
      \forall t(2,k) = sr* \forall t(2,k) + si*tO
  200 continue
      do 300 i=1,2
      do 300 j=1,2
      if(ident, ne, 3) \times t(i, j) = \forall t(i, j)
      if(ident.eq.3) yt(i,j)=vt(i,j)
  300 continue
      return
      end
C,
      subroutine source(ms, dt)
      enter source time function.
C
C
      common/srctim/ src(500)
      character*50 names
      do 50 i=1,500
      src(i)=0.0
  50 continue
      go to (100,200), ms-1
  100 write(6,*) 'enter tl for bell type of source'
      read(5,*) tl
      call pulse(src, 500, dt, t1)
      return
  200 write(6,*)
     * 'input source time function from here(1) or file(2)?'
      read(5,*) iO
      if(i0.eq.1) go to 250
      write(6,*) 'enter file name storing source function: '
      read(5,5) names
  5
      format(a)
      open(9, file=names, status='old', form='formatted')
      J== 1
  220 continue
      read(9, *, end=230) src(j)
      1=1+1
      if(j.eq.500)
     * write(6,*) 'source timeseries too long! <500'
      if(j.eq.500) go to 230
      go to 220
  230 continue
      close(9)
      return
  250 continue
      write(6, *)
          'enter source time function for every (i-1)*dt:'
      write(6,*) '(use -1000.0 to stop)'
      1=1
  260 continue
```

```
read(5, 4) src(j)
      1=1+1
      if(src(j), 1t, -999, 0) go to 280
      go to 260
  280 src(j)=0.0
      return
      end
C
C
      subroutine pulse(f,n,dt,tl)
      give a source time function
C
C
      dimension f(1)
      t1 = 0.0
      t2 = t1 + t1
      t3 = t2 + t1
      t4 = t3 + t1
      t5 = t4 + t1
      do 100 i = 1, n
      y = (i-1)*dt
      z = y - t1
      f(i) = 0.0
      if(y.gt.t1) go to 101
      go to 100
  101 if(y.gt.t2) go to 102
      f(i) = (z/t1)*(z/t1)*0.5
      go to 100
  102 if(y.gt.t3) go to 103
      f(i) = -(z/t1)*(z/t1)*0.5 + 2.0*(z/t1) - 1.0
      go to 100
  103 if(y.gt.t4) go to 104
      f(i) = -(z/t1)*(z/t1)*0.5 + 2.0*(z/t1) - 1.
      90 to 100
  104 if(y.gt.t5) go to 105
      f(i) = (z/t1)*(z/t1) * 0.5 -4.0 * (z/t1) + 8.0
      go to 100
  105 f(i) = 0.0
  100 continue
      area of pulse normalized to unity
      do 200 i = 1, n
  200 f(i) = f(i) /(2.*t1)
      しゃさりしてい
      end
****
```

```
gle81. f
Ċ
C
      f77 -i -I2 gle81 f -o gle81 -lcalcompI2
C
      This program follows the program wig81 and generates
      the seismogram after passing through an instrument.
C
r
      The instrument response is imposed in the frequency
      domain.
C
C
      The source methanism should be given here.
C
C
      The dimension needed has been reduced to the minimum,
C
      and time series point can be as large as 8192.
C
      The restriction for this point dimension comes from
C
      the FFT program 'bigfft', which should exist in
C
      the present working directory.
C
      -Oct 10, 1981.
C
      dimension (15), ar(100), bar(100)
      complex data(10), rr, tt, ff
      common/dimen/ x(1026), y(1026)
      common/pltses/ tO, yend, y1, iplt, jctrl, bb, cc
      common/keep/ s1,s2,s3,t1,t2,cos1,sin1,df
      character*2 form(5),bb,cc
      character*50 names
      data form/'Z ','R ','SH','NS','EW'/
      format(a)
     format(/1x, a2, ' component for station at R=', f8. 1,
  10
                        ' Az=', f6.1,' is being processed.')
      format(/1x, 'ins=', i2, 9x, 'peak=', f7.1, ' iresp=',
  20
         i2/' dt= ',f9.5,' ndist= ',i2,7x,'dphsrc= ',f6.1/
     -8-
              ' Z, R, SH, N, E:
                             ',5i2,5x,' Quake,Expl: ',i2/
            ' dip= ', f5, 1, 5x, 'slip= ', f5, 1, 5x, 'strk= ', f5, 1}
  30
     format(a2,1x,'component')
      write(6,*) ' '
      write(6,*) 'BE SURE ',
     * 'the program bigfft exist in the present directory,'
      write(6,*)
        'and not be used by other similar jobs.'
      write(6,*) ' '
      write(6,*) 'calculate only(1),',
     * 'plot only(2), or calculate and plot(3): '
      read(5,*) iplt
      write(6,*) 'enter the input file name: '
      read(5,5) names
      open(1, file=names, status='old', form='unformatted')
      rewind 1
      write(6,*) ' '
      write(6,*) 'store the seismogram data: (y/n)'
      read(5,5) bb
      if(bb.eq.'n') go to 40
```

```
write(6, *)
    * 'data (x,y) will be stored in the format of (2e15.8). '
     write(6, *)
     * 'enter the file name for storing seismogram data: '
      read(5,5) names
      open(9, file=names, status='new', form='formatted')
      rewind 9
 60 continue
      if(iplt.eq.2) go to 70
      open(2, file='bigfft.d', status='new',
                         form='unformatted')
      open(4, file='tmpp.d', status='scratch',
                            form='unformatted')
      if(iplt, ne. 1) go to 70
      write(6,*) ' '
      write(6,*) 'enter output file name '.
           'for storing the ploting data: '
      read(5,5) names
      open(3, file=names, status='new', form='unformatted')
      rewind 3
  70 continue
c-120 input control parameters.
C
      jctrl=-1
      if(iplt.eq.1) go to 80
      call plots(0,0,7)
      write(6,*) 'original pen move: (1.8,10.3)'
      read(5,*) x1, q1
      call plot(x1,y1,-3)
      qend=q1
      ipen=1
      write(6,*) 'choose the pen:'
      read(5,*) ipen
      call newpen(ipen)
      if(iplt. eq. 2) go to 140
  80
      continue
      write(6,*) 'choose the type of instrument'
      write(6,*) '(=0 15-100 WWSSN
                                        =1 30-100 WWSSN'
      write(6,*) / =2 6824-13 LP SYSTEM =3 6824-2 LP SYSTEM'
      write(6,*) ' =4 WWSSN SP
                                         =5 USGS SP '
      write(6,*) ' =6 CSSN SP'
      write(6,*) '=-1 do NOT pass through any instrument ) '
      read(5,*) ins
      write(6, *)
          'enter the peak value for instrument response: '
      read(5,*) peak
      write(6, *)
              'Take integration(-1), derivative(1), or not(0):'
      read(5,*) iresp
      read(1) dt, ndist, kkr, kkl
      kk=kkr+kk1
      read(1) dphsrc, nper
```

```
write(6,*) ' '
     if(kk.eq.2) write(6,*)
          'Both Rayleigh and Love eigens been generated.'
     if(kkr.eq. 1. and. kkl.eq. 0) write(6,*)
          'Only Rayleigh eigens been generated.'
     if(kkr.eq.O. and.kkl.eq.1) write(6,*)
          'Only Love eigens been generated.'
     write(6,*) ′ ′
     write(6,*)
           'Which components to be plotted? (Z,R,SH,NS,EW)'
     write(6,*)
    st 'if yes answer 1, no answer 0 (e.g. 1,0,0,1,1)'
     read(5,*) (jj(i), i=1,5)
      if(kkr. ne. 0) go to 100
      j=jj(1)+jj(2)
      if(j.eq.0) go to 100
      write(6,*) 'no P-SV wave ready, run again.'
      go to 1000
  100 continue
      if(kkl.ne.0) go to 120
      if(jj(3), eq. 0) go to 110
      write(6,*) 'no SH wave ready, run again.'
      go to 1000
  110 continue
· 120 continue
c-150 input source mechanism.
C
      write(6,*) ' '
      write(6,*)
          'no. of epicentral distance generated = ', ndist
      write(6,*) 'source depth = ',dphsrc,' dt = ',dt
      write(6,*) '
      write(6,*) 'earthquake source(1) or explosion(2)?'
      read(5, *) m3
      degrad=3.141592653/180.0
      if(m3.eq.2) go to 130
      write(6,*) 'enter the source mechanism dip.slip.strk:'
      read(5,*) dipO, slipO, strkO
  130 continue
C
      if(iplt.eq. 1) write(3) ins, peak, iresp, dt, ndist, dphsrc,
                                jj, m3, dipO, slipO, strkO
      go to 150
  140 continue
      read(1) ins.peak.iresp.dt.ndist.dphsrc.jj.m3.dipO.
                 slipO, strkO
      write(6,20) ins.peak.iresp.dt.ndist.dphsrc.jj.m3.
                   dipO, slipO, strkO
  150 continue
c-900 the main loop.
```

```
iazm=0
      iaz=-1
C
      stations at different epicentral diatances.
C
      do 900 kdist=1,ndist
      iaz=iaz+1
      if(iplt.eq.2) go to 210
      read(1) r0, t0, np0
      write(6,*) ' '
      write(6,*) 'r=',r0,' t0=',t0,' np0=',np0
      write(6,*) ' '
      np2=np0/2+1
      df=1.0/(npO*dt)
      if(iaz. lt. iazm) go to 180
      iaz=-1
      write(6,*) 'enter Az,bAz (in degree): '
      write(6,*) '(use bAz=0.0 to set bAz=Az+180.0'
      write(6,*) ' use -1,-1 to stop)'
      k = 1
  160 continue
      read(5, *) az(k), baz(k)
      if(az(k), eq. -1, 0, and, baz(k), eq. -1, 0) go to 170
      if(baz(k), eq. 0.0) baz(k)=az(k)+180.0
      if(baz(k), ge, 360, 0) baz(k)=baz(k)-360, 0
      k=k+1
      go to 160
  170 naz=k-1
      write(6,*)
     * 'how many sets of different distances FOLLOWED will use'
      write(6, *)
     * 'the above azimuth data: (if not know, enter 100)'
       read(5, *) iazm
  180 continue
      write(6,*) ' '
      write(6,*) 'wait.'
       rewind 4
       do 200 k=np2,1,-1
       read(1) (data(i), i=1, 10)
       write(4) (data(i), i=1, 10)
  200 continue
       if(iplt.eq.1)
      * write(3) rO, tO, npO, naz, (az(k), baz(k), k=1, naz)
  210 continue
       if(iplt.eq.2)
      * read(1) r0, t0, np0, naz, (az(k), baz(k), k=1, naz)
       n1=npO/1024
       nm=npO-n1*1024
       n1=n1+1
       if(nm, eq. 0) n1=n1-1
C
       stations at different azimuths.
C
       do 900 kaz=1, naz
       phiO=ar(kar)
```

```
phil=bar(kar)
      if(iplt.eq.2) go to 310
      phi=phil*degrad
      cos1=cos(phi)
      sin1=sin(phi)
      if(m3.eq.2) go to 300-
c-300 dislocation source.
      strk1=phiO-strk0
      dip=dipO*degrad
      slip=slipO*degrad
      strk1=strk1*degrad
      strk2=2. *strk1
      dip2=2. *dip
      s1=sin(slip)*sin(dip2)
      s2=-(cos(slip)*sin(dip)*sin(strk2)+
             O. 5*sin(slip)*sin(dip2)*cos(strk2))
      s3=sin(slip)*cos(dip2)*sin(strk1)-
         cos(slip)*cos(dip)*cos(strk1)
      t1=sin(slip)*cos(dip2)*cos(strk1)+
         cos(slip)*cos(dip)*sin(strk1)
      t2=cos(slip)*sin(dip)*cos(strk2)-
            O. 5*sin(slip)*sin(dip2)*sin(strk2)
  300 continue
  310 continue
c-500 store the spectrum data in bigfft.d
C
       different components at the same station.
C
       do 900 j=1,5
       if(jj(j).eq.O) go to 900
       cc=form(1)
       write(6,10) cc,r0,phi0
       if(iplt.eq.2) go to 650
       rewind 2
       rewind 4
       isign=+2
       write(2) npO, dt, df, isign
       do 500 k=np2, 2, -1
       read(4) (data(i), i=1, 10)
       freq=(k-1)*df
       go to (400,320),m3
   320 go to (330,340,350,360,370),j
   330 ff=data(4)
       go to 460
   340 ff≕data(8)
       go to 460
   350 write(6,*) 'explosive source no SH wave. run again.'
       go to 1000
   360 ff=-cos1*data(8)
       go to 460
   370 ff=-sin1*data(8)
```

```
go to 460 -
 400 go to (410, 420, 430, 440, 440), j
 410 ff=s1*data(1)+s2*data(2)+s3*data(3)
      go to 460
 420 ff=s1*data(5)+s2*data(6)+s3*data(7)
      go to 460
 430 ff=t1*data(9)+t2*data(10)
      go to 460
 440 rr=s1*data(5)+s2*data(6)+s3*data(7)
      tt=t1*data(9)+t2*data(10)
      cti=-cosi
      st1=sin1
      if(j.eq.4) go to 450
      ct1 = -sin1
      st1=-cosi
 450 ff=ct1*rr+st1*tt
  460 continue
      call resp(ff, freq, ins, peak, iresp)
      write(2) ff
  500 continue
      ff=cmp1x(0,0,0,0)
      write(2) ff
C
      system call: perform big fft
C .
      close(2)
      call system('bigfft', kturn)
      open(2, file='biqfft.d', status='old',
                              form='unformatted')
      rewind 2
      read(2) npx, dtx, dfx, isx
c-600 find overall max/min
      y_{max} = -1.0e + 38
      ymin= 1 0e+38
      do 600 k=1, np0, 2
      read(2) u1,u2
      if(y1.gt.ymax) ymax≕y1
      if(y2.gt.ymax) ymax=y2
      if(y1.lt.ymin) ymin=y1
      if(y2.1t.ymin) ymin=y2
  600 continue
       xmin=0.0
      xmax=(npO-1)*dt
       if(iplt.eq.1) write(3) xmin, xmax, ymin, ymax
      rewind 2
      read(2) npx,dtx,dfx,isx
       if(iplt. eq. 1) go to 660
  650 continue
       if(iplt.eq.2) read(1) xmin, xmax, ymin, ymax
       plot axis
  660 continue
       call seiplt(xmin, xmax, ymin, ymax, xmin1, xmax1)
```

```
if(bb.eq.'n') go to 680
      write(9,30) cc
      write(9, *)
       'dt=',dt,' nn=',npO,' from:',xmin1,' to ',xmax1,' sec'
      if(m3. eq. 1) write(9,*)
       'source depth=',dphsrc,' dip=',dipO,' slip=',
         slipO, ' strk=', strkO
      if(m3. eq. 2) write(9, *)
        'source depth=', dphsrc, ' explosion source'
      write(9, *)
          'r=', r0, ' phi=', phi0, ' t0=', t0, ' instrument=', ins,
          ' peak=', peak
  680 continue
c
c-800 divide the whole length into sub-lenth with 1024 as max.
      do 800 k=1, n1
      k 0 == k
      n0=1024
      if(k, eq, n1, and, nm, ne, O) n0=nm
      go to (700,750,700), iplt
  700 do 720 i=1,n0,2
      read(2) y(i), y(i+1)
  720 continue
      if(ip1t.eq. 1) write(3) (y(i), i=1, n0)
      go to 760
  750 read(1) (y(i),i=1,n0)
  760 continue
      call seical(kO, n1, nO, dt, rO, xmin1, xmax1)
  800 continue
  900 continue
 1000 continue
      if(iplt.ne.1) call plot(8.0,0.0,999)
      close(i)
      close(2)
      close(3)
      close(4, status='delete')
      close(9)
      write(6,*) ' '
      write(6,*) 'gle81 finished'
      write(6,*) ' '
      stop
      end
C
C
C
      subroutine resp(ff, freq, ins, peak, iresp)
      This routine imposes the instrument response on the
C
      seismogram by frequency domain multiplication.
C
      complex ff
      fr=real(ff)
      fi≃aimag(ff)
      if(ins. lt. 0) go to 250
```

```
= 1 WWSSN 30-100
      ins = 0 WWSSN 15-100
С
                                      = 3 LP SYSTEM 6824-2
               LP SYSTEM 6824-13
C
                                      == 5
                                           USGS SP
               WWSSN SP
          = 4
C
          = £
                CSSN SP
C
      go to (110, 110, 120, 120, 130, 130, 130), ins+1
  110 call wwssn(freq.peak,ins,pr,pi)
      go to 200
  120 call lpsys(freq.peak,ins,pr,pi)
      go to 200
  130 call sperd(freq.peak,ins.pr.pi)
  200 continue
      tmp=fr
      fr=fr*pr-fi*pi
      fi=tmp*pi+fi*pr
  250 if(iresp. eq. 0) go to 400
      omega=2. *3, 141592653*freq
      if(iresp. eq. -1) go to 300
      tmp=fr
       fr≕-fi∻omega
      fi≕tmp*omega
       go to 400
  300 continue
       tmp=fr
       fr=fi/omega
       fi=-tmp/omega
  400 ff=cmplx(fr,fi)
       return
       end
C
C,
       subroutine www.sn(freq.peak.ins.xr,xi)
                15-100 WWEEN
       ins eq 0
C
                 30-100 WWSSN
       ins eq 1
c
       peak magnifications are 350, 420, 1400, 2800, 5600
 c
       we=6.2831853*freq
       index=(peak+1)/375
       if(ins.gt.O) go to 200
   100 continue
       go to(1,2,2,3,3,3,4,4,4,4,4,4,4,4,5), index
     1 fmag=278.
       sigma=0.003
       go to 6
     2 fmag=556.0
       sigma=0.013
       go to 6
     3 fmag=1110.
       sigma=0.047
       go to 6
     4 fmag=2190.
       sigma=0.204
       go to 6
     5 fmag=3950.
```

```
sigma=0.805
  6 zeta=0.93
    reta1=1.
    wn = .418879
    wn1 = .062831853
    go to 300
200 continue
    ge to(10, 20, 20, 30, 30, 30, 30, 40, 40, 40, 40, 40, 40, 40, 40, 50),
         index
 10 \text{ fmag} = 251.9
    sigma = 0.003
    go to 60
 20 \text{ fmag} = 503.1
    sigma = 0.012
    go to 60
 30 \text{ fmag} = 1001.5
    sigma = 0.044
    go to 60
 40 fmag = 1941.9
    sigma = 0.195
    go to 60
 50 \text{ fmag} = 2241.8
    sigma # .767
 60 \text{ zeta} = 1.5
    zeta1 = 1.0
    wn = .2094395
    wn1 = .062831853
300 continue
    ar= (we*we-wn*wn)*(we*we-wn1*wn1)-4. *reta*reta1*wn*wn1*
         (1, -sigma)*we*we
    ai=2.*we*(retai*wn1*(wn*wn-we*we)+reta*wn*
        (wn1*wn1-we*we))
    factor = fmag*we*we*we / (ai*ai + ar*ar)
    xr =-ai * factor
    xi =-factor * ar
    return
    end
    subroutine lpsys(freq,peak,ins,xr,xi)
    LRSM response for LP system with filter 6824-2 ins=3
    LRSM response for LP system with filter 6824-13 ins=2
    phase response obtained from hilbert transform of amplitude
    response. gain normalized to 1.0 at 25 seconds.
    dimension fre(28), p(56), phi(56)
     the first 28 p and phi are for 6824-2, next
    28 are for 6824-13
     data fre/.001,.002,.003,.004,.005,.006,.007,.008,.009,
    1, 01, , 02, , 03, , 04, , 05, , 06, , 07, , 08, , 09, , 1, , 2, , 3, , 4, , 5, , 6,
   2.7, 8, 9, 1.7
     data phi/263. 9, 257. 9, 251. 9, 245. 6, 239. 3, 233. 1, 227. 1, 221. 1,
    1214, 6, 208, 3, 134, 9, 73, 2, 15, 3, -33, 2, -71, 2, -100, 1, -122, 4,
```

c c

C

C

C

C

C

```
2-140.0,-153.8,-213.1,-232.6,-242.8,-249.2,-253.0,-256.9,
    3-259. 3, -261. 2, -262. 9, 265. 0, 259. 1, 253. 4, 248. 1, 242. 8, 236. 9,
    4 230. 3, 223. 4, 216. 4, 209. 7, 153. 4, 102. 0, 54. 6, 14. 4, -16. 8,
    5 -40, 4, -58, 2, -73, 3, -85, 4, -146, 4, -153, 9, -155, 3, -156, 4,
    6-157, 7, -159, 2, -160, 0, -162, 3, -163, 8/
     data p/.00005,.00040,.00135,.00321,.00625,.01077,.01706,.02546,
    1. 03631, . 05006, . 34117, . 73904, 1. 0000, . 97633, . 79807, . 61417, . 46105,
    2.34464,.26315,.03583,.01097,.00468,.00240,.00139,.00087,.00058,
     3, 000407, 000297
    4 .000015,.000391,.00131,.00310,.00609,.01068,.01713,.02556,
    5 . 03618, . 04848, . 28030, . 67553, 1. 0, 1. 09448, 1. 0044, . 86969, . 74511,
    6 . 63557, . 54818, . 12153, . 04830, . 02741, . 01827, . 01328, . 01017,
     7 .008052 .006545 ..005404/
      m = 28
      degrad = 0.01745329
      if(freq. gt. 0.5) freq = 0.5
      if(freq. 1t. 0.005) freq = 0.005
      do\ 200\ i\ =\ 1.m
      if(freq.ge.fre(i).and.freq.le.fre(i+1)) go to 160
 200 continue
 160 continue
      if(ins.eq.2) j=i+28
      if(ins.eq.3)_{j=1}
      pf= phi(j)+(phi(j+1)-phi(j))/(fre(i+1)-fre(i))*(freq-fre(i))
      ph = p(j) + (p(j+1) - p(j))/(fre(i+1)-fre(i))*(freq-fre(i))
      pf = pf * degrad
      xr = cos(pf) * ph * peak
      xi = sin(pf) * ph * peak
      return
      end
C
C
C
      subroutine sperd(freq.peak.ins.xr.xi)
                WWSSN SP
      ins = 4
C
      ins = 5
                USGS
                       SP
C
                       SP
                CSSN
      ins = 6
C
      The values for WWSSN and USGS from Luh (1977, BSSA, p. 950)
C
r
      double complex ss(20), www.tt
      double precision xnorm, a(20)
      go to (100,200,300), ins-3
  100 continue
      m#3
      n=5
       xnorm=1.007d+1
       a(6)=1.0d+0
       a(5)=5.684d+0
       a(4)=1.510d+1
       a(3)=2.217d+1
       a(2)=1.846d+1
       a(1)=7.220d+0
       go to 400
```

```
200 continue
      01:24
      n=10
      xnorm=2.921d+15
      a(11)=1 Od+O
      a(10)=4.315d+2
      a(9) = 8.449d+4
      a(8) = 7.520d + 6
      a(7) = 4.032d + 8
      a(6) = 1.161d + 10
      a(5) = 2.041d + 11
      a(4) = 1.482d + 12
      a(3) = 2.348d + 12
      a(2) = 1.589d + 12
      a(1) = 2.665d + 11
      go to 400
  300 continue
      Ts=1.0 Tg=0.23 hs=0.8 hg=0.8 sigma=0.0
C
      Emm
      n = 4
      xnorm=. 307321224d+2
      a(5)=1.0d+0
      a(4) = .855652d + 1
      a(3) = .310340d+2
      a(2) = .372023d+2
      a(1) = .189036d + 2
  400 continue
      ww=dcmplx(O.Od+O,dble(freq))
      ss(1)=ww
      do 500 i=2,n
      55(i)=55(i-1)*ww
  500 continue
       tt=demplx(a(i), 0, 0)
       do 600 i=1, n
       tt=tt+a(i+1)%ss(i)
  400 continue
       tt=ss(m)/tt
       xnorm=xnorm*peak
       xr=real(tt)*xnorm
       xi=dimag(tt)*xnorm
       return
       end
C.
C
C
       subroutine seiplt(xmin, xmax, ymin, ymax, xmini, xmaxi)
       CALCOMP plot.
C
       plot axis only.
C
       common/pltses/ tO, yend, y1, iplt, jctrl, bb, cc
       common/plott/ xnp1, xnp2, ynp1, ynp2, xr, yr
       character*2 bb,cc
       xinut: how many unit per inch
C
       xseg: how many unit for one segment in time axis
c
```

```
if(ketrl.le.jetrl) go to 100
   write(b) * '
   write(6.4) (Yotal time axis from ', xmin, ' to ', xmax
   write(6, *)
  a 'enter time window to be plotted: (start and end time)'
    write(6,4) ' (use -1,-1 to plot the whole time span)'
    read(5,*) xmin1, xmax1
    if(xmin1, 1t, 0, 0, and, xmax1, 1t, 0, 0) then
        xmin1=xmin
        xmax1=xmax
    endif
    xmn1=abs(xmax1-xmin1)
    if(iplt.eq.1) go to 80
    write(6,*)
   4 'how many inches for time axis, (total= ',xmn1,'sec)'
    write(6,*) 'how many inches for amplitude axis,'
    write(6,%)
   * 'and how many units for one segment? (as 5.0.0.8.10.0)'
    read(5,*) xl,yl,xseg
   continue
80
    write(6, *)
   * 'how many plots FOLLOWED will use above numbers?'
    write(6,*) '(if not know, enter 100)'
    read(5, *) jctrl
    kctr1=0
100 continue
    kctrl=kctrl+1
    if(iplt.eq.1) return
    x & mu=x x mu
    if(abs(umin), gt. umax) ymxx=abs(ymin)
    gend1=gend-g1-0.68
    if(yend1, 1t, 0, 5) call plot(8, 5, y1-yend, -3)
    1bney=bney
    if(yend1.1t.0.5) yend=y1-y1-0.68
    call plot(0, 0, -y1-0, 68, -3)
    xinut=xmn1/xl
    ji=xmin1/xseq
    if(xmin1, 1t, 0, 0) ii=ii-1
    xO≕xsen⊁ii
    xnp1=xQ
    xnp2=xinut
    ynp1=-ymxx
    ynp2=2. *ymxx/yl
    power=alog10(ymxx)
    if(power. lt. O. ) power=power-1.
    power=aint(power)
    yy=ymxx/10 **power
    call plot(-0.16,0.0,3)
    call plot(-0.1,0 0,2)
    call plot(-0.1, y1, 2)
    call plot(-0.16, yl, 2)
    call symbol(-0.175,-0.025,0.09,1h-,90.,1)
    call number (999., 999., 0.09, yy, 90., 1)
```

```
tall number (999., yl-0.19, 0.09, yy, 90., 1)
     call symbol(-0.35, y1*0.25, 0.11, 3h*10, 90., 3)
     call number(-0.4,997.,0.06,power,90.,-1)
     call symbol(-0.8, gl*0.45, 0.12, cc, 0.0, 2)
     xr = x1 + 0.03
     ur=0.5*41-0.05
     xlen=abs(xmax1-x0)/xinut
     orid=xseq/xinut
     call plot(0.0,-0.1,3)
     call plot(0.0,-0.14,2)
     if(x0, 1t, 0, 0) xshif=-0, 15
     if(x0, ge, 0, 0) xshif=0.03
     if(x0, ge, 10, 0) xshif=-0, 08
     if(x0, ge, 100, 0) xshif=-0.11
     call number(xshif, -0.25, 0.08, x0, 0.0, -1)
     call plot(0, 0, -0, 1, 3)
     x i = 1.
 400 xx=xi*grid
     call plot(xx, -0.1,2)
     call plot(xx,-0.14,2)
      xsym=xO+xseg*xi
      if(xsym. lt. 0. 0) xshif=-0.15
      if(xsym.ge. 0. 0) xshif=-0.03
      if(xsym.ge. 10.0) xshif=-0.08
      if(xsym.ge. 100.0) xshif=-0.11
      if(xsym. ge, 1000.0) xshif=-0.15
      call number(xx+xshif, -0, 25, 0, 08, xsym, 0, 0, -1)
      call plot(xx, -0.1, 3)
      xi = xi+1.
      if(abs(xx), gt. xlen) go to 500
      go to 400
 500 continue
      if(tO.le.O.O) go to 520
      call symbol(xx/2,3/-0,45/0,14/17-1/0,0/2)
      call number (999., 999., 0.1, t0, 0.0, 2)
      go to 600
  520 call symbol(xx/2.05,-0.45,0.14,'T',0.0,1)
  600 continue
      call plot(0.0,0.0,3)
      return
      end
C,
C
      subroutine seical(j0, n1, nn, dt, r0, xmin1, xmax1)
      plot seismogram.
C
      common/dimen/ x(1026), y(1026)
      common/plott/ xnp1, xnp2, ynp1, ynp2, xr, yr
      common/pltses/ tO, yend, y1, iplt, jctrl, bb, cc
      character*2 bb,cc
      i0=(j0-1)*1024*1
      i00=i0+1023
       if(nn.lt.1024) i00=i0+nn-1
```

```
xQ=(iO-1)*dt
     x00=(i00-1)*dt
     if(xO, le, xmin1, and, xOO, le, xmin1) go to 500
     if(xQ, ge, xmax1, and, xQQ, ge, xmax1) go to 500
     do 100 i=1, nn
     x(i)=xO+(i-1)*dt
 100 continue
     do 200 i=1, nn
     if(x(i), ge, xmin1) go to 250
 200 continue
 250 continue
     mm=i-1
     do 300 i=1, nn
     if(x(i), gt, xmax1) go to 350
 300 continue
 350 continue
     mm1=i-1
      kk=mm1-mm
      do 400 i=1, kk
      j≖mm+i
      x(i) = x(j)
      y(i)=y(j)
 400 continue
      if(iplt. eq. 1) go to 420
      x(kk+1)=xnp1
      x(kk+2)=xnp2
      g(kk+1)=ynp1
      u(kk+2)=ynp2
      call line(x, y, kk, 1, 0, 0)
 420 continue
      if(bb, eq. 'n') go to 500
      do 450 i=1, kk
      write(9,10) x(i),y(i)
     format(2e15.8)
 450 continue
      if(iplt.eq. 1) return
  500 continue
      if(j0, eq. n1) call number(xr, yr, 0, 1, r0, 0, 0, -1)
      neturn
      end
*****
```

- 85 -

```
program spec81.f
C
c
      f77 -i -I2 spec81 f -o spec81 -lcalcompI2
C
C
      This program follows the program wig81 and plots the
C
      spectrum after passing through an instrument.
C
C
      The source mechanism is included here.
C
      The whole processes are treated in the freq domain.
C
C
C
      The dimension needed has been reduced to the minimum,
      and the number of series point is not limitted.
C
C
      -NOV 1, 1981
C
C.
      dimension jj(5), az(100), baz(100)
      complex data(10,2), rr, tt, ff
      common/dimen/ x(1026), y(1026)
      common/ctrl/ ictrl/size/xy(6)
      character*2 form(5),cc,yorn
      character*50 names
      data form/'Z ', 'R ', 'SH', 'NS', 'EW'/
      format(a)
  10
      format(/1x,a2,' component for station at R=',f8.1,' Az=',
                 f6.1, ' is being processed.')
      format(/1x, 'ins=', i2, 9x, 'peak= ', f7, 1, '
  20
                                                  iresp= ', i2/
                 / dt= /, f9.5, /
                                _ ndist≕ ',i2,7x,'dphsrc≕ ',f6,1/
                 ′ Z, R, SH, N, E:
                                  ',5i2,5x,' Quake,Expl: ',i2/
                 ' Spectrum Plot: ', i2/
                 / dip= ', f5, 1, 5x, 'slip= ', f5, 1, 5x, 'strk= ', f5, 1)
      write(6, 3) ' '
      write(6,*) 'calculation only(1), plot only(2), or both(3): '
      read(5,*) iplt
      write(6,*) ′ ′
      write(6,%) 'enter the input file name:'
      read(5,5) names
      if(iplt.ne.2)
     # open(1, file=names, status='old', form='unformatted')
      if(iplt.eq.2)
        open(9, file=names, status='old', form='formatted')
      rewind 1
      rewind 9
      if(iplt.eq.2) go to 60
      open(2,file='tmpO.d',status='scratch',form='unformatted')
      open(4, file='tmpp, d', status='scratch', form='unformatted')
      if(iplt.ne.1) go to 60
      write(6,*) 'enter the output file name: '
      read(5,5) names
      open(3,file=names,status='new',form='formatted')
      rewind 3
  60
     continue
```

c-120 input control parameters.

```
if(iplt. eq. 1) go to 70
   call plots(0,0,7)
   write(6,%) 'original pen move: (2.0,2.0)'
   read(5,4) x1,41
   call plot(x1, y1, -3)
    ictrl=0
   write(6,*) 'choose the pen: '
   read(5, ≥) ipen
    call newpen(ipen)
   write(6, *) 'enter size (e, g. 1.0): '
   read(5,*) size
    call factor(size)
    write(6,*) 'plot the observed data? (y/n): '
    read(5,5) yorn
    if(yorn.eq.'n') go to 70
    write(6,*) 'enter the file name of observed data:'
    read(5,5) names
    open(8, file=names, status='old', form='formatted')
    rewind 8
70
    continue
    if(iplt.eq.2) go to 140
    write(6,*) (X-axis to be period(1) or freq(2): '
    read(S,*) iaxis
    if(ip1t.eq. 2) go to 140
    write(6,*) 'choose the type of instrument'
                                       =1 30-100 WWSSN'
    write(6,*) '(=0 15-100 WWESN
    write(6,*) / =2 6824-13 LP SYSTEM =3 6824-2 LP SYSTEM/
                                        =5 USGS SP'
    write(6,*) / =4 WWSSN SP
    write(6,*) / =6 CSSN SP/
    write(6, *) ′ =-) do NOT pass through any instrument )′
    read(5,*) ins
   write(6, *)
       'enter the peak value for instrument response: '
    read(5,%) peak
    write(6, *)
             'take the integration(-1), derivative(1), or not(0)'
    read(5,*) iresp
    read(1) dt.ndist.kkr.kkl.kkf
    kk=kkr+kkl
    read(1) dphsrc, nper
    write(6, *) ' '>
    if(kk, eq. 2) write(6,*)
        'Both Rayleigh and Love eigens been generated.'
    if(kkr. eq. 1. and, kkl. eq. 0) write(6, 4)
         'Only Rayleigh eigens been generated.'
    if(kkr. eq. 0, and, kkl. eq. 1) write(6,*)
         'Only Love eigens been generated.'
    write(6,*) ' '
    write(6,*)
        'Which components to be plotted? (Z,R,SH,NS,EW)'
    write(6,*)
        'if yes answer 1, no answer 0 (e.g. 1,0,0,1,1)'
```

```
read(5,*) (jj(i), i=1,5)
      if(kkr.ne.O) go to 100
      j=jj(1)+jj(2)
      if(j.eq.0) go to 90
      write(6,*) 'no P-SV wave ready, run again, '
      go to 1000
 90
      continue
 100 if(kkl.ne.0) go to 120
      if())(3), eq. 0) go to 110
      write(6,*) 'no SH wave ready, run again, '
      go to 1000
  110 continue
  120 continue
c-130 input source mechanism.
      write(6,4) ' '
      write(6,%) 'no. of distance = (,ndist, ' dt=',dt
      write(6,4)
        _ 'source depth == ', dphsrc, ' spectrum type == ', kkf
      write(6, *) 'earthquake source(1) or explosion(2)?'
      read(5, *) m3
      degrad=3.141592653/180.0
      if(m3.eq.2) go to 130
      write(6,*) 'enter the source mechanism dip.slip.strk:'
      read(5,*) dipO, slipO, strkO
  130 continue
C
      if(iplt.eq.1) write(3,*) 'control parameter:'
      if(iplt.eq. 1) write(3,*) ins.peak, iresp.dt, ndist, dphsrc,
                                   kkf, jj, m3, dipO, slipO, strkO
      go to 150
  140 continua
      read(9,5) names
      read(9,*) ins.peak.iresp.dt.ndist.dphsrc.kkf.jj.m3.
                   dipO, slipO, strkO
      write(6,20) ins.peak.iresp.dt.ndist.dphsrc.jj.m3.kkf.
                     dipO, slipO, strkO
  150 continue
C
c-900 the main loop.
C
       iarm=0
       iaz≕-1
       do 900 kdist=1, ndist
       iaz=iaz+1
       if(iplt.eq. 2) go to 210
       read(1) r0, t0, np0
       write(6,*) ' '
       write(6,*) 'r=(,r0,' t0=',t0,' np0=',np0
       urite(6)*) / /
       if(iaz. lt. iazm) go to 175
       iaz=-1
```

```
write(6,*) (enter Az, bAz (in degree): (
     write(6, *) '(use bAz=0.0 to set bAz=Az+180.0'
      write(6,*) ' use -1,-1 to stop)'
      k = 1
 160 continue
      read(5,*) az(k),baz(k)
      if(a_2(k), e_3, -1, 0, a_1d, b_2(k), e_3, -1, 0) go to 170
      if(baz(k), eq. 0, 0) baz(k)=az(k)+180.0
      if(baz(k), ge, 360, 0) baz(k)=baz(k)-360, 0
      k = k + 1
      go to 160
  170 \text{ naz=k-1}
     write(6,*)
         'how many sets of different distances (r) FOLLOWED'
      write(6,*) ' will use ',
         'the above azimuth data: (if not know, enter 100)'
      read(5,*) iazm
  175 continue
      write(6,*) ' '
      write(6,*) 'wait'
      rewind 4
      k == 1
  180 continue
      do 190 j=1, kkf
      read(1) per,(data(i,j),i=1,10)
  190 continue
      if(per. le. 0. 0) go to 200
      write(4) per, ((data(i, j), i=1, 10), j=1, kkf)
      k = k + 1
      go to 180
  200 continue
      np2=k-1
     if(iplt.eq.1)
     write(3,*) 'r, t0, npt, nstation, az1, baz1, az2, baz2...'
      if(iplt.eq.1)
         write(3,*) rO, tO, npO, naz, (az(k), baz(k), k=1, naz)
  210 continue
      if(iplt.eq.2) read(9,5) names
      if(iplt.eq.2)
         read(9,*) rO, tO, npO, naz, (az(k), baz(k), k=1, naz)
      do 900 kaz=1, naz
      phiO=az(kaz)
      phi1=baz(kaz)
      if(iplt.eq.2) go to 310
      phi=phil*degrad
      cos1=cos(phi)
      sin1=sin(phi)
      if(m3.eq.2) go to 300
c-300 dislocation source.
      strk1=phiO-strkO
      dip=dipO*degrad
```

```
slip=slipO*degrad
     strk1=strk1*degrad
     strk2=2. *strk1
     dip2=2. *dip
     s1=sin(slip)*sin(dip2)
     s2=-(cos(slip)*sin(dip)*sin(strk2)+
             O. 5*sin(slip)*sin(dip2)*cos(strk2))
      s3≕sin(slip)*cos(dip2)*sin(strk1)-
         cos(slip)*cos(dip)*cos(strk1)
      tl=sin(slip)*cos(dip2)*cos(strk1)+

    cos(slip)*cos(dip)*sin(strk1)

      t2=cos(slip)*sin(dip)*cos(strk2)-
           Q. S*sin(slip)*sin(dip2)*sin(strk2)
  300 continue
  310 continue
C
      do 900 j=1,5
      if(jj(j), eq. 0) go to 900
      cc=form(j)
      write(6,10) cc,r0,phi0
      do 900 kf=1, kkf
      if(iplt.eq.2) go to 640
      rewind 2
      rewind 4
      kzero=0
      do 500 k#1/np2
      read(4) per, ((data(i, j0), i=1, 10), j0=1, kkf)
      freq=1./per
      go to (400,320),m3
  320 go to (330,340,350,360,370), j
  330 ff=data(4,kf)
      go to 460
  340 ff=data(8,kf)
      go to 460
  350 write(6,*) 'explosive source no SH wave, run again, '
      go to 1000
  360 ff=-cosi*data(8,kf)
      go to 460
  370 ff=-sin1*data(8,kf)
      go to 460
  400 go to (410, 420, 430, 440, 440), J
  410 ff=s1*data(1,kf)+s2*data(2,kf)+s3*data(3,kf)
      go to 460
  420 ff=s1*data(5,kf)+s2*data(6,kf)+s3*data(7,kf)
      go to 460
  430 ff=t1*data(9,kf)+t2*data(10,kf)
      go to 460
  440 rr=s1*data(5, kf)+s2*data(6, kf)+s3*data(7, kf)
      tt=t1*data(9,kf)+t2*data(10,kf)
      ct1=-cos1
      sti=sin1
      if(j.eq.4) go to 450
      ct1=-sin1
```

```
sti=-cosi
 450 ff=ct1*rr+st1*tt
 460 continue
      call resp(ff, freq, ins, peak, iresp)
      gg=sqrt(real(ff)*real(ff)+aimag(ff)*aimag(ff))
      if(gg.le.1.e-35) kiero=kzero+1
      if(gg.le.1.e-35) go to 500
      write(2) freq.gg
  500 continue
      npp=np2-krero
c-600 find the max min
      rewind 2
      ymax = -1.0e + 38
      ymin= 1.0e+38
      xmax=-1. Oe+38
      xmin = 1.0e + 38
      do 600 k=1, npp
      read(2) freq.gg
      if(gg.gt.ymax) ymax=gg
      if(qq. lt. ymin) ymin≖gg
      if(freq.gt.xmax) xmax=freq
      if (freq. lt. xmin) xmin=freq
  600 continue
      tmp=xmin
      if(iaxis.eq.1) xmin≕1./xmax
      if(iaxis.eq.1) xmax≕1./tmp
      rewind 2
      if(iplt.ne.1) go to 640
      write(3.*) 'nptO, xmin, xmax, ymin, ymax, T/f: '
      write(3)*) npp,xmin,xmax,ymin,ymax,iaxis
                                                    amp '
      if(iaxis, eq. 1) write(3,*) ' period
                                                    amp '
      if(iaxis, eq. 2) write(3,*) '
      go to 660
  640 continue
      if(iplt.ne.2) go to 650
      read(9,5) names
      read(9,*) npp, xmin, xmax, ymin, ymax, iaxis
      read(9,5) names
  450 continue
       plot axis
ς
       if(kf.eq. 1) call sppplt(xmin, xmax, ymin, ymax, cc, iaxis)
       if(kf.eq.1.and.yorn.eq.'y') call obs(iaxis)
  660 continue
      n1=npp/1024
       nm=npp-n1*1024
       n1=n1+1
       if(nm, eq. O) n1=n1-1
c-800 divide the whole length into sub-lenth with 1024 as max.
       do 800 k=1, n1
       n0=1024
       if(k.eq.nl.and.nm.ne.O) nO=nm
       if(iplt.eq.2) go to 750
```

```
do 700 i=1,n0
      read(2) x(i), y(i)
      if(iaxis, eq. 1) x(i)=1, /x(i)
  700 continue
      if(iplt, eq.1) write(3,40) (x(i),y(i),i=1,n0)
      if(ip1t.eq.1) go to 800
  750 continue
      if(iplt.eq 2) read(9,40) (x(i), y(i), i=1, n0)
      format(2e13.7)
      call sppcal(nO)
  800 continue
  900 continue
  1000
               continue
      if(iplt.ne.1) call plot(8.0,0.0,999)
      close(1)
      close(2, status='delete')
      close(3)
      close(4, status='delete')
      if(yorn.eq. 'y') close(8)
      if(iplt.eq.2) close(9)
      write(6,*) / /
      write(6,*) 'spec81 finished'
      write(6,*) ' '
      stop
      ≋ಗರ
C
C
C
      subroutine resp(ff, freq, ins, peak, iresp)
      This routine imposes the instrument response on the
C
c
      seismogram by frequency domain multiplication.
      complex ff
      Prareal(ff)
      fi=aimag(ff)
      if(ins. lt. 0) go to 250
      in = 0
C
              001-51 NRBWW
                                      = 1
                                           WWSSN 30-100
          = 2
C
               LP SYSTEM 6824-13
                                      ## (B
                                           LP SYSTEM 6824-2
C
          =: 4
               WWSSN SP
                                      ≈ 5
                                           USGS SP
C.
          = 6
               CGSN SP
      go to (110,110,120,120,130,130,130), ins+1
  110 call www.sn(freq.peak.ins.pr.pi)
      go to 200
  120 call lpsys(freq.peak,ins.pr.pi)
      go to 200
  130 call spend(freq.peak, ins.pr.pi)
  200 continue
      tmp=fr
      fr=fr*pr-fi*pi
      fi=tmp*pi+fi*pr
  250 if(iresp. eq. 0) go to 400
      omega=2. *3, 141592653*freq
      if(iresp. eq. -1) go to 300
      tmp=fr
```

```
fr=-fi*omega
      fi≕tmp*omega
      go to 400
  300 continue
      tmp=fr
      fr≕fi/omega
      fi=-tmp/omega
  400 ff=cmplx(fr,fi)
      return
      end
C
C
        subroutine www.sn(freq.peak.ins.xr.xi)
C
      ins eq 0 - 15-100 wwssn
      ins eq 1 30-100 wwssn
C
      peak magnification is fixed at 350,700,1400,2800,5600
C
        we=6.2831853*freq
        index=(peak+1.)/375.
        if(ins.gt.O) go to 200
        go to (1,2,2,3,3,3,3,4,4,4,4,4,4,4,4,5), index
      1 fmag=278.
        sigma=0.003
        go to 6
      2 fmag=556.0
        sigma#0.013
        go to 6
      3 fmag=1110.0
        .sigma=0.047
        go to 6
      4 fmag=2190.0
        sigma=0.204
        go to 6
      5 fmag=3950.0
        sigma=0.805
      6 reta=0.93
        zetai=1. O
        wn=. 418879
        wn1=.062831853
        go to 300
    200 continue
        go to (10,20,20,30,30,30,30,40,40,40,40,40,40,40,40,50), index
     10 fmag=251.9
        sigma=0.003
        go to 60
     20 fmag=503.1
        sigma=0.012
        go to 60
     30 fmag=1001.5
        sigma≃0.044
        go to 60
     40 fmag=1941.9
```

sigma=0.195

```
50 fmag=2241.8
        sigma=0.767
     60 reta=1.5
        retai=1.0
        wn=. 2094395
        wn1 = .062831853
    300 continue
        ar= (we*we-wn*wn)*(we*we-wn1*wn1)-4.*reta*reta1*wn*wn1*
             (1.-sigma)*we*we
        ai=2.*we*(zeta1*wn1*(wn*wn-we*we)+zeta*wn*(wn1*wn1-we*we))
        factor = fmag*we*we*we / (ai*ai + ar*ar)
        xr =-ai * factor
        xi =-ar * factor
        return
        end
C.
£
C.
        subroutine lpsys(freq.peak,ins,xr,xi)
      lrsm response for 1p system with filter 6824-2 ins=3
C
      lrsm response for 1p system with filter 6824-13 ins=2
C
      phase response obtained from hilbert transform of amplitude
C
                  gain normalized to 1.0 at 25 seconds.
      response.
C
         dimension fre(28),p(56),phi(56)
      the first 28 p and phi are for 6824-2, next
C
      28 are for 6824-13
C
         data fre/.001,.002,.003,.004,.005,.006,.007,.008,.009,
        1.01,.02,.03,.04,.05,.06,.07,.08,.09,.1,.2,.3,.4,.5,.6,
        2.7, 8, 9,1./
         data phi/263. 9, 257. 9, 251. 9, 245. 6, 239. 3, 233. 1, 227. 1, 221. 1,
        1214. 6, 208. 3, 134. 9, 73. 2, 15. 3, -33. 2, -71. 2, -100. 1, -122. 4,
        2-140.0,-153.8,-213.1,-232.6,-242.8,-249.2,-253.0,-256.9,
        3-259, 3, -261, 2, -262, 9, 265, 0, 259, 1, 253, 4, 248, 1, 242, 8, 236, 9,
        4 230, 3, 223, 4, 216, 4, 209, 7, 153, 4, 102, 0, 54, 6, 14, 4, -16, 8,
        5 -40, 4, -58, 2, -73, 3, -85, 4, -146, 4, -153, 9, -155, 3, -156, 4,
        6-157. 7, -159. 2, -160. 0, -162. 3, -163. 8/
         data p/.00005,.00040,.00135,.00321,.00625,.01077,.01706,.02546,
        1,03631,05006,34117,73904,1,0000,97633,79807,61417,46105,
        2 34464, 26315, 03583, 01097, 00468, 00240, 00139, 00087, 00058,
        3.00040,.00029,
        4 .000015,.000391,.00131,.00310,.00609,.01068,.01713,.02556,
        5 . 03618, . 04848, . 28030, . 67553, 1. 0, 1. 09448, 1. 0044, . 86969, . 74511,
        6 . 63557, . 54818, . 12153, . 04830, . 02741, . 01827, . 01328, . 01017,
        7 .008052,.006545,.005404/
         u = 58
         degrad = 0.01745329
         if(freq. gt. 0.5) freq = 0.5
         if(freq. lt. 0.005) freq = 0.005
         do 200 i = 1 m
         if(freq.ge.fre(i).and.freq.le.fre(i+1)) go to 160
     200 continue
     160 continue
```

go to 60

```
if(ins. eq. 2) j=i+28
        if(ins.eq.3)j≕i
        pf= phi(j)+(phi(j+1)-phi(j))/(fre(i+1)-fre(i))*(freq-fre(i))
        ph = p(j) + (p(j+1) - p(j))/(fre(i+1)-fre(i))*(freq-fre(i))
        pf = pf * degrad
        ar = cos(pf) > ph * peak
        xi = sin(pf) * ph * peak
        end
C
C
C
      subroutine spend(freq.peak, ins.xr,xi)
      ins = 4 WWSSN SP
C
      ins = 5 USGS
C
      ins = 6
                CSSN
                      SP
      The values for WWSSN and USGS from Luh (1977, BSSA, p. 950)
C
C
      double complex ss(20), www.tt
      double precision knorm, a(20)
      go to (100,200,300), ins-3
  100 continue
      m=3
      n = 5
      xnorm=1.007d+1
      a(6)=1.0d+0
      a(5) = 5.684d + 0
      a(4)=1.510d+1
      a(3)=2.217d+1
      a(2)=1.846d+1
      a(1)=7.220d+0
      go to 400
  200 continue
      m=4
      n=10
       xnorm=2 921d+15
       a(11)=1.0d+0
       a(10)=4.315d+2
       a(9) = 8.449d + 4
       a(8) = 7.5206 + 6
       a(7) = 4.032d + 8
       a(6) = 1.161d + 10
       a(5) = 2.041d+11
       a(4) = 1:482d+12
       a(3) = 2.3486 + 12
       a(2) = 1.589d + 12
       a(1) = 2.665d + 11
       go to 400
  300 continue
       T_{5}=1.0 T_{0}=0.23
                         hs=0.8 hg=0.8 sigma=0.0
       m=3
       xnorm=. 307321224d+2
```

```
a(5)=1.0d+0
      a(4) = .855652d + 1
      a(3) = 310340d+2
      a(2) = .372023d+2
      a(1) = 189036d+2
  400 continue
      ww=dcmplx(O.Od+O,dble(freq))
      ss(1) =ww
      do 500 i=2, n
      55(i)=55(i-1)*ww
  500 continue
      tt=dcmplx(a(1),0.0)
      do 600 i=1, n
      tt=tt+a(i+1)*ss(i)
  600 continue
      tt=ss(m)/tt
      xnorm=xnorm*peak
      xr=real(tt)3xnorm
      xi#dimag(tt)*xnorm
      return
      end
C
C
C
        subroutine obs(iaxis)
      common/ctrl/ ictrl, size, xy(6)
        read(8,10) n
     10 format (5x, i5)
        do 100 i=1, n
        inteq=3
        read(8,20) t, y, m
     20 format (f10.8,8x, f12.6,25x, i5)
        if(iaxis.eq.2) t=1./t
        if (q. 1t. xq(3)) go to 100
        x = alog10(t)
        y=(x-xy(1))/xy(5)
        q=(q-xq(3))/xq(6)
        if (m. ne. 1) inteq=1
      call symbol(x, y, 0.07, inteq, 0.07, -1)
    100 continue
        return
         end
C
C
      subroutine sppplt(xmin, xmax, ymin, ymax, cc, iaxis)
      common/ctrl/ ictrl, size, xy(6)
      character*2 cc, ans
      character*50 names
      ictrl=ictrl+1
      if(ictrl. eq. 1) go to 100
      write(6,8) 'enter new origin for the next figure w.r.t.'
      write(6,*) 'the origin of the present figure (8.5,0.0): '
```

```
read(5,*) x0,40
    xO=xO/size
    uO=yO/sire
    call plot(x0, y0, -3)
100 continue
    write(6, a) (enter the station identification: '
   read(5,5) names
5 format(a)
    if(xmin. lt. 1. e-30) xmin=1. e-30
    if(ymin. lt. 1. e-30) ymin=1. e-30
    xu(1) = alog 10(xmin)
    xy(2) = alog 10(xmax)
    xy(3) = alog 10(ymin)
    xy(4) = alog 10(ymax)
    do 150 i=1.4
    k=int(xu(i))
    t=k
    1 + (xy(i), 1t, 0, 0) go to 140
    tt=xu(i)-t
    xy(i)=t
    if(i.eq. 1. or. i.eq. 3) go to 150
    if(tt, gt, 0, 001) xy(i)=t+1.0
    go to 150
140 tt=t-xy(i)
    xy(i)=t
    if(i, eq. 2, or. i, eq. 4) go to 150
    if(tt, gt, 0, 001) xy(i)=t-1, 0
             continue
150
    write(6,%)
      'xmin= ',xmin,'xmax= ',xmax,'ymin= ',ymin,'ymax= ',ymax
    i1=xq(1)
     i2=xy(2)
     i3=xy(3)
     i4=xu(4)
    write(6, *)
        'Log-Log range- X:', i1,' to ', i2,' Y:', i3,' to ', i4
    write((E,*)) 'give the range for spectrum plotting? (y/n)'
     read(5,5) ans
     if(ans. eq. 'n') go to 190
     write(6, *)
        'enter your range for X-axis, then for Y-axis: (0,2,-2,3)'
     read(5, *) xy(1), xy(2), xy(3), xy(4)
 190 continue
     n1=int(xy(2)-xy(1))
     n2=int(xy(4)-xy(3))
     write(6,*) 'enter the length of x- and y-axis: (5,3.75)'
     read(5,*) x1, y1
     write(6,*) 'plot the frame? (y/n): '
     read(5,5) ans
     if(ans.eq.'n') go to 360
     x12=0.5*x1
     u12=0.5*u1
     xseg=x1/n1
```

```
yseg=y1/n2
     call plot(x12,0.0,2)
     call plot(x1,0.0,2)
     call plot(0.0, yl ,3)
     call plot(0.0, y12,2)
     call plot(0,0,0,0,2)
     do 300 1=1, n1+1
     xt=xseg*(i-1)
     call number(xt-0.09, -0.2, 0.1, 10., 0.0, -1)
      yp = xy(1) + i - 1.
     call number (999., -0.12, 0.06, xp, 0.0, -1)
      if(i.eq. n1+1) go to 300
      do 300 j=2.10
      xtt=xt+xseg*alog10(real(j))
      call plot(xtt, 0.0,3)
      ytt=0.04
      if(j. eq. 10) ytt=0.07
      call plot(xtt,-ytt,2)
 300 continue
      do 350 i=1, n2+1
      yt=yseg*(i-1)
      call number(-0.38, yt-0.07, 0.1, 10.0, 0.0, -1)
      yp=xy(3)+i-1
      call number (999., yt-0.04, 0.06, yp, 0.0, -1)
      if(i.eq. n2+1) go to 350
      do 350 j=2,10
      utt=ut+useg*alog10(real(j))
      call plot(0.0, ytt,3)
      xtt=0.04
      if(j, eq. 10) xtt=0.07
      call plot("xtt, ytt, 2)
 350 continue
      call symbol(x1/3.0,y1/1.05,0.12,names,0.0,20)
      call symbol(-0.45, y1/2.1, 0.14, 'AMP', 90.0, 3)
      names='FREQ (Hz)
      if(iaxis.eq.1) names='PERIOD (sec)'
      call symbol(x1/2, 5, -0, 45, 0, 14, names, 0, 0, 12)
  360 continue
      xu(5)=real(n1)/xi
      xy(6) = real(n2)/y1
      return
      end
C
      subroutine sppcal(nn)
      common/dimen/ x(1026), y(1026)
common/ctrl/ ictrl, size, xy(6)
      do 100 i=1, nn
      x(i) = alog1O(x(i))
      y(i)=alogiO(y(i))
      if(x(i), lt, xy(1)) x(i)=xy(1)
      if(x(i), gt, xy(2)) \times (i) = xy(2)
```

```
program bigfft
C
      This performs a fft of length m=2n to yield a
C
      2n series by doing only an n length fft
      I/O is accomplished by means of the file 'bigfft d'.
c
      The data in 'bigfft.d' should be:
c
               m, dt, df, isign
C
               real, imag
                                           point1, point2
C
                                 OT
                                           point3, point4
C
C
C
         <total: m/2+1 complex data>
                                           <total: m time data>
C
                                          do xxx i=1, m, 2
      use do xxx i=1, m/2+1
c
                                     xxx write(1) time(i), time(i+1)
       xxx write(1) data(i)
C
c
      to set up the data file 'bigfft.d' in main program.
      isign=-1 : 'forward' FFT, output the lowest
                                                      freq first.
           =-2 : 'forward' FFT, output the highest freq first.
C
           =+1 : 'inverse' FFT, input
                                         the lowest freq first.
            =+2 : 'inverse' FFT, input
                                         the highest freq first.
      dt, df defined in the fft program four. f are
      if(isign. 1t, 0. or. df. eq. 0. 0) df=1. /(nn*dt)
      if(isign.gt.O.or.dt.eq.O.O) dt=1./(nn*df)
      This shows how to use 'bigfft':
      Input:
             m, dt, df
             time(i), i=1, m
                                --time series
      Program:
C
      open(1, file='bigfft.d', status='new', form='unformatted')
      rewind 1
C
C
      ision=-1
      write(1) m, dt, df, isign
C
C
      do 100 i=1, m, 2
      write(1) time(i), time(i+1)
c 100
      close(1)
C
      call system('bigfft', kturn)
C
      open(1,file='bigfft.d',status='old',form='unformatted')
ε
c
      rewind 1
      read(1) m, dt, df, isign
C
      do 200 i=1, n/2+1
C
c 200
      read(1) spec(1,i), spec(2,i)
C
      Output:
c
              m, dt, df, isign
              spec(i), i=1, m/2+1 -- complex spectrum
ር ጵጵጵጵጵጵጵጵ
      complex data(4096), x, y, z, twidle, twidel
```

```
dimension data2(2,4096)
   equivalence (data(1),data2(1,1))
    double precision cosO, cosd, sinO, sind, pp, r12, g12, rr, gg
   open(1, file='bigfft.d', status='old', form='unformatted')
   rewind 1
   read(1) m, dt, df, isign
   n=m/2
   n2=n/2
    np1=n+1
    pp=3.14159265358979d+O/float(n)
    cosd=dcos(pp)
    sind=dsin(pp)
    twidle=cmplx(cosd, sind)
    if(isign. lt. 0) go to 1000
    open(2,file='fft,l',status='scratch',form='unformatted')
    rewind 2
    if(isign.ne.+2) go to 200
    do 100 i=np1,1,-1
    read(1) data(i)
100 continue
    rewind 1
    write(1) m, dt, df, isign
    do 150 i=1, np1
    write(1) data(i)
150 continue
200 continue
    do the work
    do 500 j=1,2
    i = 1
    rewind 1
    read(1) m, dt, df, isign
    read(1) x
    if(j.eq.1) then
             data(i)=x
             i=i+1
    endif
    do 300 k=1, n2
    ii=n+2-i
    read(1) u
    read(1) x
    if(j, eq. 1) then
             data(i)≕x
             data(ii)=conjg(x)
    else
             data(i)=u
             data(ii-1)=conjg(y)
    endif
    i = i + 1
300 continue
    call four(data, n, +1, dt, df)
    dt=0.5*dt
    if(j.ne.1) go to 500
    do 400 i=1, n
```

```
write(2) data(i)
  400 continue
  500 continue
      rewind 1
      rewind 2
      twide1=cmp1x(1.0,0.0)
      do 600 i=1, n
      read(2) x
      z=twidel*data(i)
      y = x + z
      z = x - z
      data(i)=cmplx(real(y),real(z))
      twidel=twidel*twidle
  400 continue
      now consider data as a real time series of length 2n
C
      the order of time values is 1 n+1 2 n+2 3 n+3 .... n 2n
C
      we now have to demultiplex
      rewind 2
      do 700 i=1, n, 2
      write(2) data(i), data(i+1)
  700 continue
      rewind 2
      do 800 i=1, n2
      ii=i+n2
      read(2) x/y
      data2(1,i)=real(x)
      data2(2, i)=real(y)
      data2(1, ii)=aimag(x)
      data2(2, ii)=aimaq(q)
  800 continue
      rewind 1
      write(1) m, dt, df, isign
      do 900 i=1, n
      write(1) data2(1, i), data2(2, i)
  900 continue
      close(2)
      go to 2000
 1000 continue
      forward Fourier transform
      do 1100 i=1, n
      read(1) data2(1,i), data2(2,i)
 1100 continue
      call four (data, n, -1, dt, df)
      df=0.5*df
      cos0=1. Od+0
      sin0=0.0d+0
      data2(1, n+1)=data2(1, 1)-data2(2, 1)
      data2(1,1)=data2(1,1)+data2(2,1)
      data2(2, n+1)=0.0
                                    112 = m/2
      data2(2,1)=0.0
      do 1200 i=2.n2+1
           =cos0*cosd-sin0*sind
       sinO=sinO*cosd+cosO*sind
```

```
cosO≔pp
      i1 = n + 2 - i
      r1=data2(1,i)
      r2=data2(1, i1)
      g1=data2(2, i)
      o2=data2(2, i1)
      r12=r1-r2
      g12=g1+g2
      rr=cos0*g12-sin0*r12
      gg#sinO*g12+cosO*r12
      gg = -gg
      r12=r1+r2
      g12=g1-g2
      data2(1,i) =r12+rr
      data2(1, i1)=r12-rr
      data2(2,i) = qq+q12
      data2(2, i1)=gg-g12
 1200 continue
      do 1300 i=2, n
      data2(1, i)=0.5*data2(1, i)
      data2(2, i)=0.5*data2(2, i)
 1300 continue
      rewind 1
      write(1) m, dt, df, isign
      do 1400 i=1, np1
      if(isign, eq. -2) i1=n+2-i
      write(1) data(i1)
 1400 continue
 2000 continue
      close(1)
      stop
      end
С
      subroutine four(data, nn, isign, dt, df)
      FFT
C
      dimension data(1)
      n = 2 * nn
      if(isign. eq. -1 . or. df. eq. 0.0) df=1.0/(nn*dt)
      if(isign.eq. +1 .or. dt.eq. 0.0) dt=1.0/(nn*df)
      J == 1
      do 5 i=1, n, 2
      if(i-j)1,2,2
      tempr = data(j)
      tempi = data(j+1)
      data(j) = .data(i)
      data(j+1)=data(i+1)
      data(i) = tempr
      data(i+1) = tempi
  2
      w = u/5
  3
      if(j-m) 5,5,4
```

```
1 = 1-4
     m = m/2
     if(m-2)5, 3, 3
     j=j+m
     mmax = 2
     if(mmax-n) 7,10,10
 6
     istep= 2 *mmax
     theta = 6.283185307/float(isign*mmax)
     sinth=sin(theta/2.)
     wstpr=-2. *sinth*sinth
     wstpi=sin(theta)
     wr=1.0
     wi=0.0
     do 9 m=1/mmax/2
     do 8 i=m, n, istep
     j≖i+mmax
     tempr=wr*data(j)-wi*data(j+1)
     tempi=wr*data(j+1)+wi*data(j)
     data(j)=data(i)-tempr
     data(j+1)=data(i+1)-tempi
     data(i)=data(i)+tempr
     data(i+1) = data(i+1)+tempi
 8
      tempr = wr
     wr = wr*wstpr-wi*wstpi + wr
     wi = wi*wstpr+tempr*wstpi + wi
      mmax = istep
      go to 6
  10
      continue
      if(isign. 1t. 0) go to 1002
      frequency to time domain
C
      do 1001 iiii = 1/n
 1001 data(iiii) = data(iiii) * df
      return
 1002 continue
      time to frequency demain
      do 1003 iiii = 1, n
 1003 data(iiii) = data(iiii) * dt
      return
      end
******
```

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